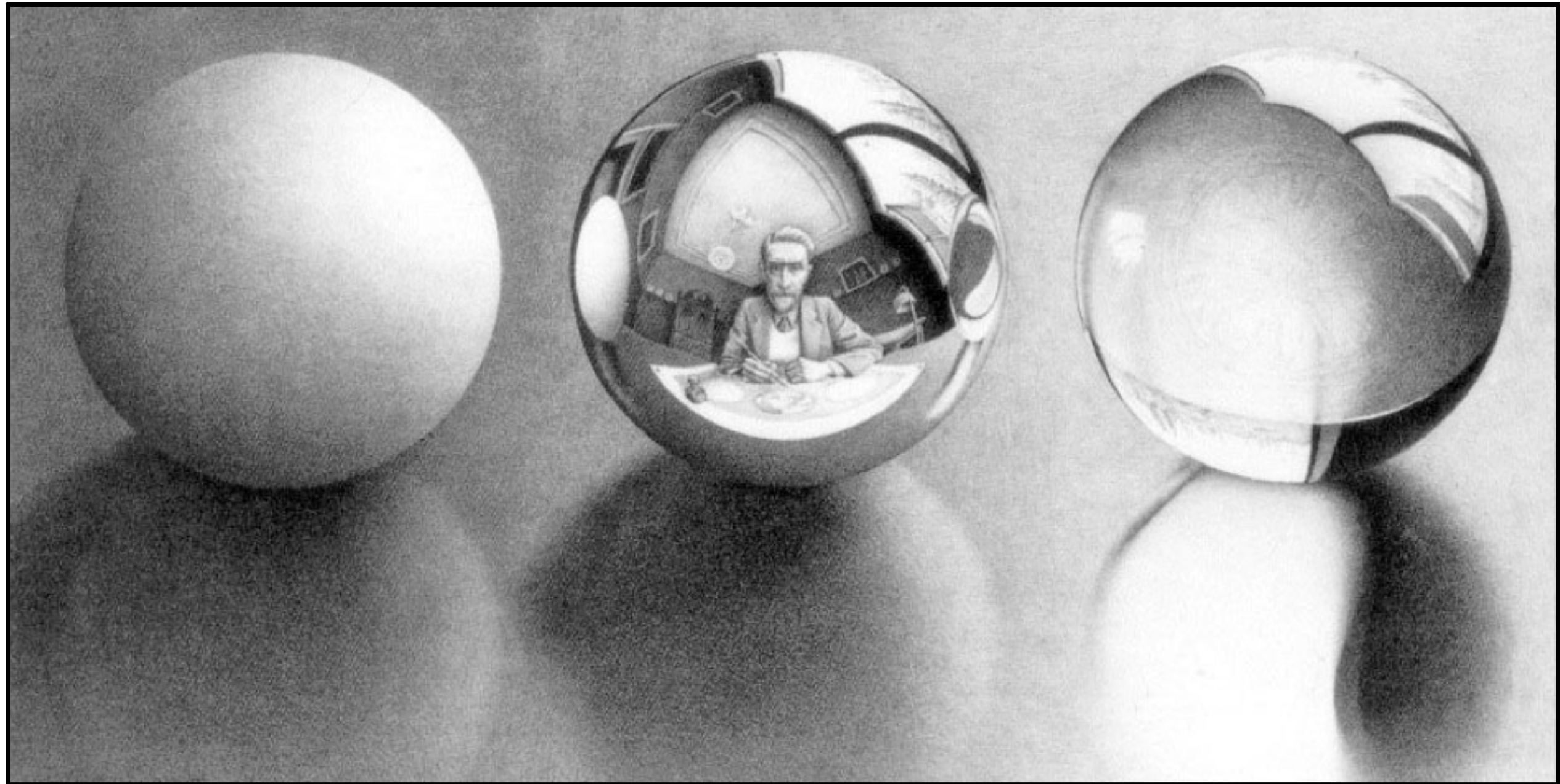


Ray tracing and simple shading



15-468, 15-668, 15-868
Physics-based Rendering
Spring 2025, Lecture 3

Course announcements

- Vote on Slack for first make-up lecture.
- Programming assignment 0 available on Canvas.
 - Ungraded, no due date.
 - Used to set up rendering environment and github-based submission system.
 - Should take no more than 1-2 hours max.
- Due dates for remaining homework shifted:
 - Programming assignment 1 will be posted on Friday, will be due two weeks later.
 - Take-home quiz 1 will be due next Tuesday.

Overview of today's lecture

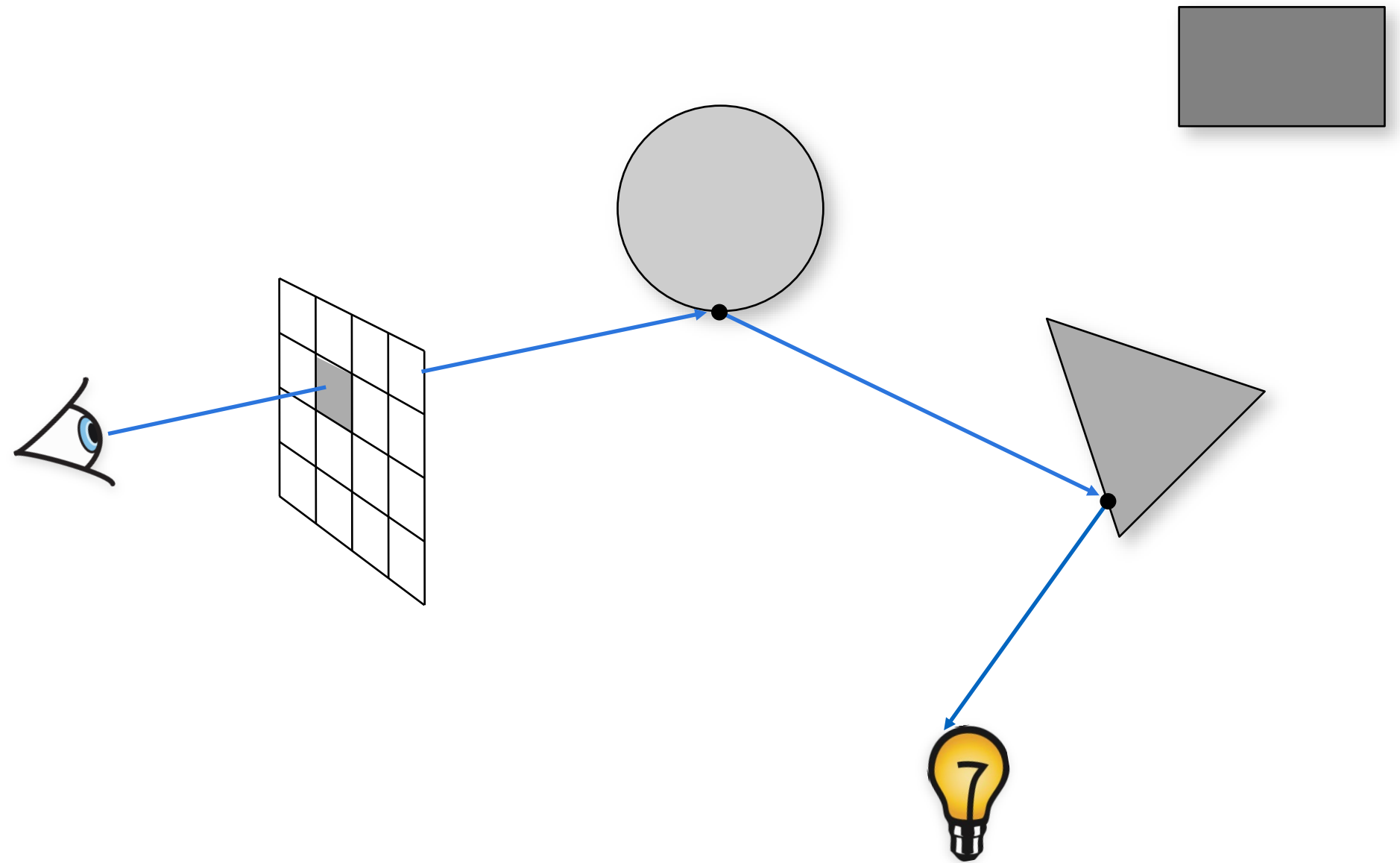
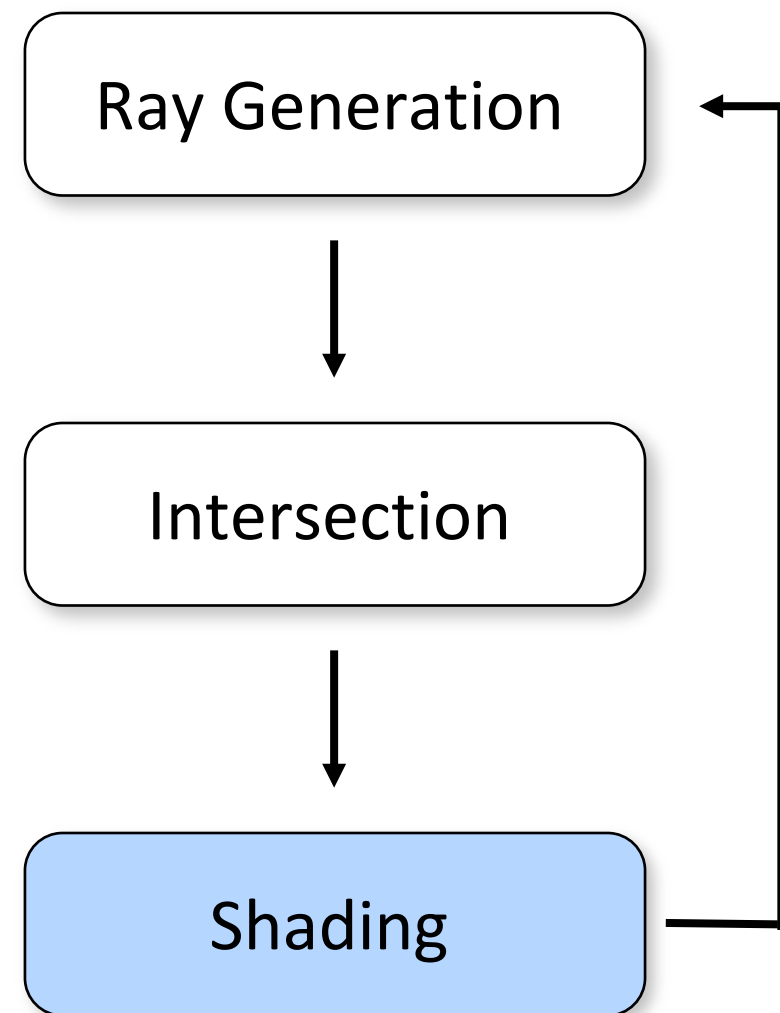
- Leftover from previous lecture: intersections, meshes, acceleration structures.
- Basics of shading.
- Basic reflection models.

Slide credits

Most of these slides were directly adapted from:

- Wojciech Jarosz (Dartmouth).

Recap: Raytracing



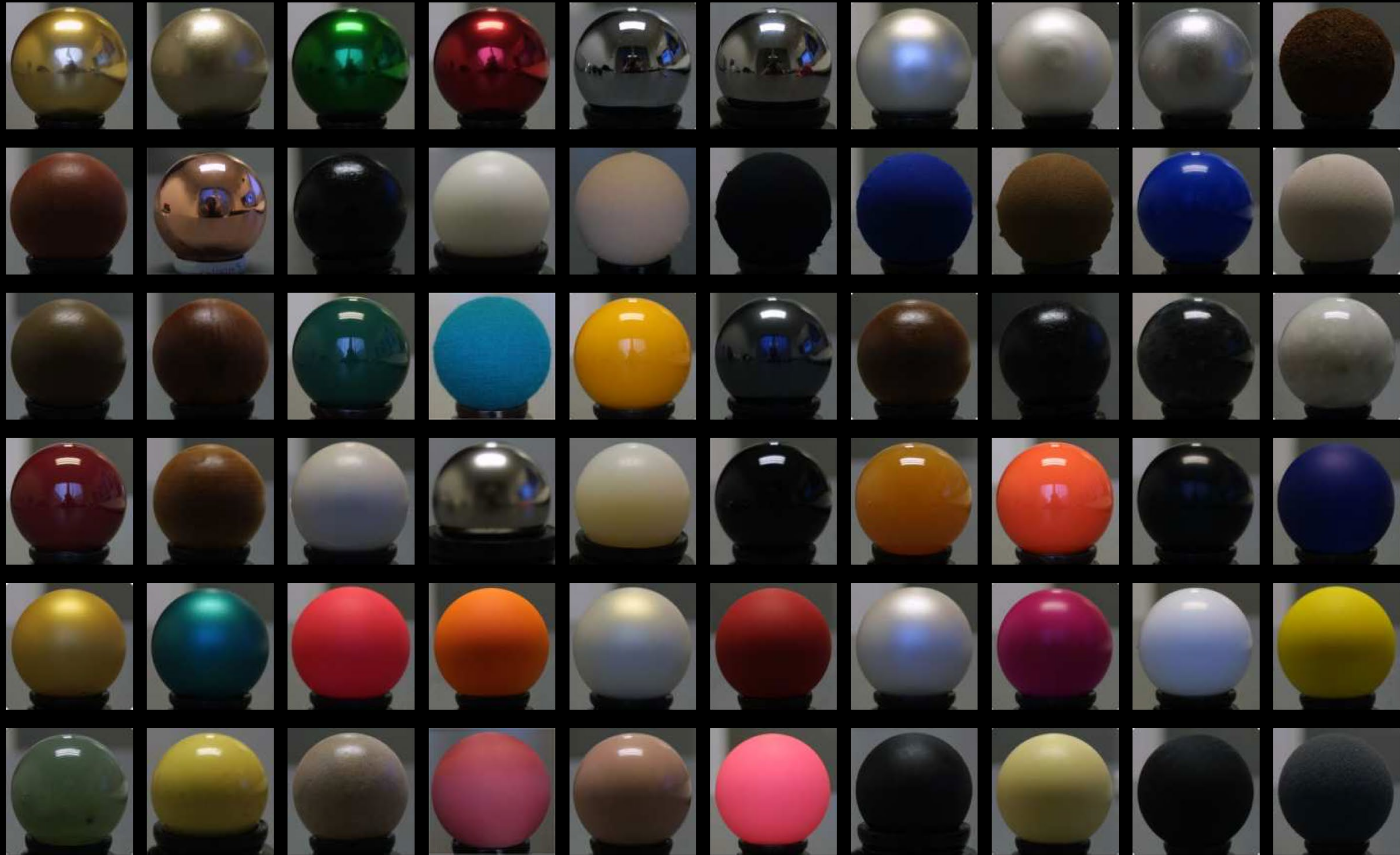
Shading

When ray hits a surface we perform *lighting/shading*

Determine “what color/light should we see at this location?”

Surfaces can scatter and/or emit light

- Surface emits light? just return emitted color
(determined by the material)
- Surface scatters/reflects/refracts light?
(recursively) trace a ray in a scattering direction
(determined by the underlying material)



Overview

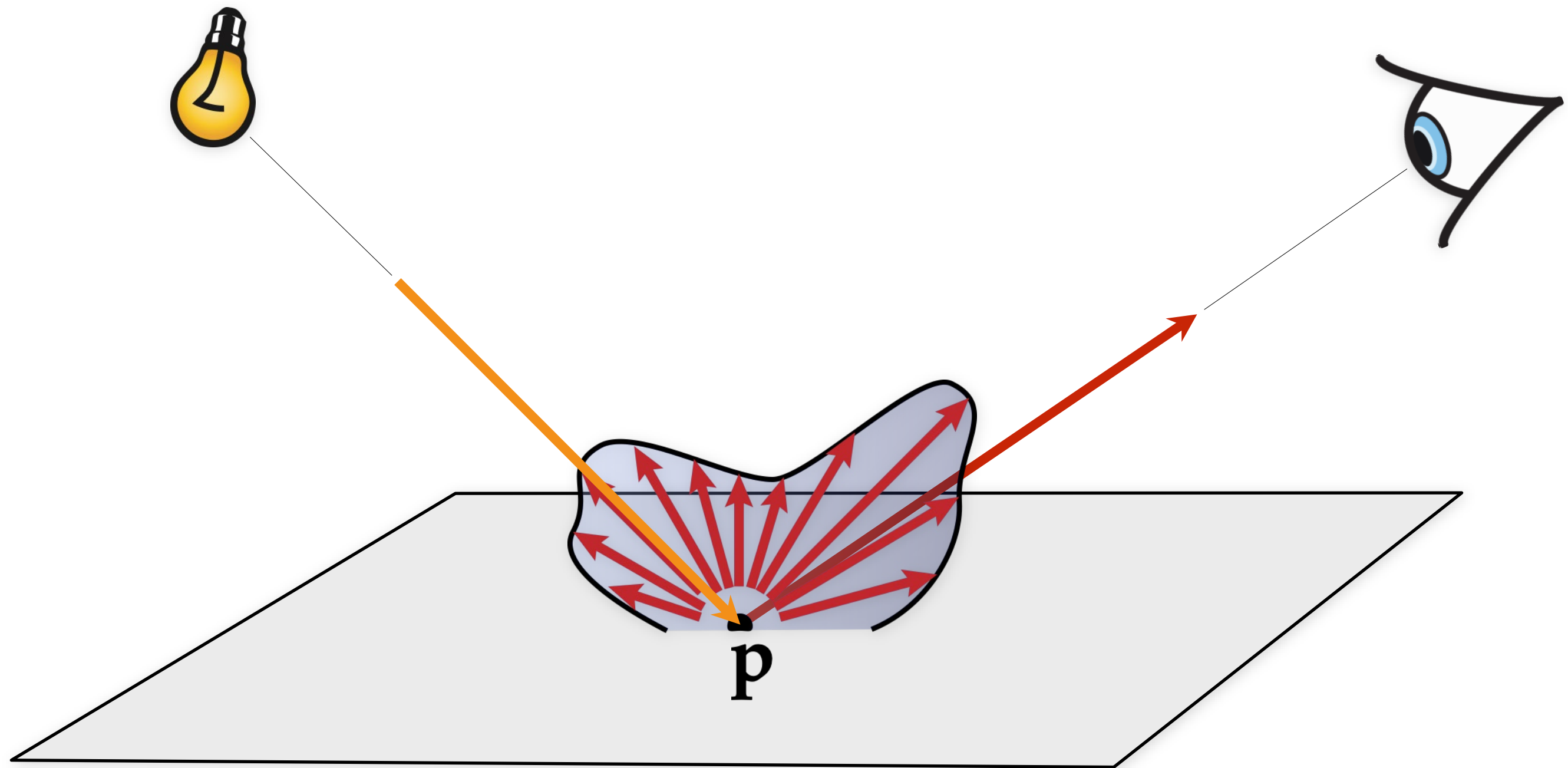
Diffuse shading

Specular reflection

Refraction

Diffuse emission

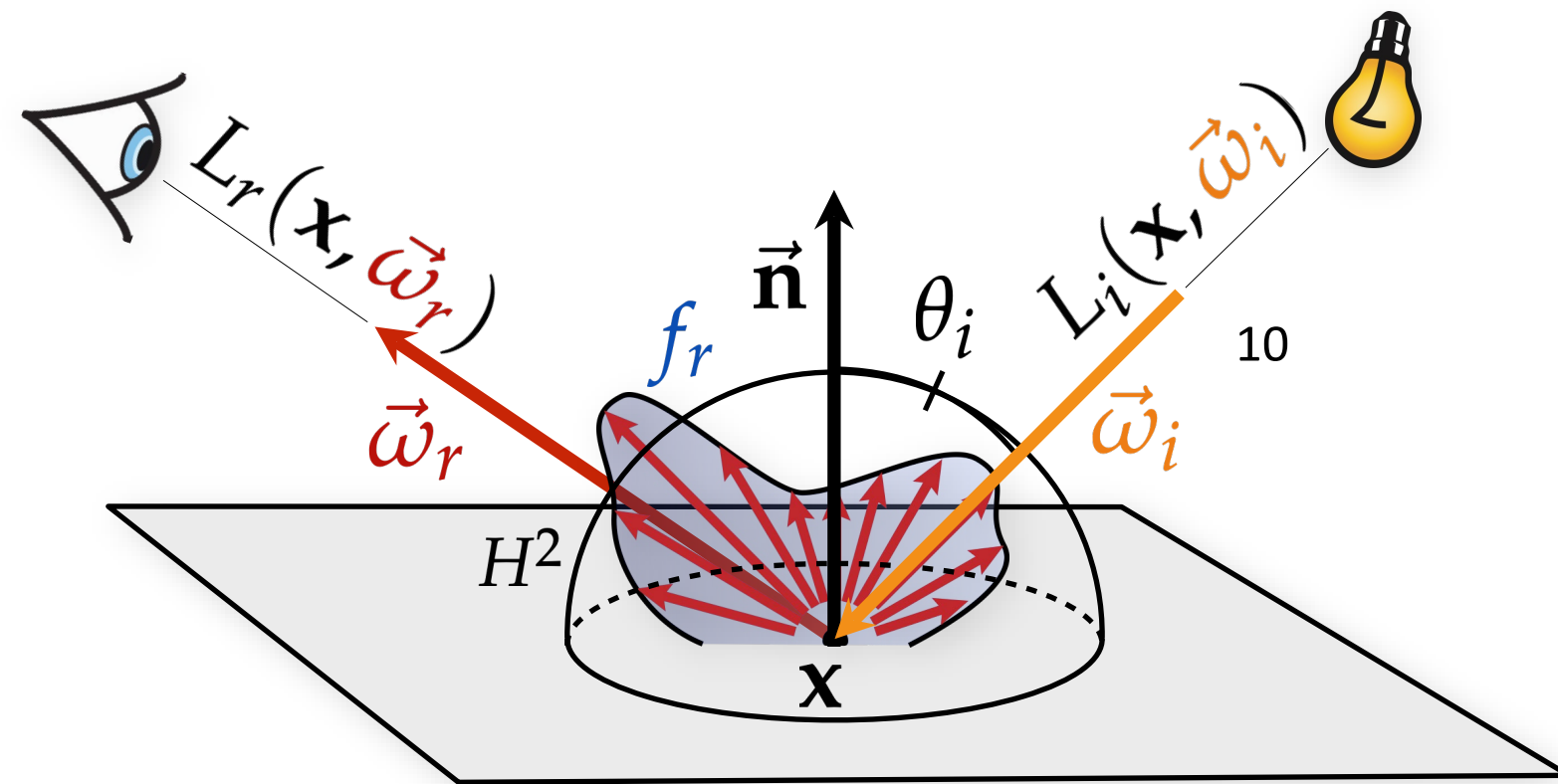
Light-material interactions



The reflection equation

Reflected radiance is a (hemi)spherical integral of incident radiance from all directions

$$L_r(\mathbf{x}, \vec{\omega}_r) = \int_{H^2} f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i$$

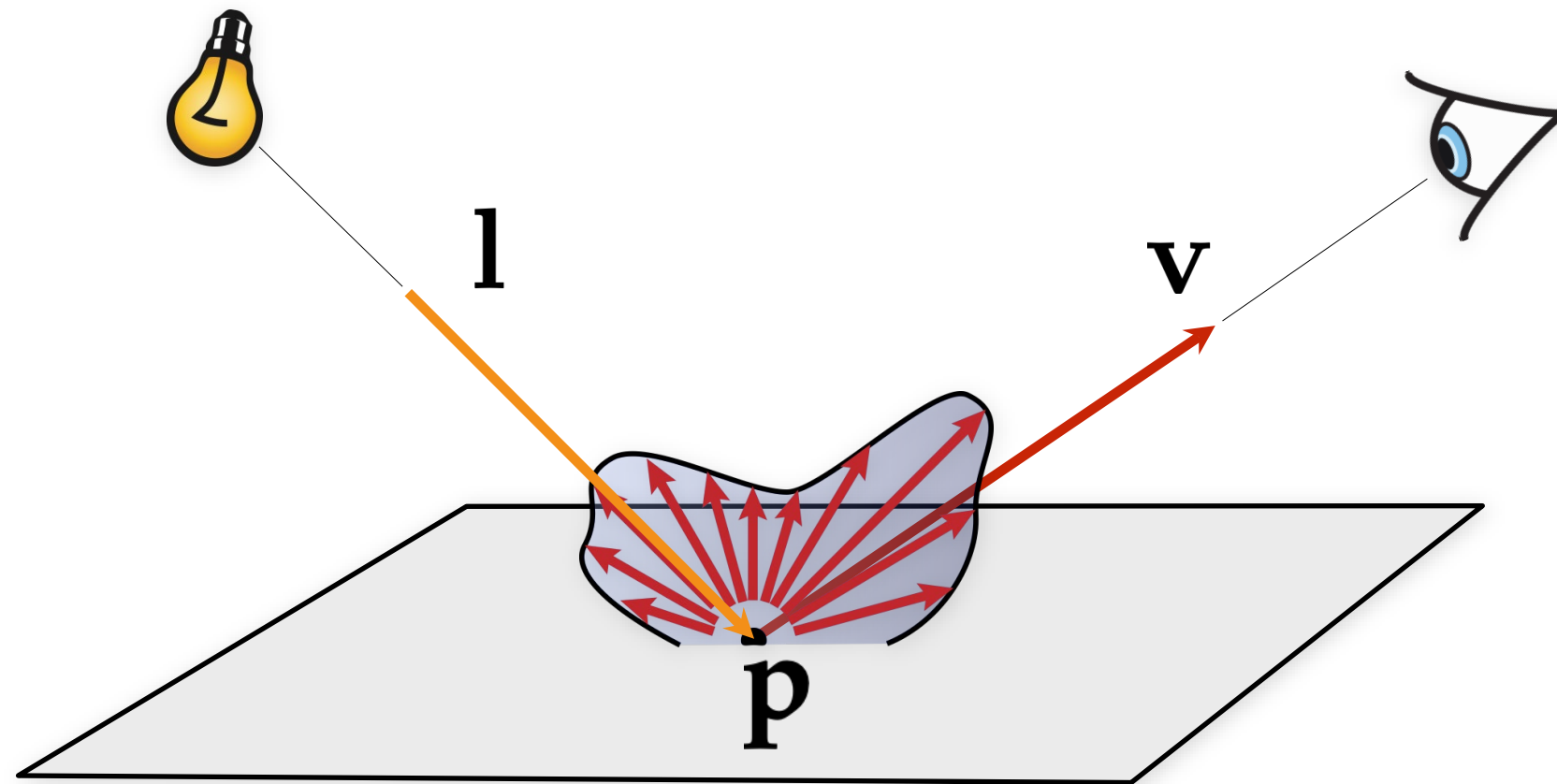


This describes a local illumination model

The BSDF

Bidirectional Scattering Distribution Function

- informally: how much the material scatters light coming from one direction \mathbf{l} into some other direction \mathbf{v} , at each point \mathbf{p}

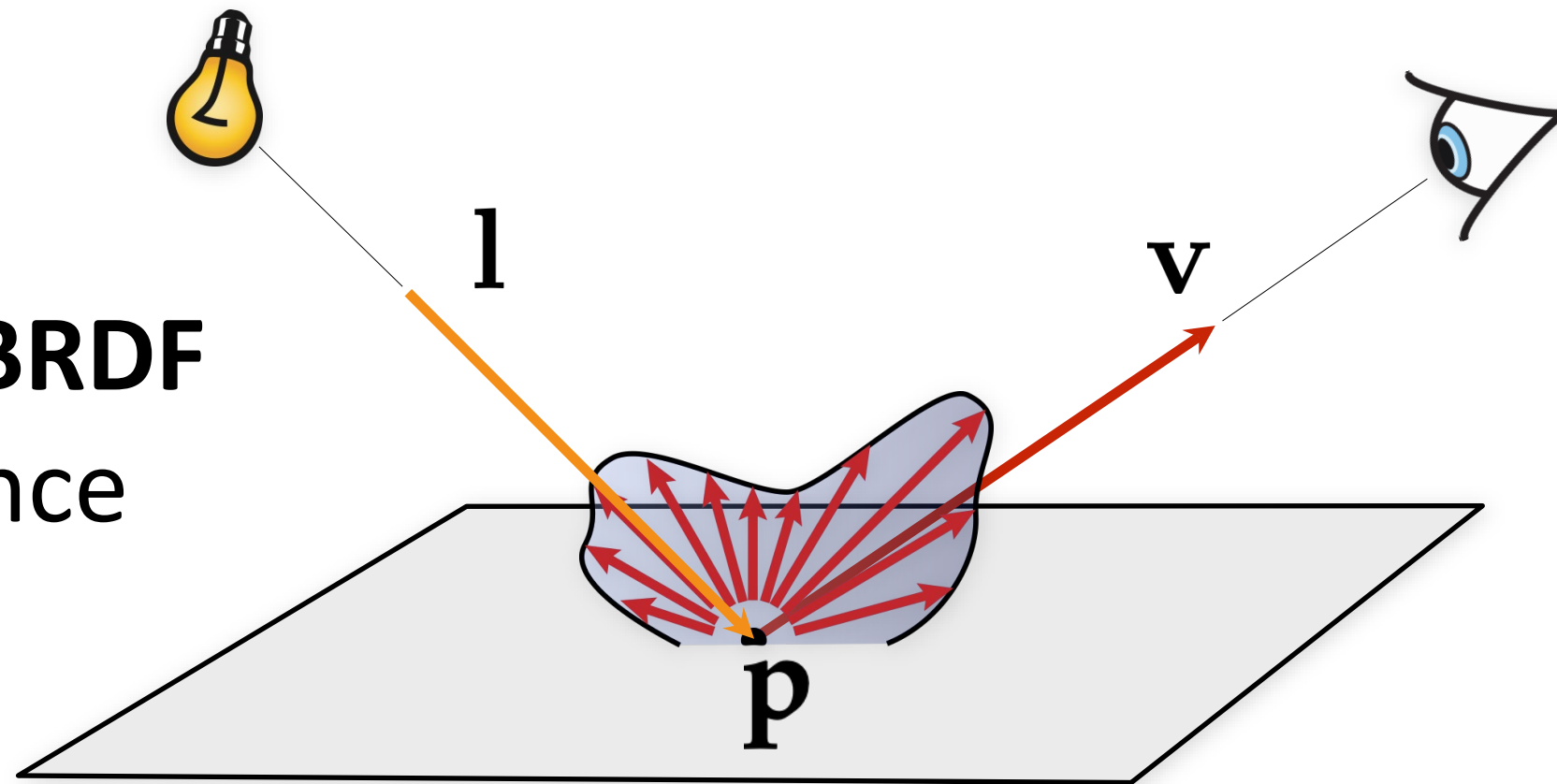


The BSDF

Bidirectional Scattering Distribution Function

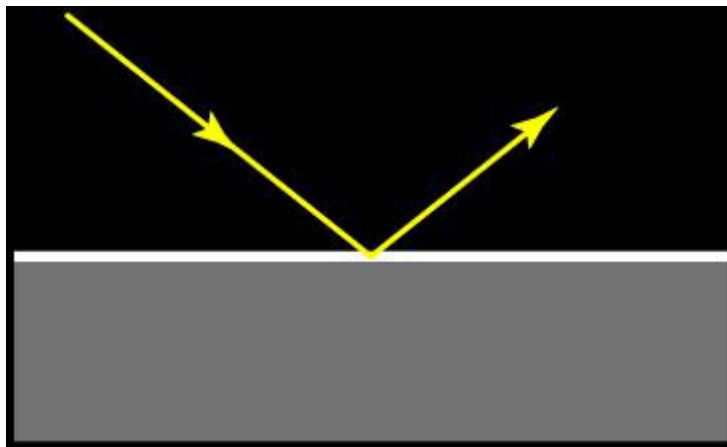
- informally: how much the material scatters light coming from one direction \mathbf{l} into some other direction \mathbf{v} , at each point \mathbf{p}

You'll also see **BRDF**
R for reflectance

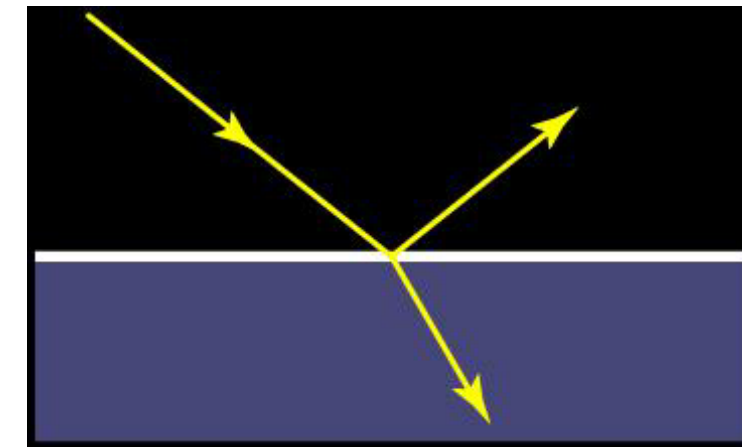


Real-world materials

Metals

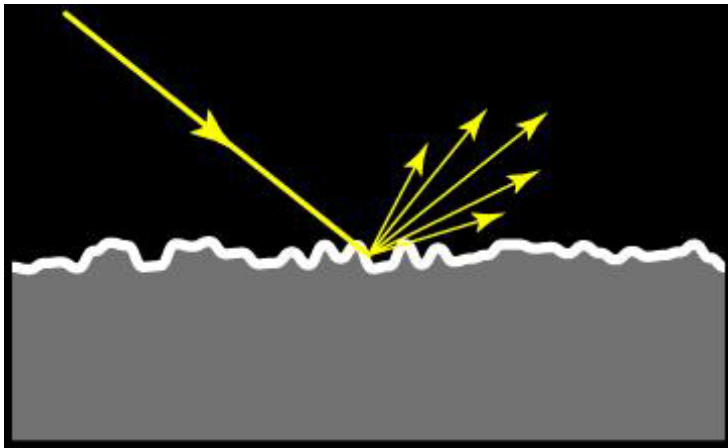


Dielectric

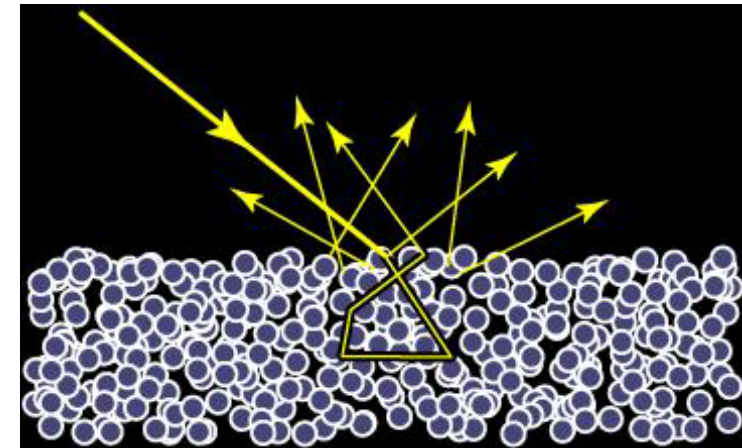


Real-world materials

Metals



Dielectric



Idealized material models

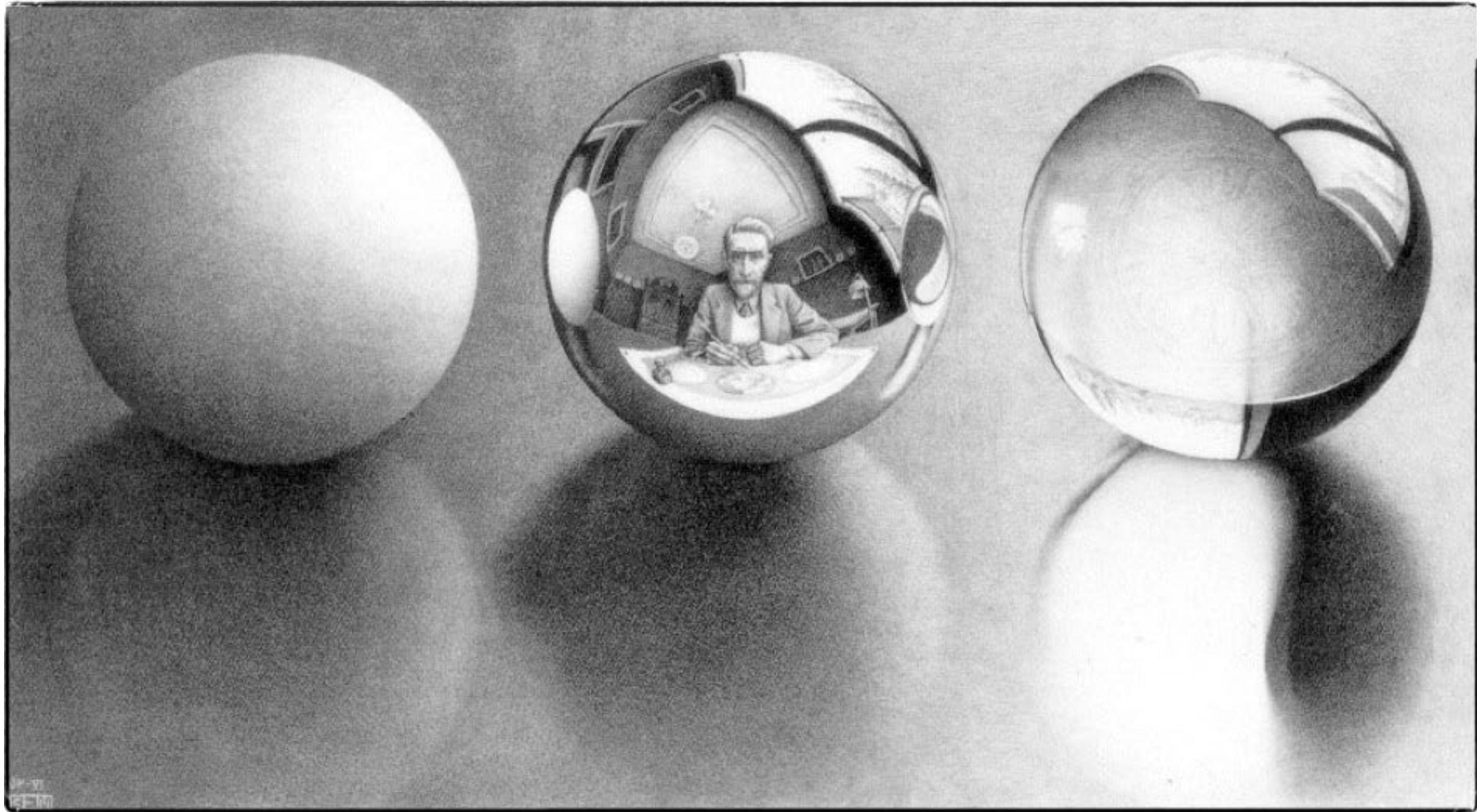
Diffuse reflection

- light is reflected in all directions
- colored by surface color

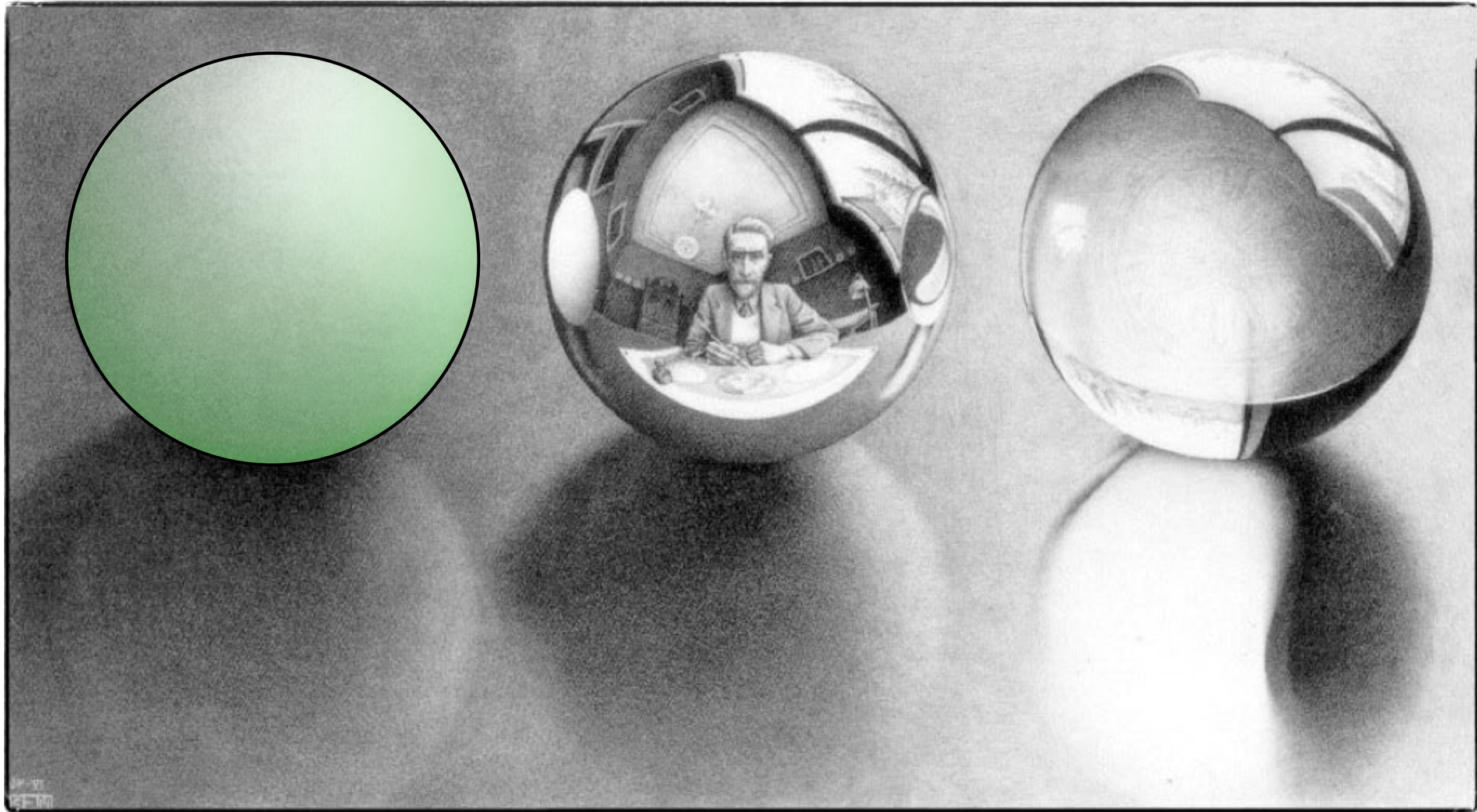
Smooth specular reflection/refraction (e.g., chrome, glass, glaze/varnish)

- light reflected/refracted only in a single direction
- colored by source color

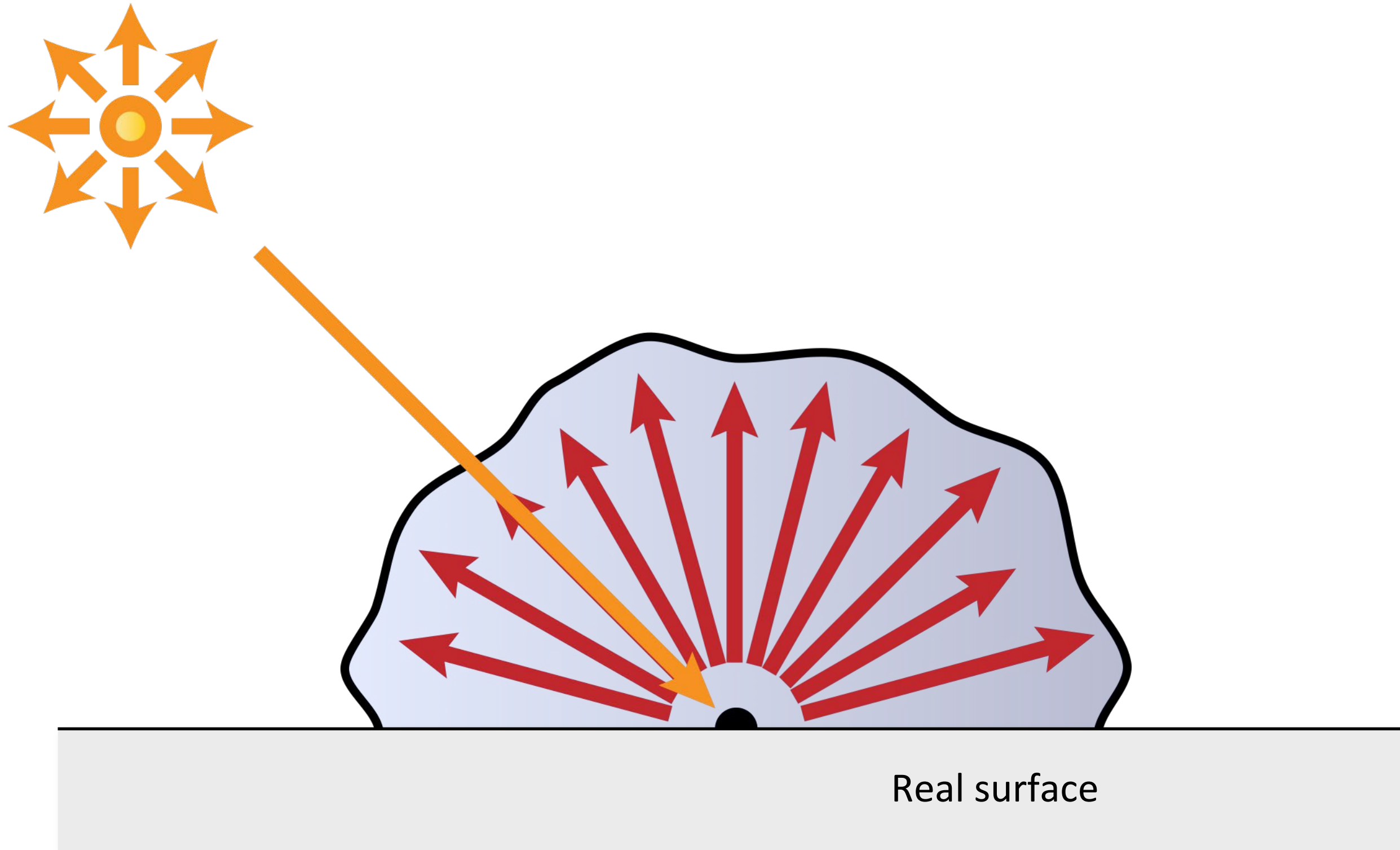
Idealized materials



Diffuse reflection



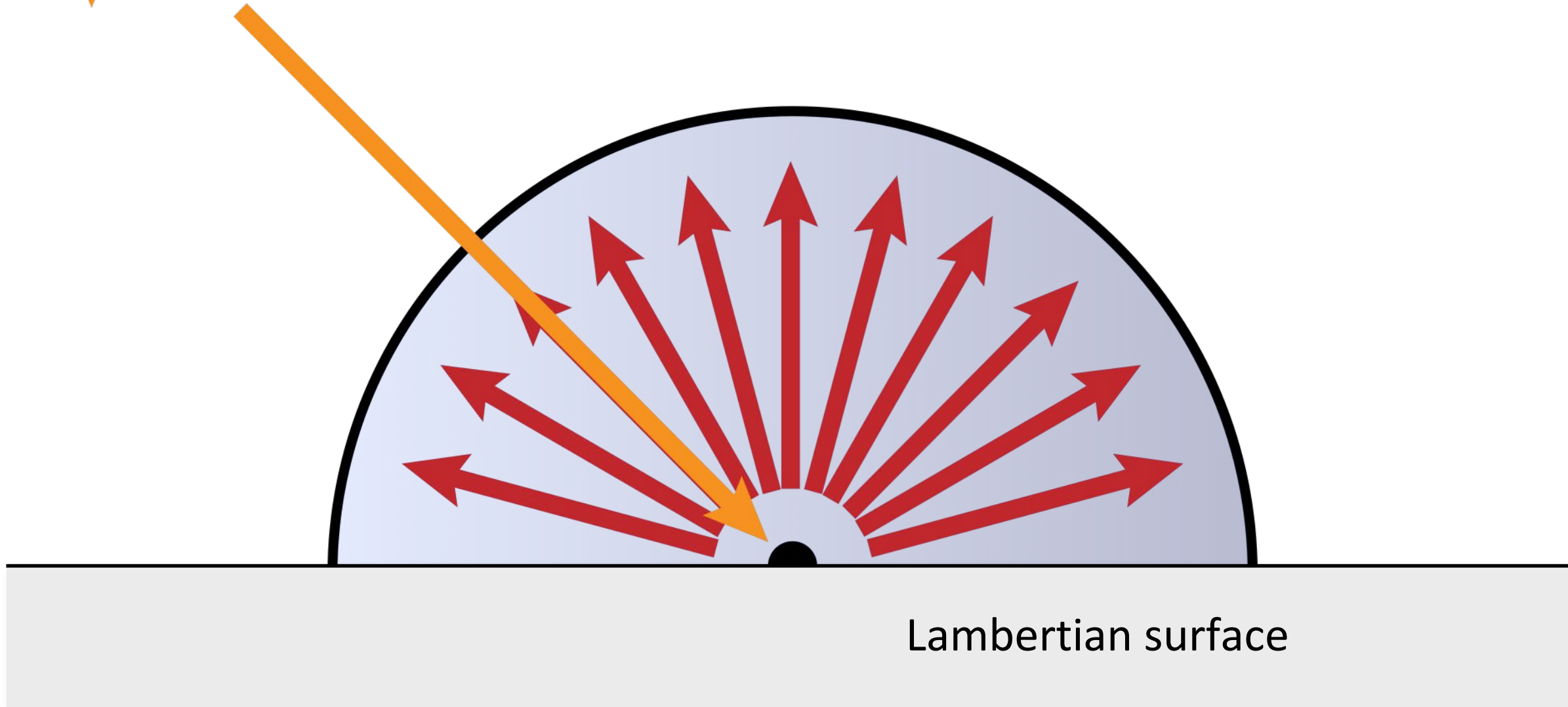
Diffuse reflection



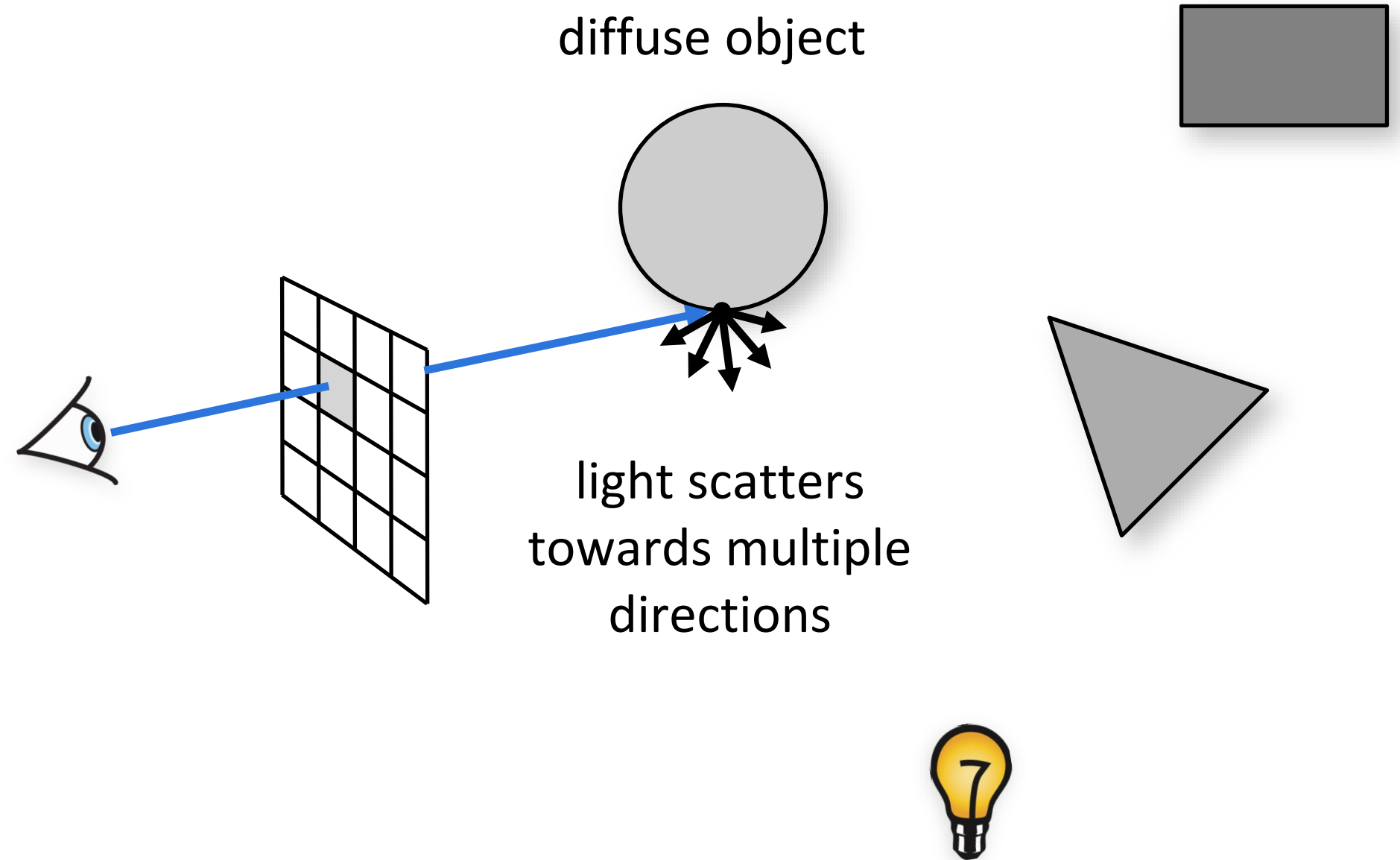
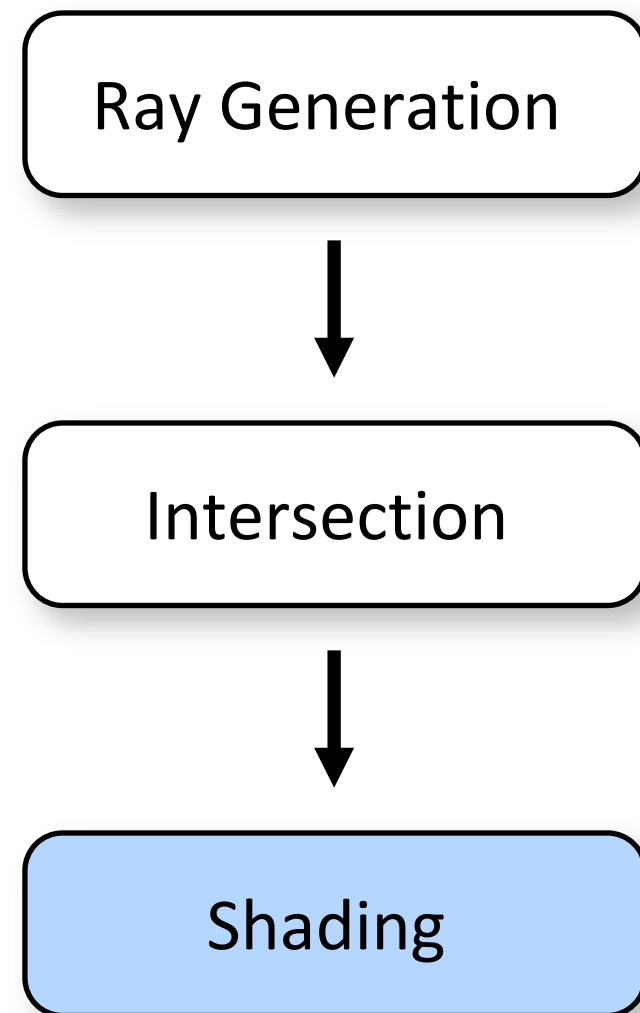
Lambertian reflection



Also called ideal diffuse reflection

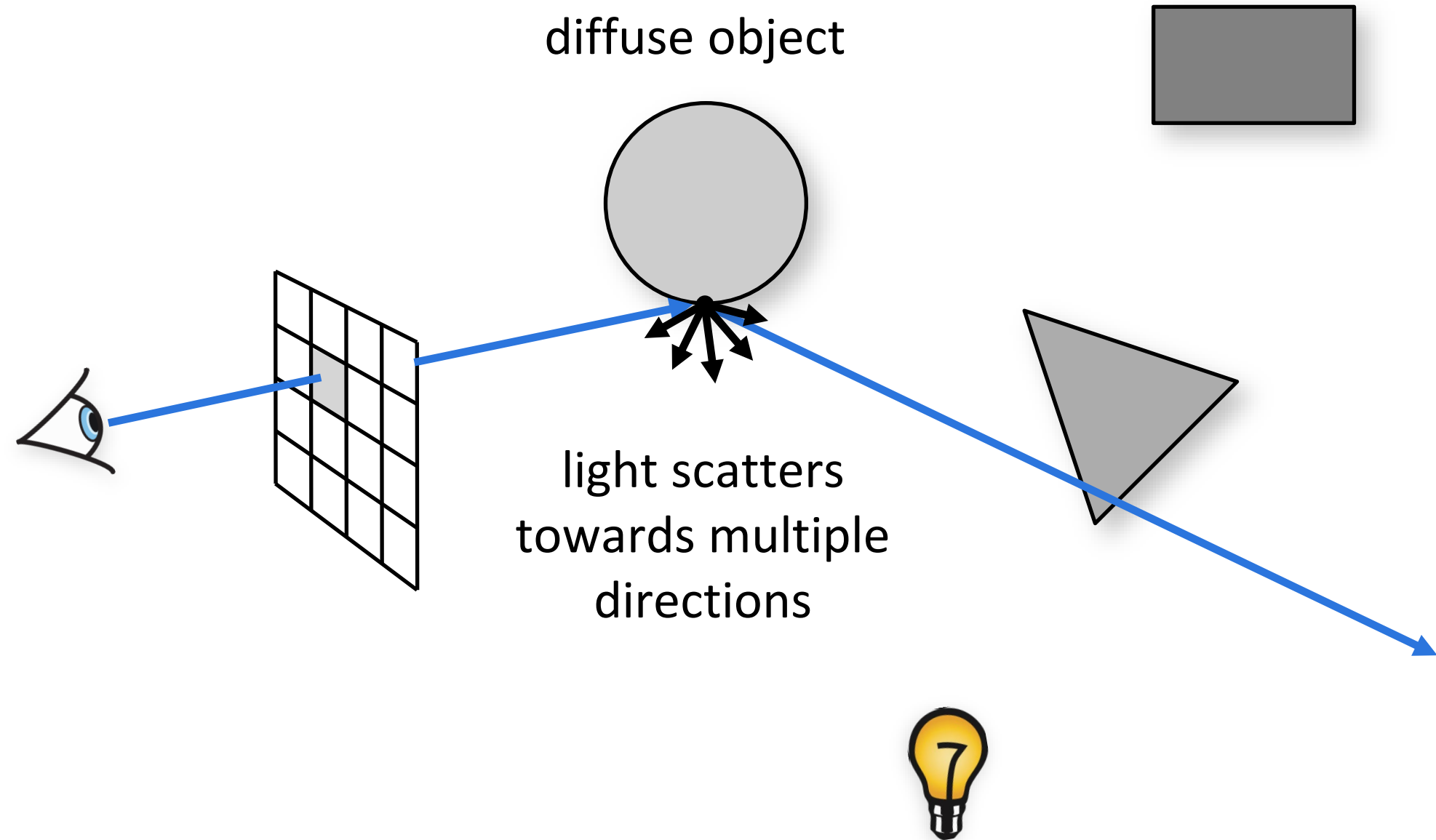
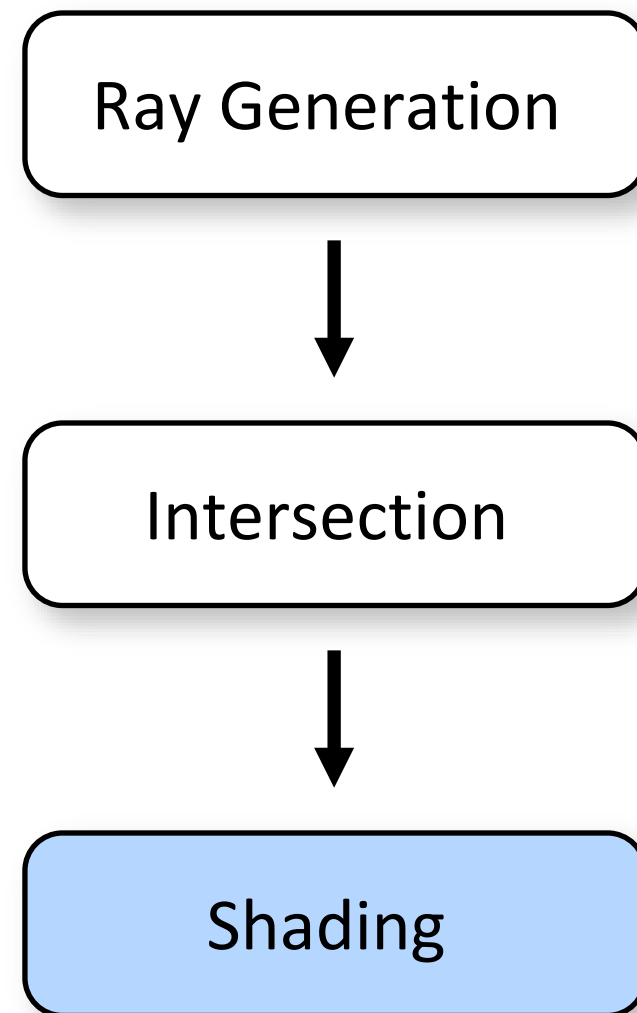


Basic Ray Tracing Pipeline



What direction should we trace a new ray towards?

Basic Ray Tracing Pipeline



What direction should we trace a new ray towards?

- Pick a direction at random!

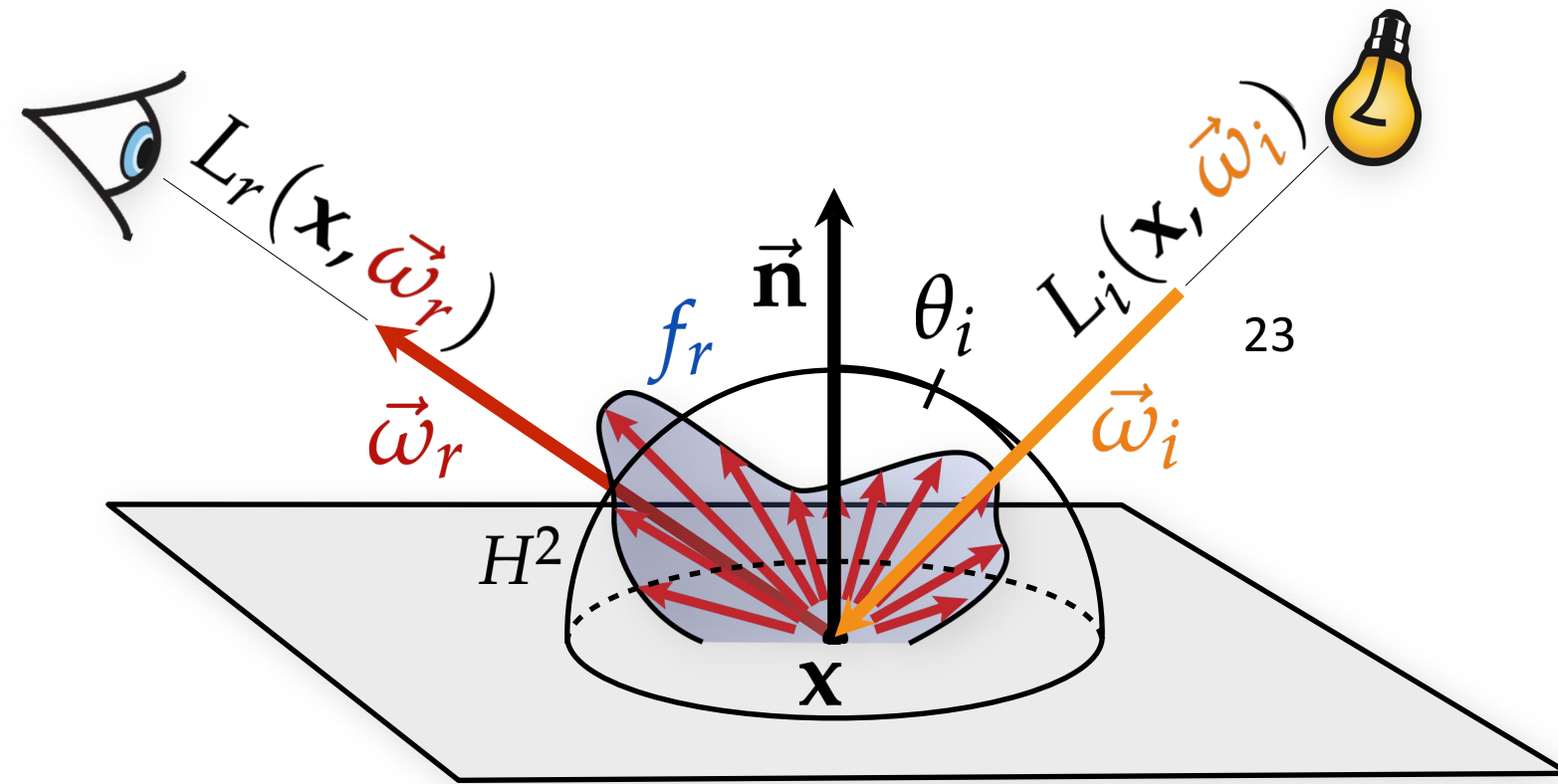
Sampling Lambertian scattering

From what distribution should we sample directions?

The reflection equation

Reflected radiance is a (hemi)spherical integral of incident radiance from all directions

$$L_r(\mathbf{x}, \vec{\omega}_r) = \int_{H^2} f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i$$



This describes a local illumination model

Sampling Lambertian scattering

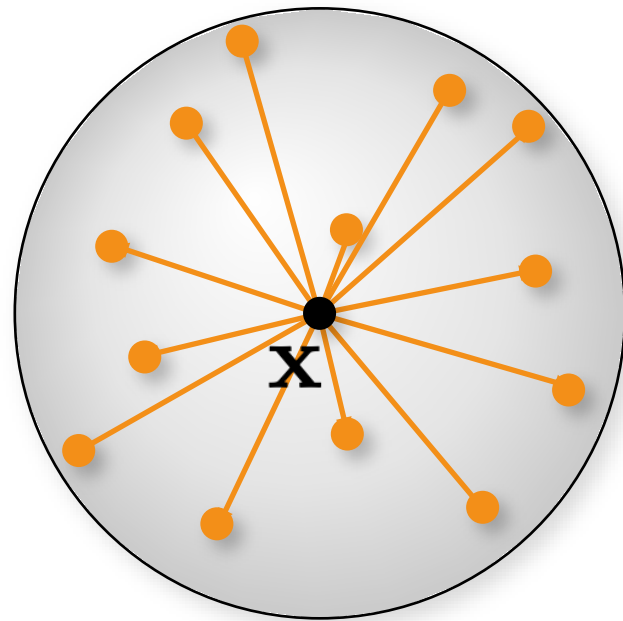
From what distribution should we sample directions?

- Probability proportional to $\cos(\hat{n} \cdot \hat{\omega})$.
- Even though BSDF scatters to all directions uniformly, we need to account for foreshortening.

How do we sample directions based on this *cosine-weighted* distribution?

Sampling Lambertian scattering

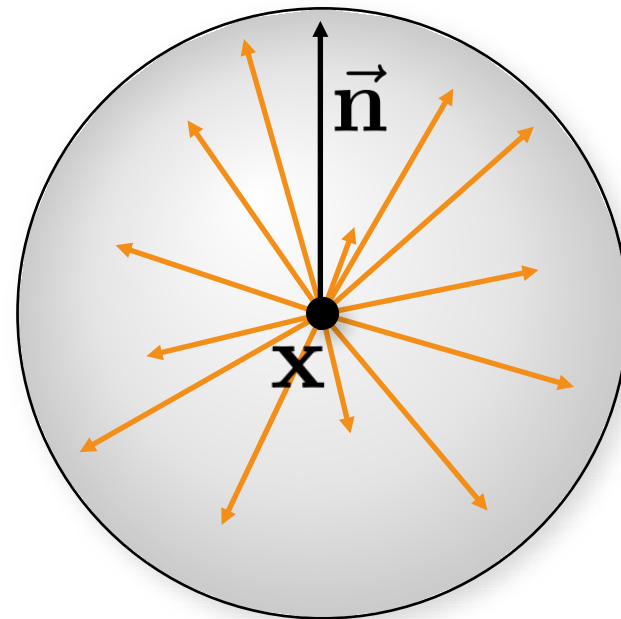
Generate points
uniformly on sphere
(unit directions)



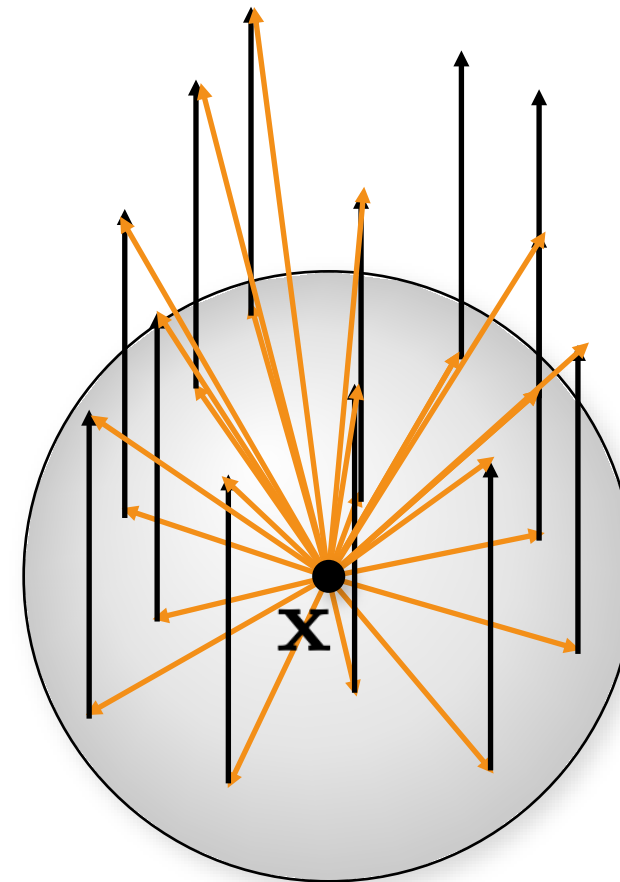
Sampling Lambertian scattering

Generate points
uniformly on sphere
(unit directions)

unit normal



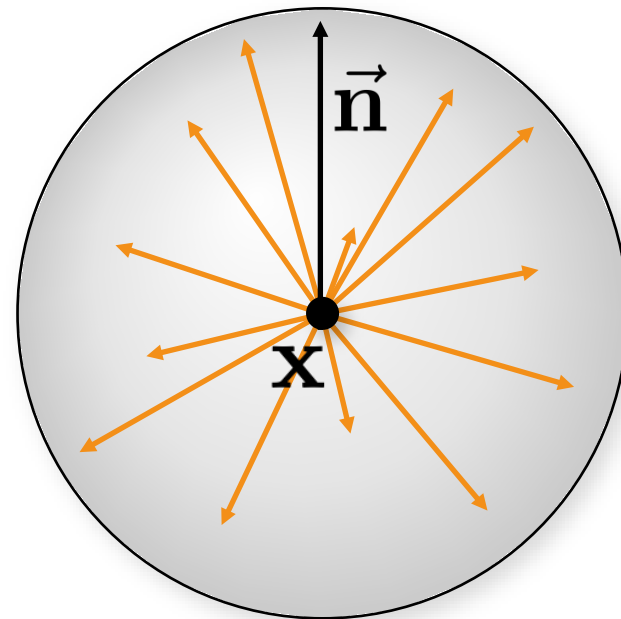
Add unit normal



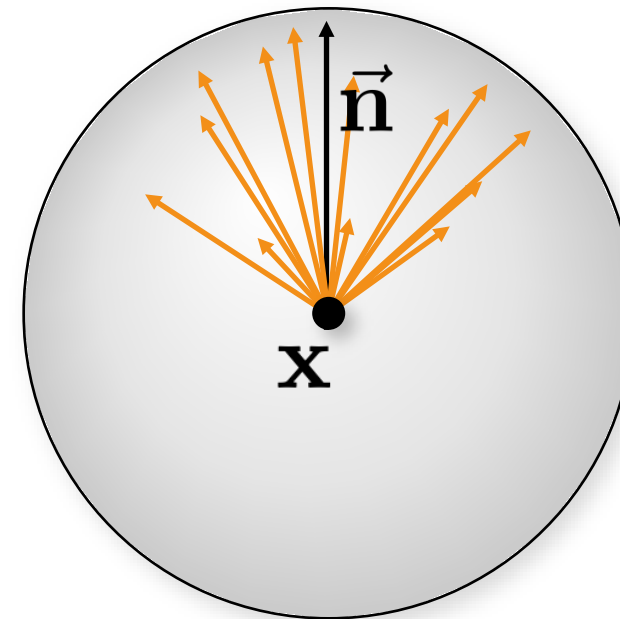
Sampling Lambertian scattering

Generate points
uniformly on sphere
(unit directions)

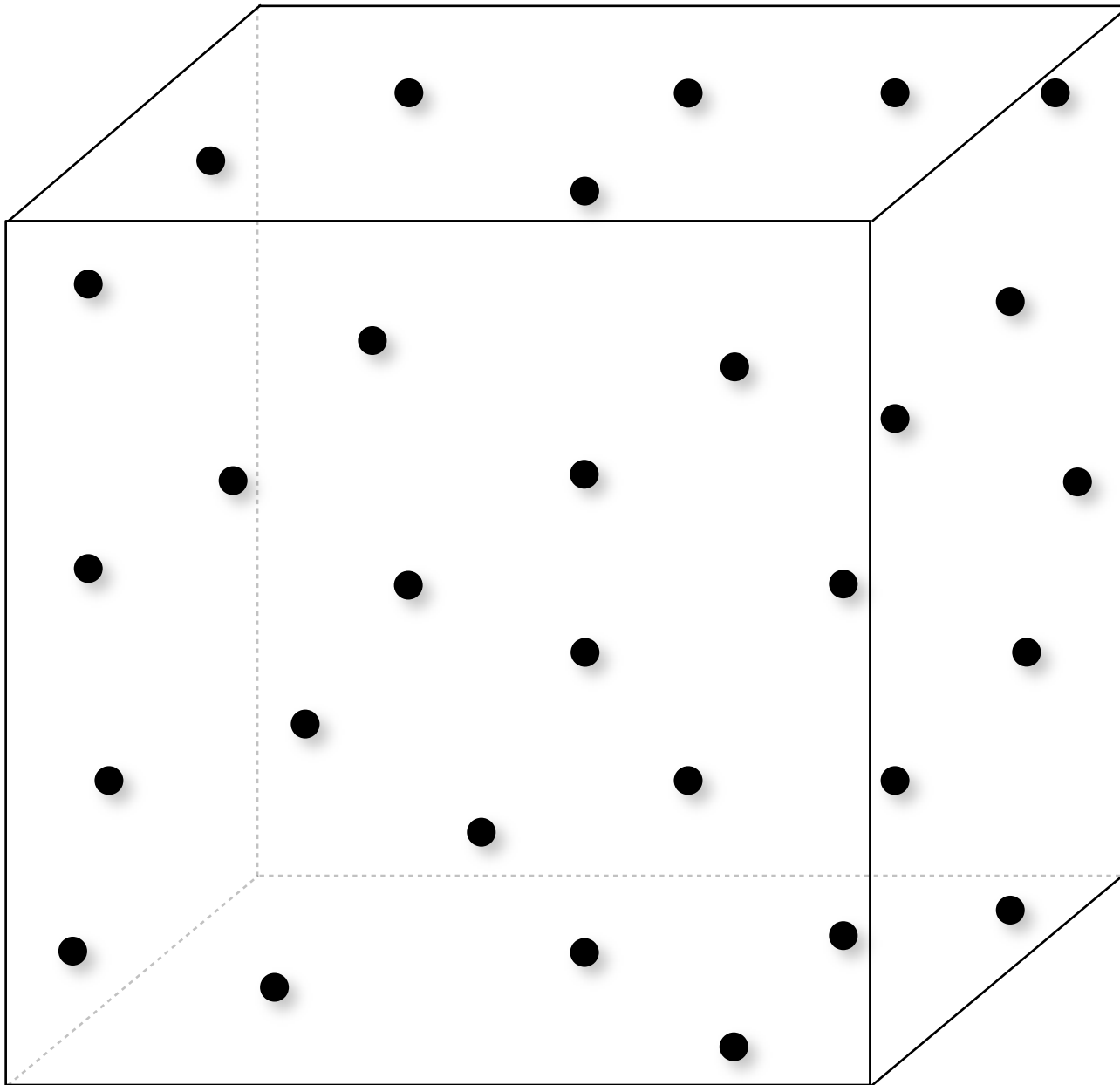
unit normal



Add unit normal
normalize



Rejection Sampling a Sphere



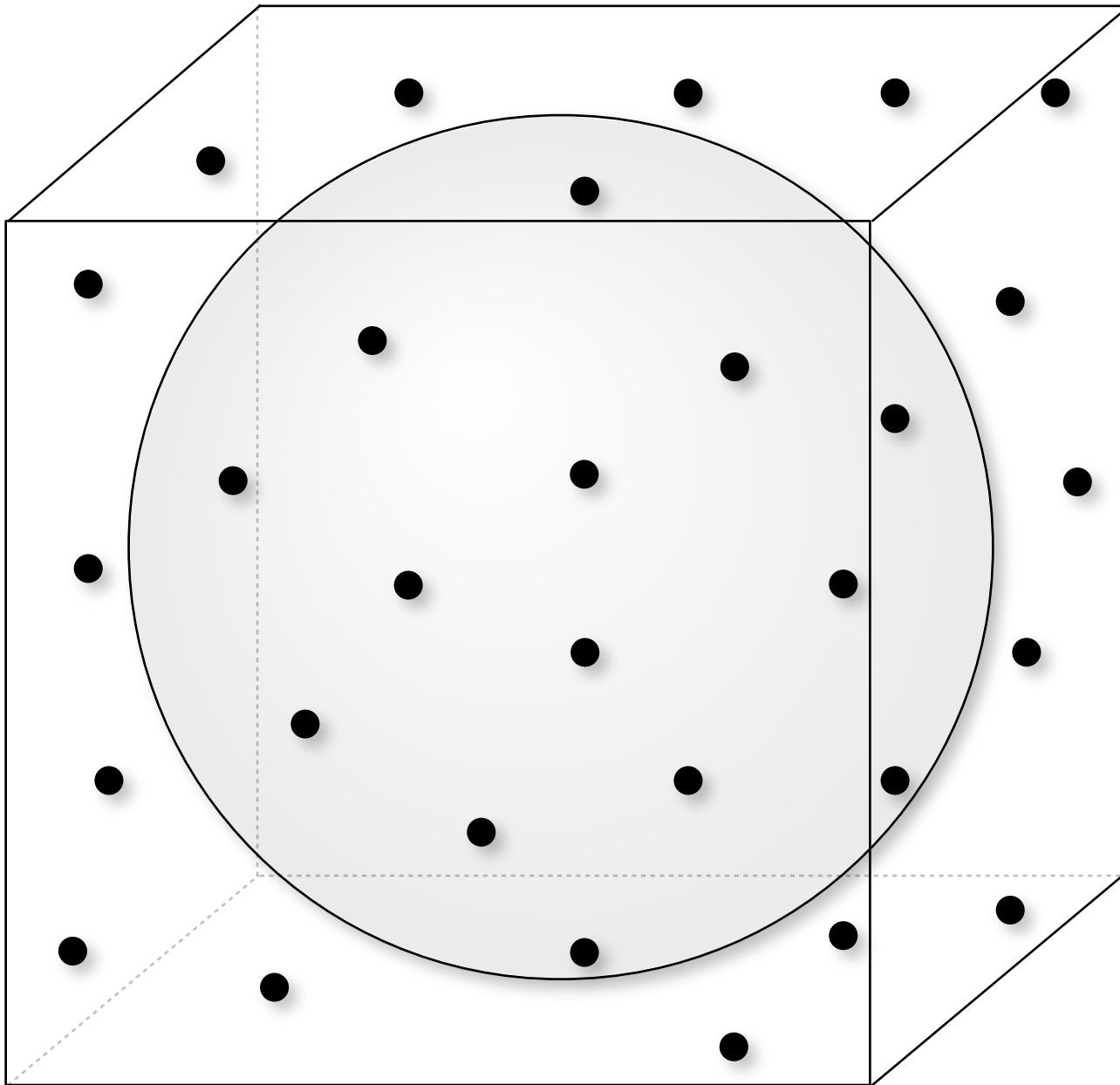
```
Vector3D v;
```

```
v.x = 1-2*randf();
```

```
v.y = 1-2*randf();
```

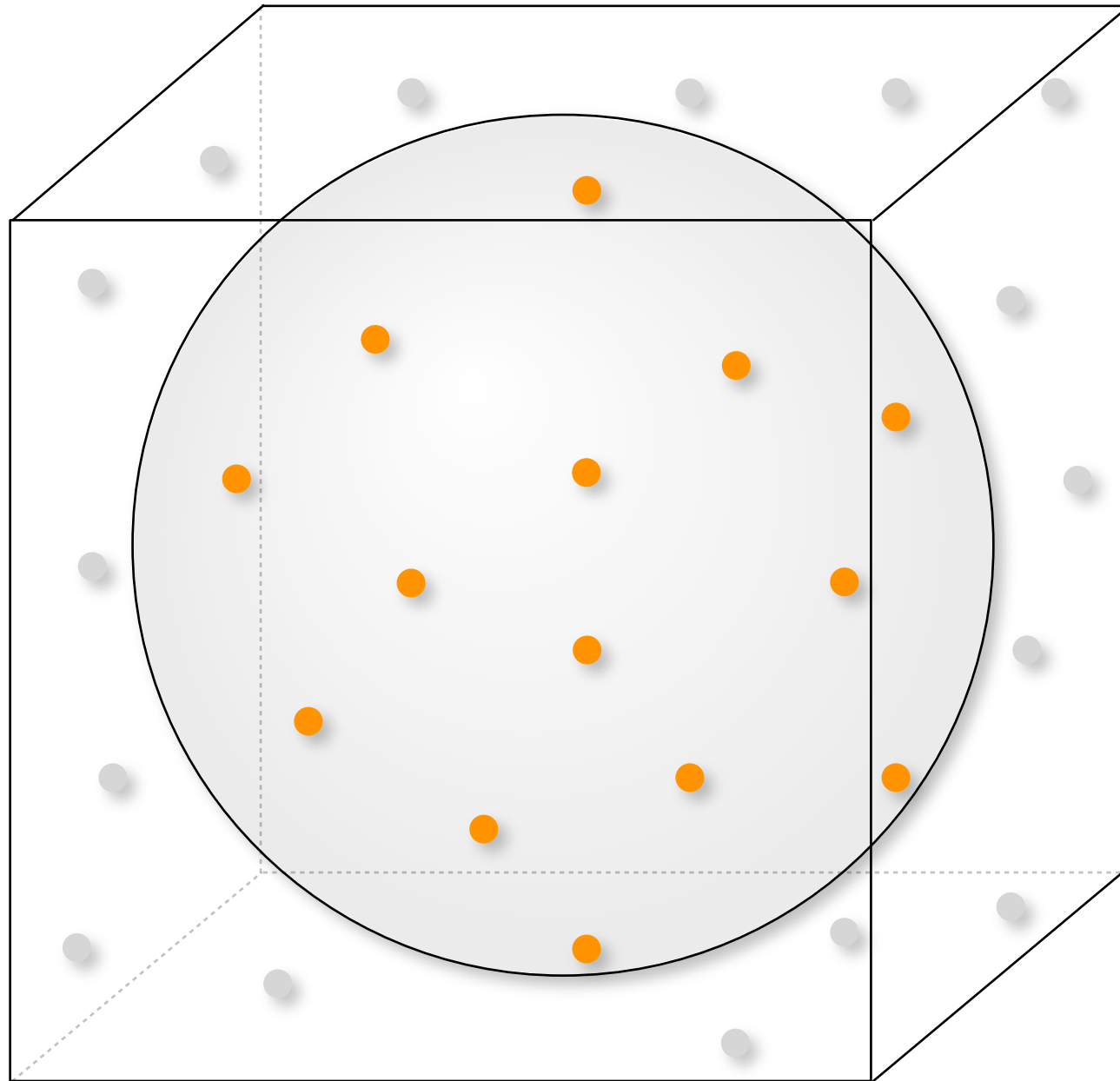
```
v.z = 1-2*randf();
```


Rejection Sampling a Sphere



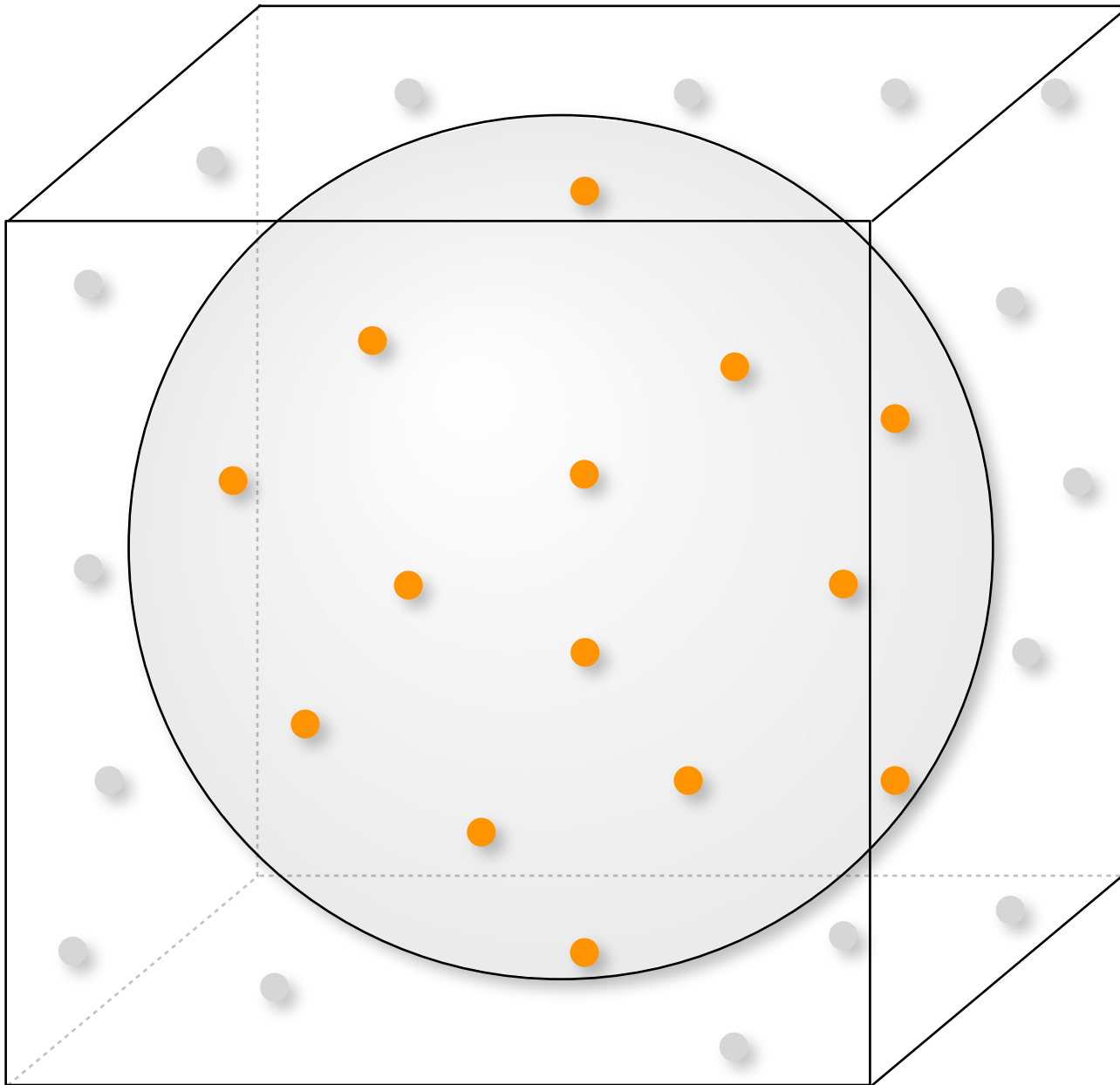
```
Vector3D v;  
do  
{  
    v.x = 1-2*randf();  
    v.y = 1-2*randf();  
    v.z = 1-2*randf();  
} while(v.length2() > 1)
```

Rejection Sampling a Sphere



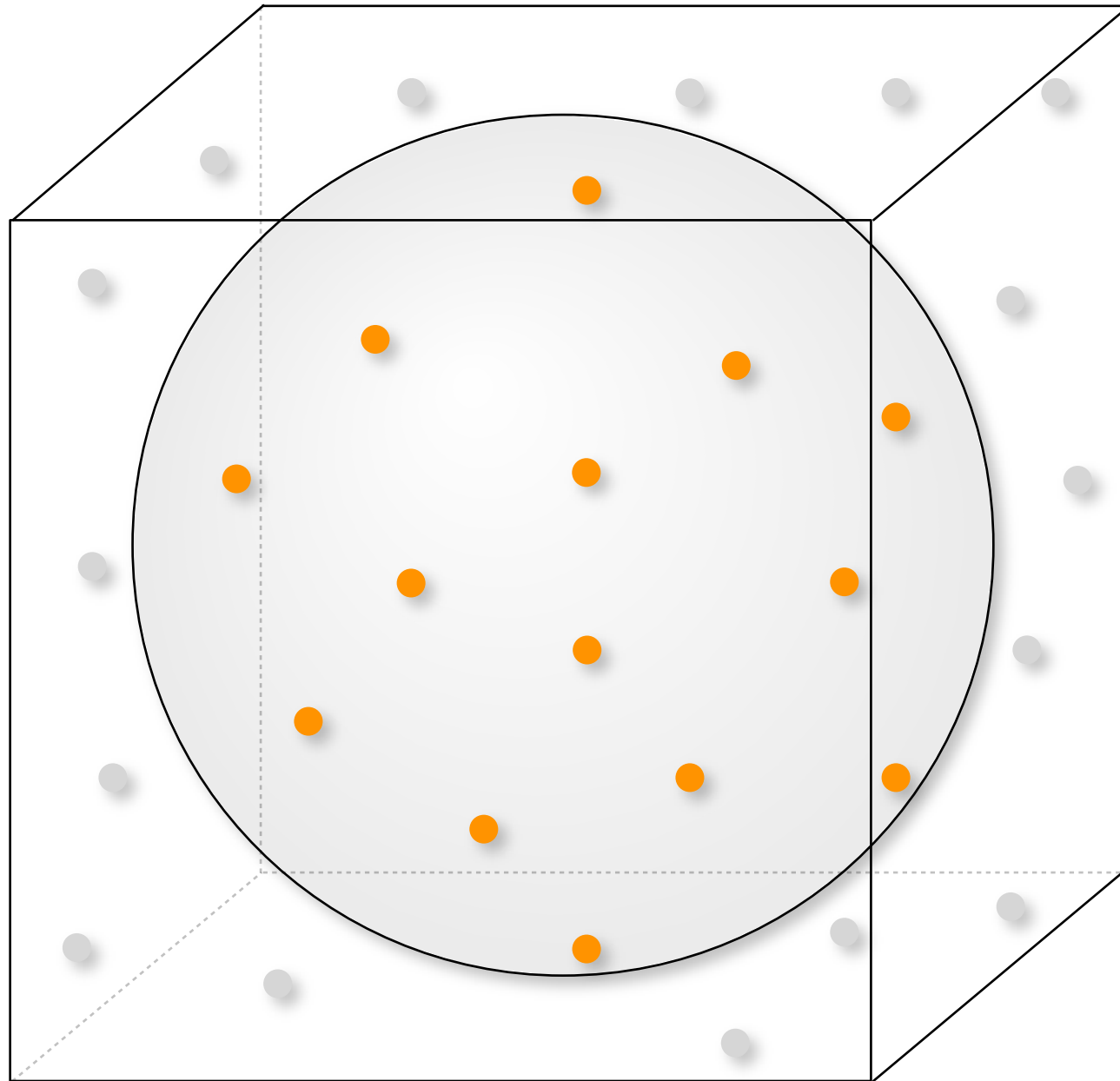
```
Vector3D v;  
do  
{  
    v.x = 1-2*randf();  
    v.y = 1-2*randf();  
    v.z = 1-2*randf();  
} while(v.length2() > 1)
```

Rejection Sampling a Sphere



```
Vector3D v;  
do  
{  
    v.x = 1-2*randf();  
    v.y = 1-2*randf();  
    v.z = 1-2*randf();  
} while(v.length2() > 1)  
// Project onto sphere  
v /= v.length();
```

Sampling a Sphere using normal samples



```
Vector3D v;  
v.x = randnf();  
v.y = randnf();  
v.z = randnf();  
// Project onto sphere  
v /= v.length();
```

- No rejection sampling required (no `while` loop).
- Need to use normal, rather than uniform, samples.

Accounting for Lambertian albedo

At each diffuse shading event, you also need to multiply by the diffuse albedo (between 0 and 1).

Diffuse shading

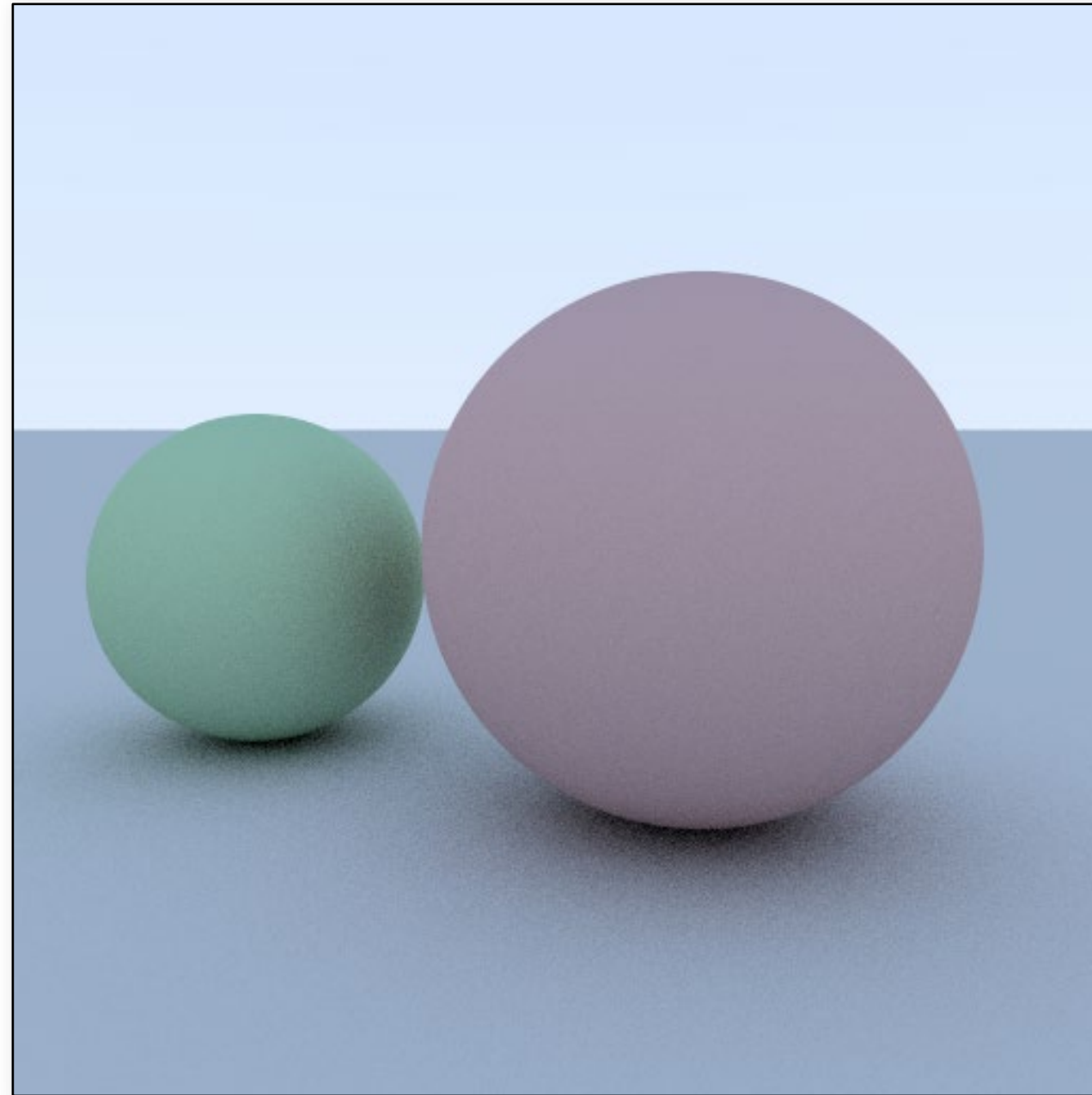
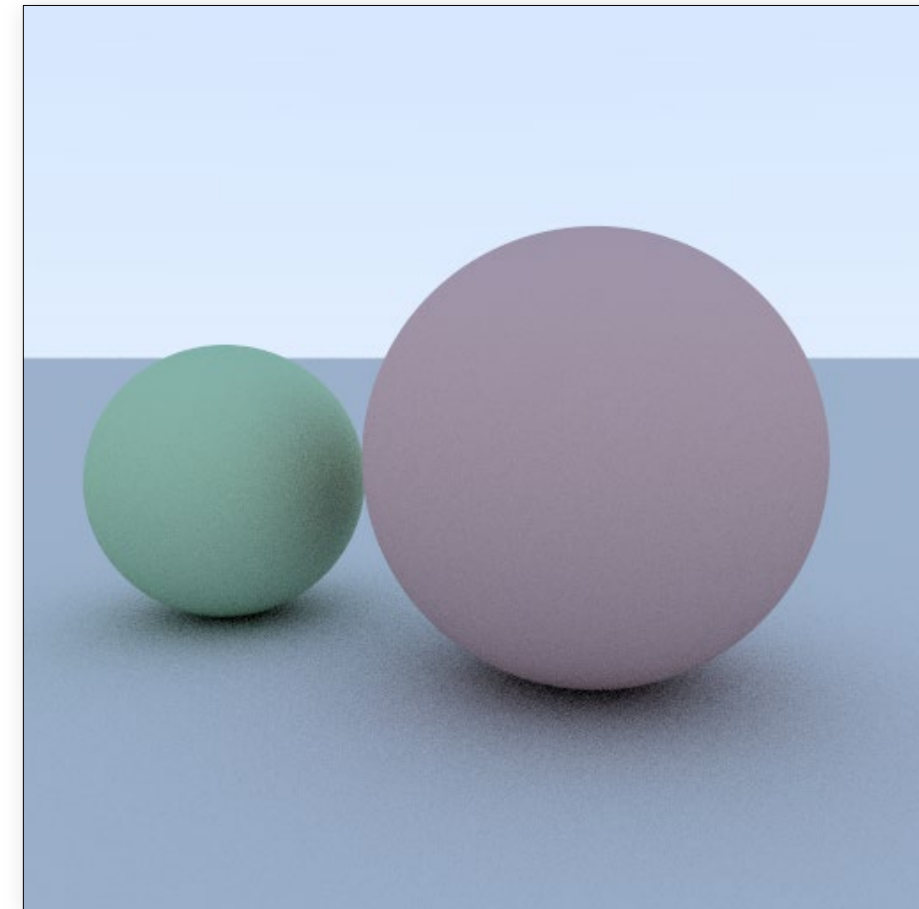


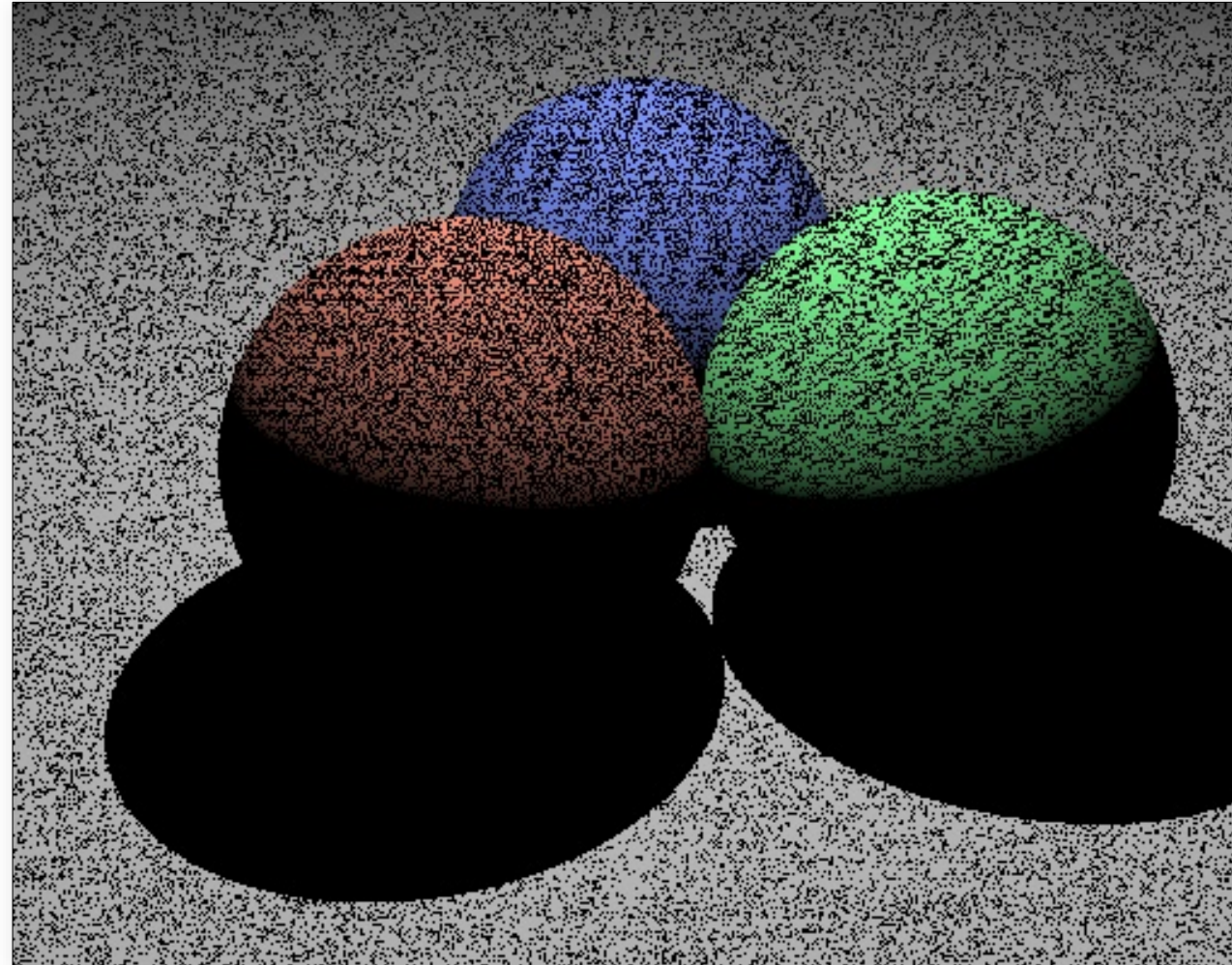
Image so far

```
Scene::trace(Ray ray)
    hit = surfaces.intersect(ray);
    if hit
        [col, sRay] = hit->mat->scatter(ray)
        return col * trace(sRay);
    else
        return backgroundColor;
```



Rounding errors

Don't fall victim to one of the classic blunders:

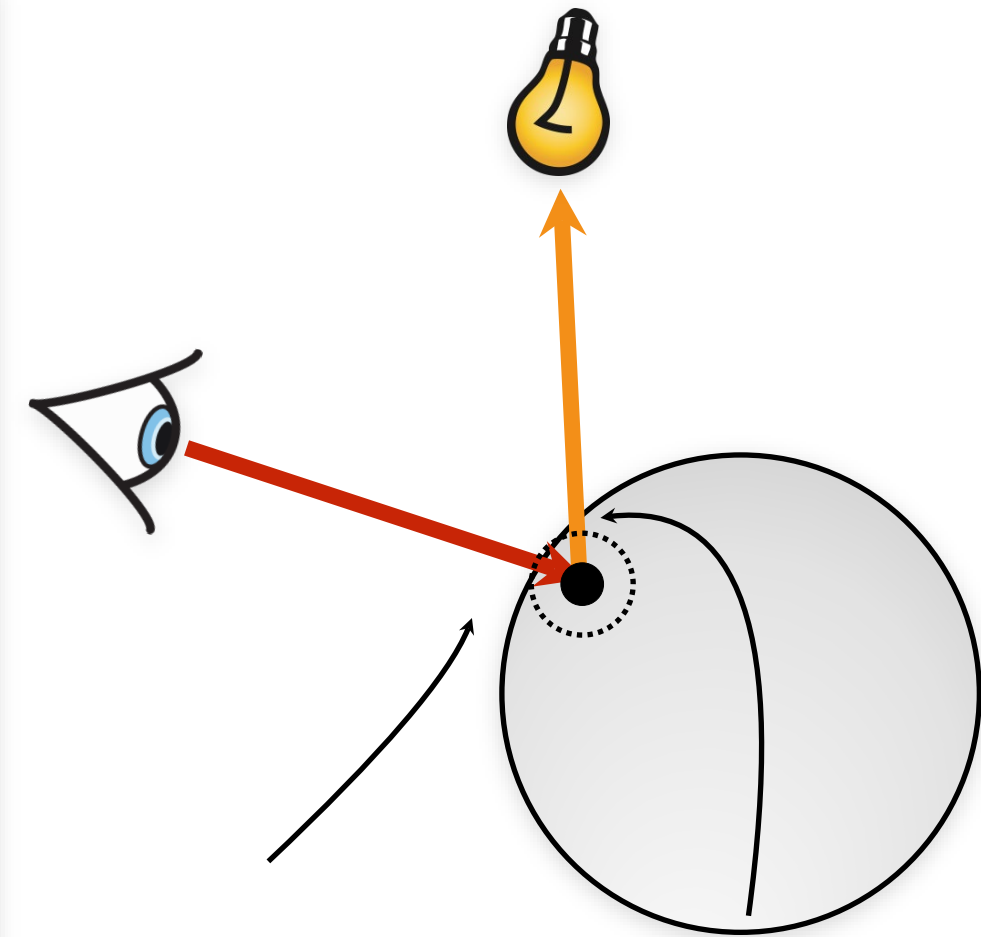
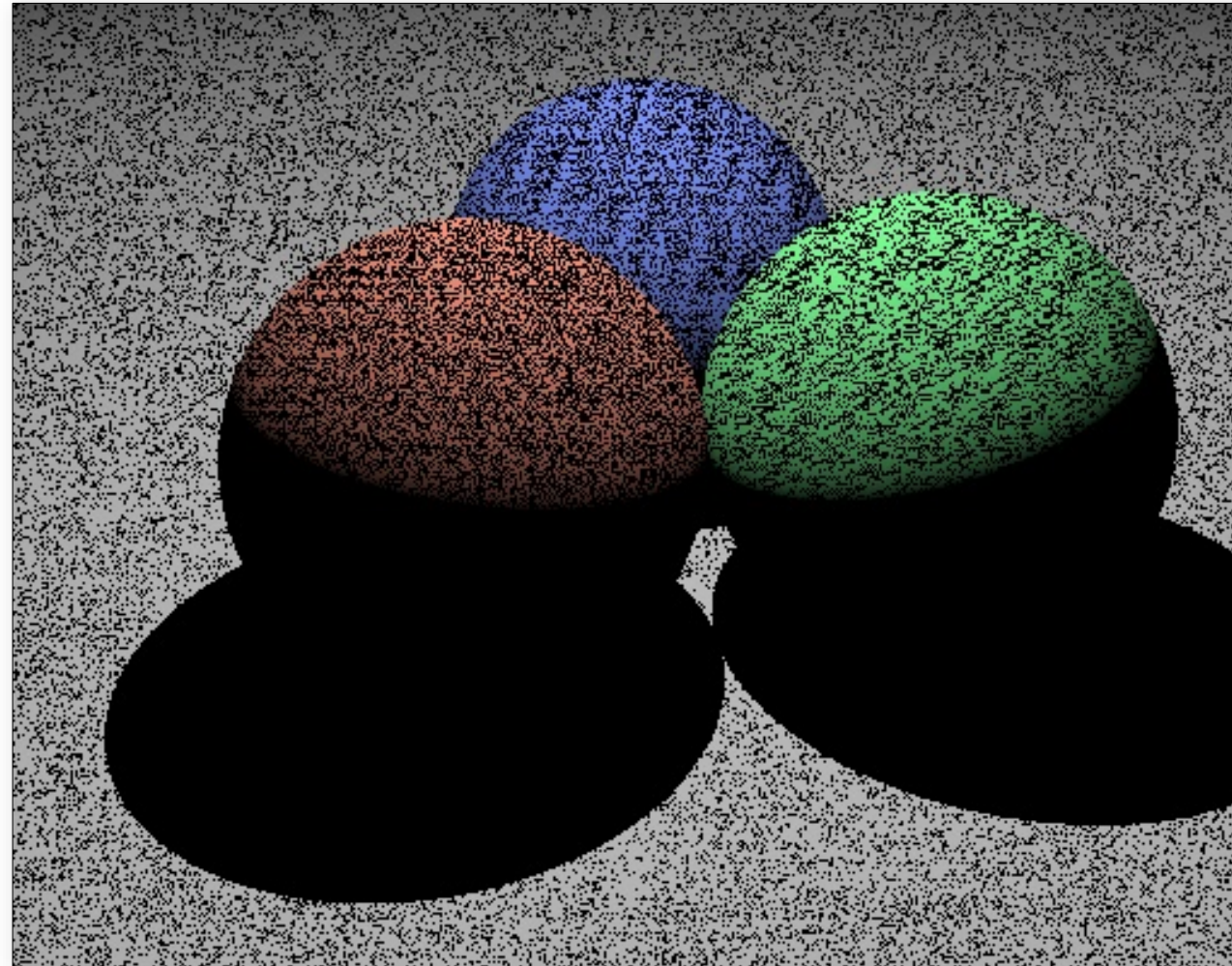
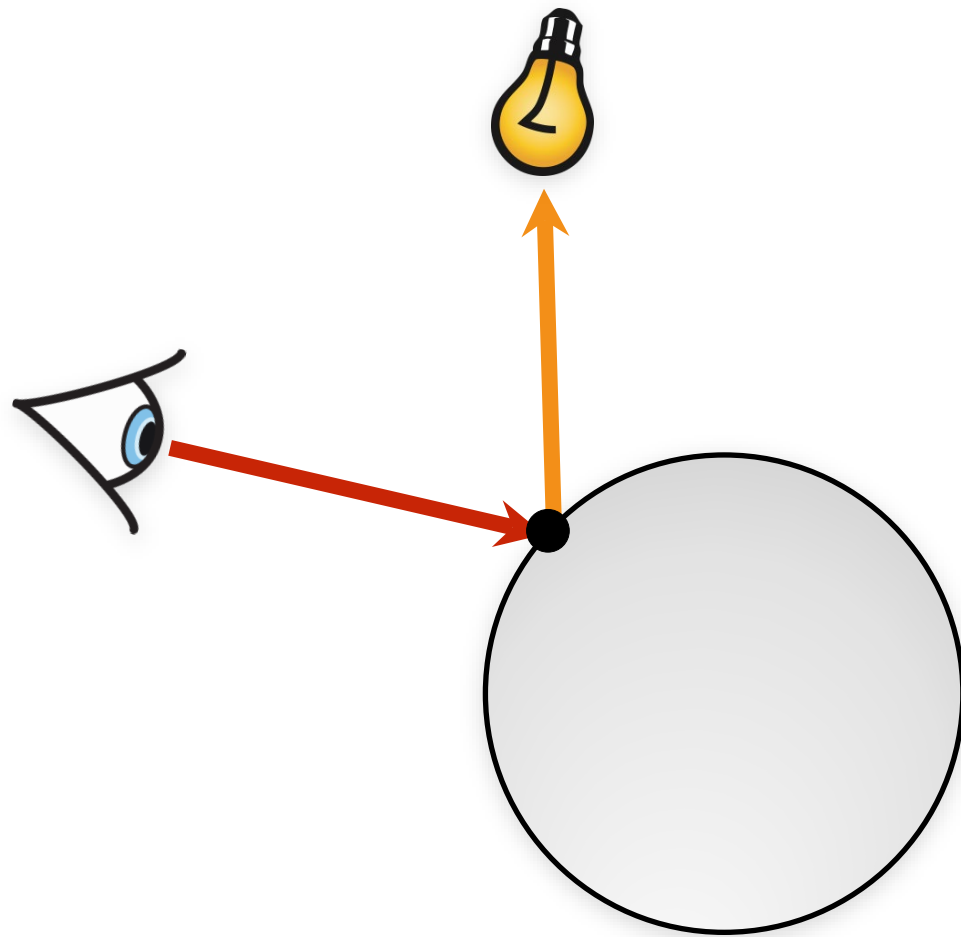


affectionately called
“shadow acne”

What's going on?

Rounding errors

Don't fall victim to one of the classic blunders:

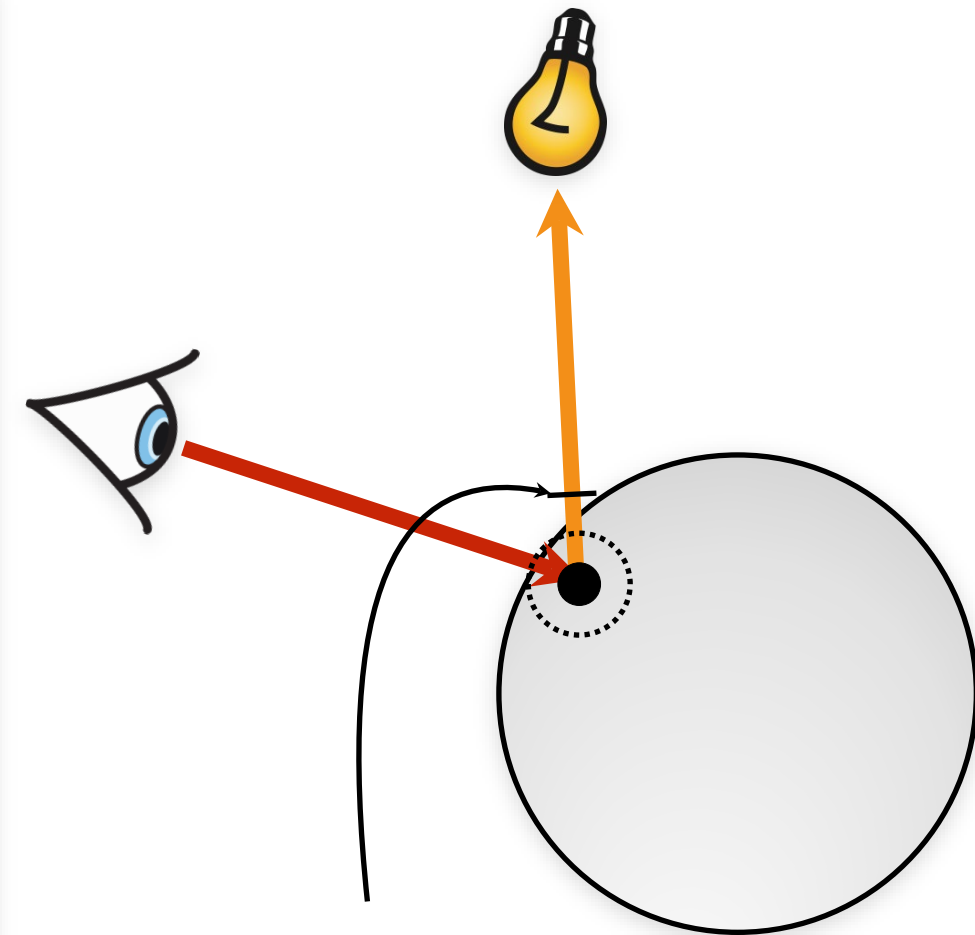
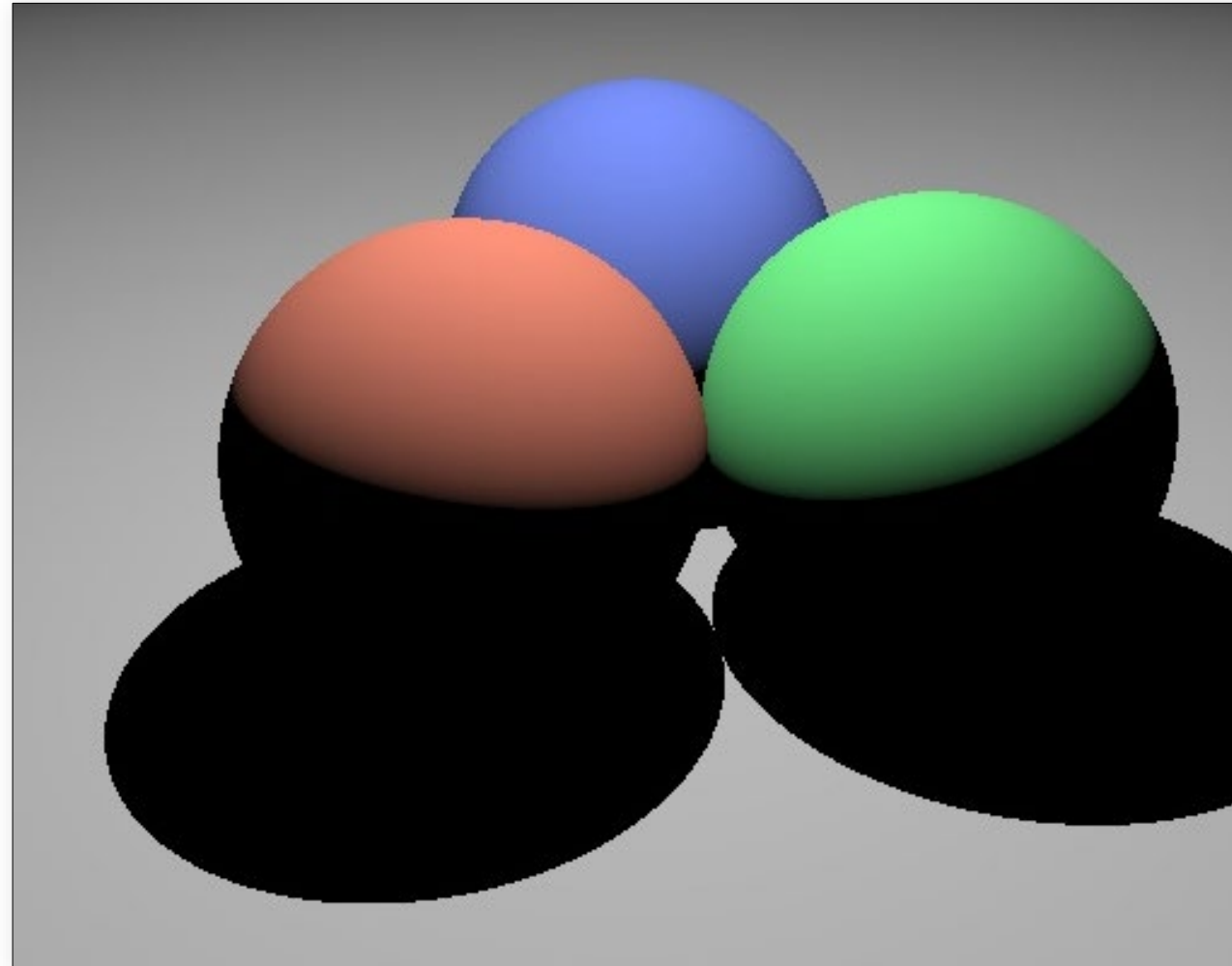


intersection inside
due to floating point
error

ray blocked by
self-intersection

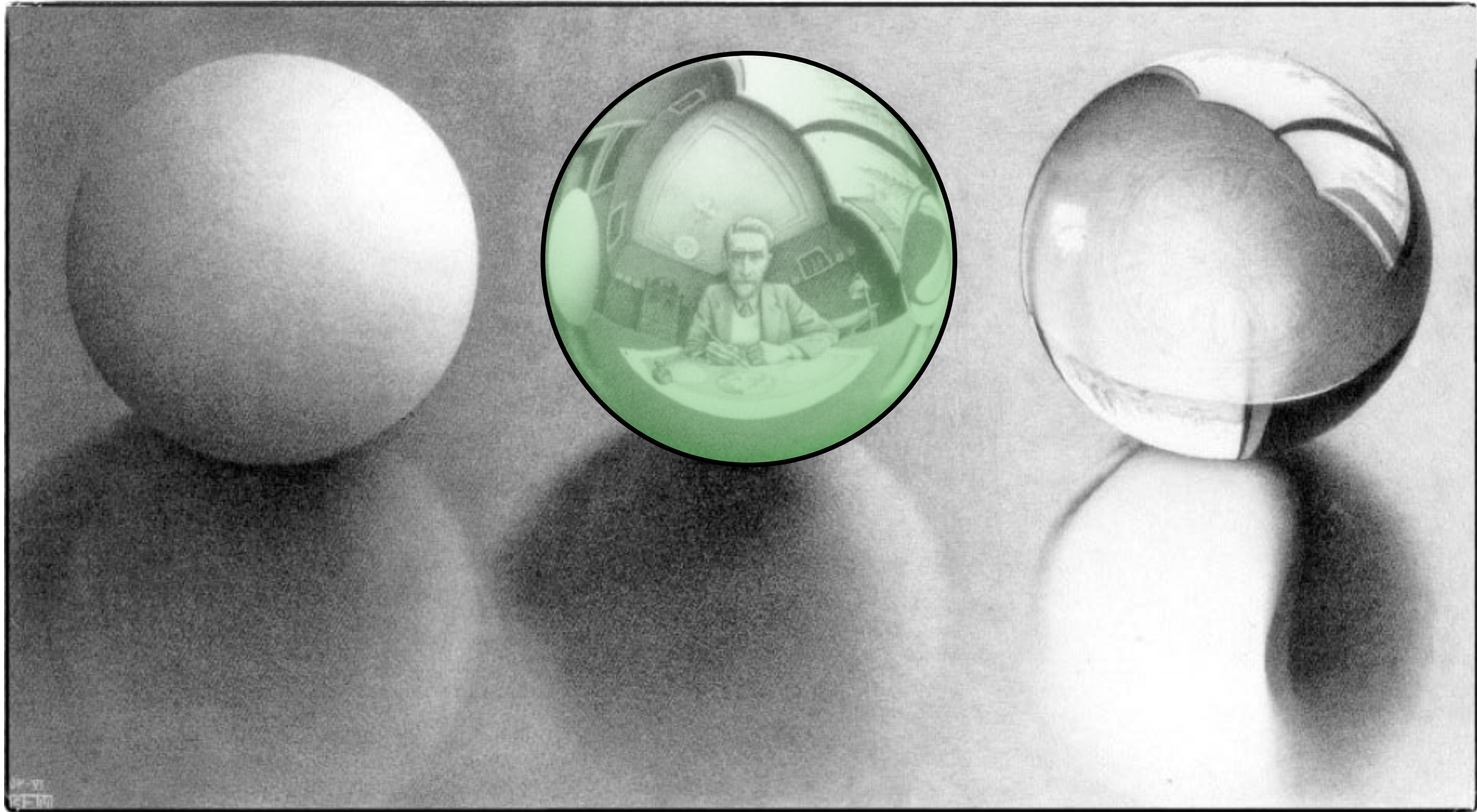
Shadow rounding errors

Solution: recursive rays start a tiny distance from the surface

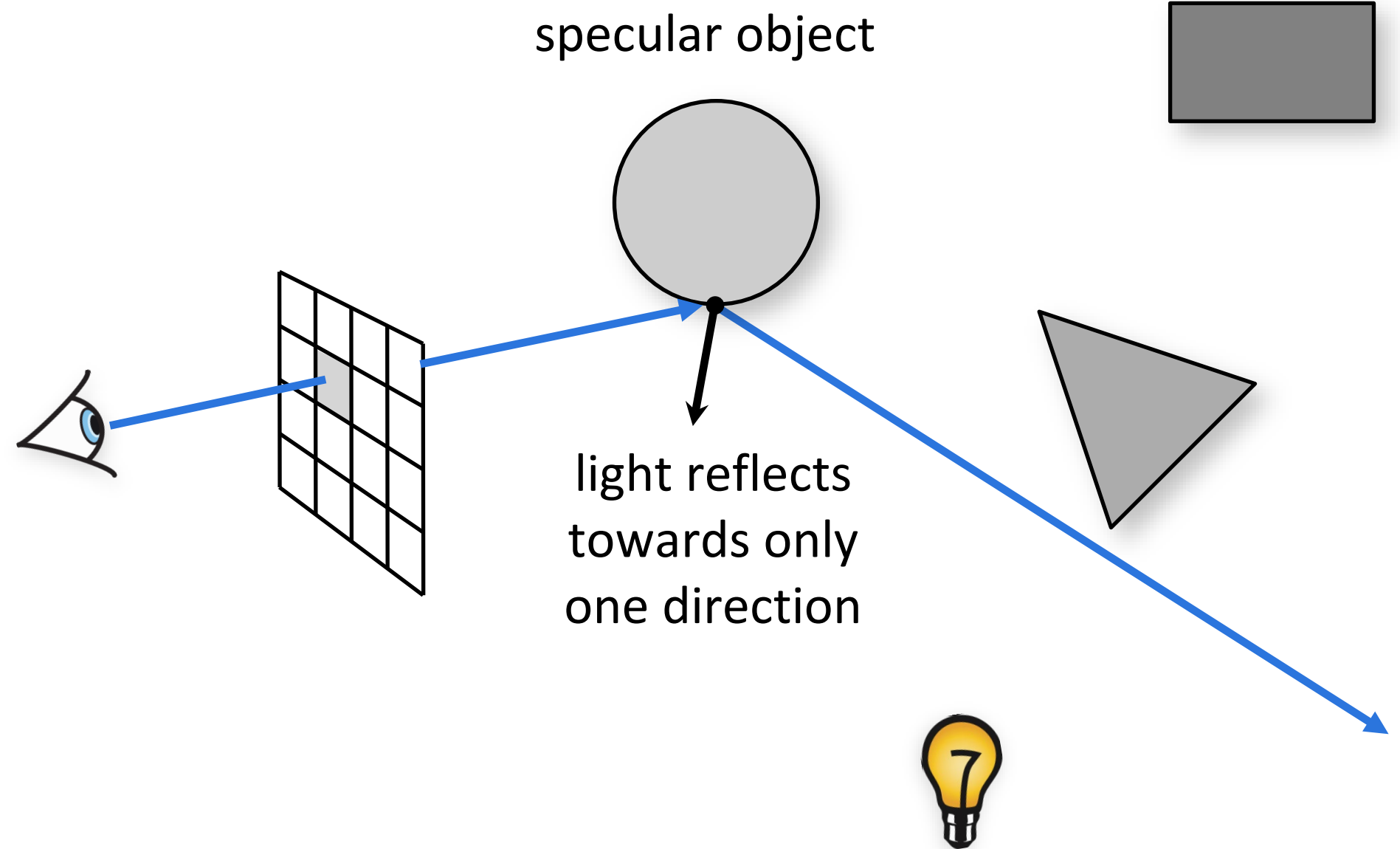
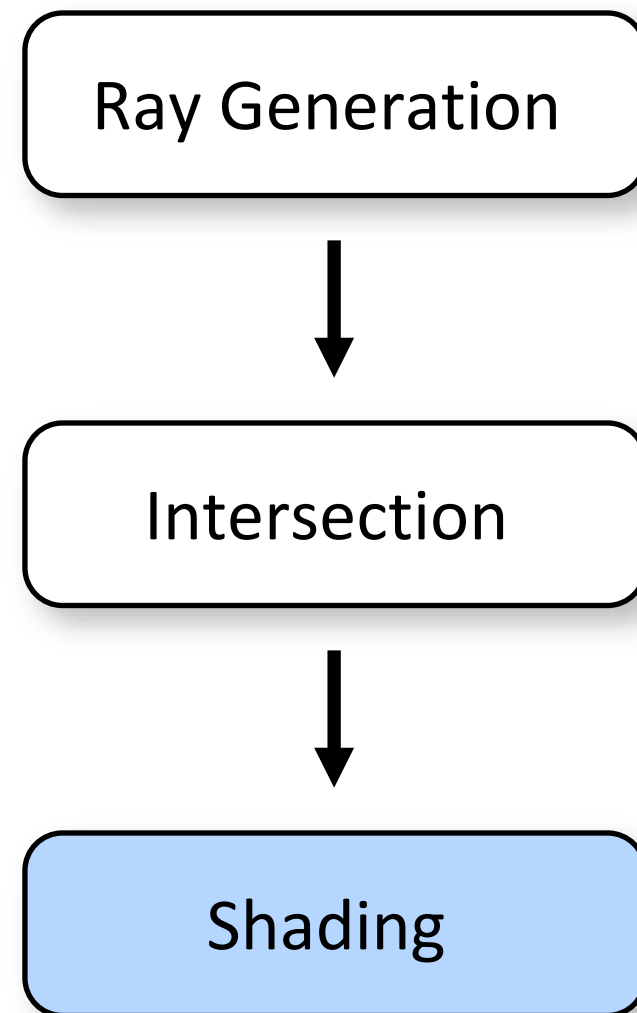


Do this by limiting
the t range

Specular/Mirror reflection

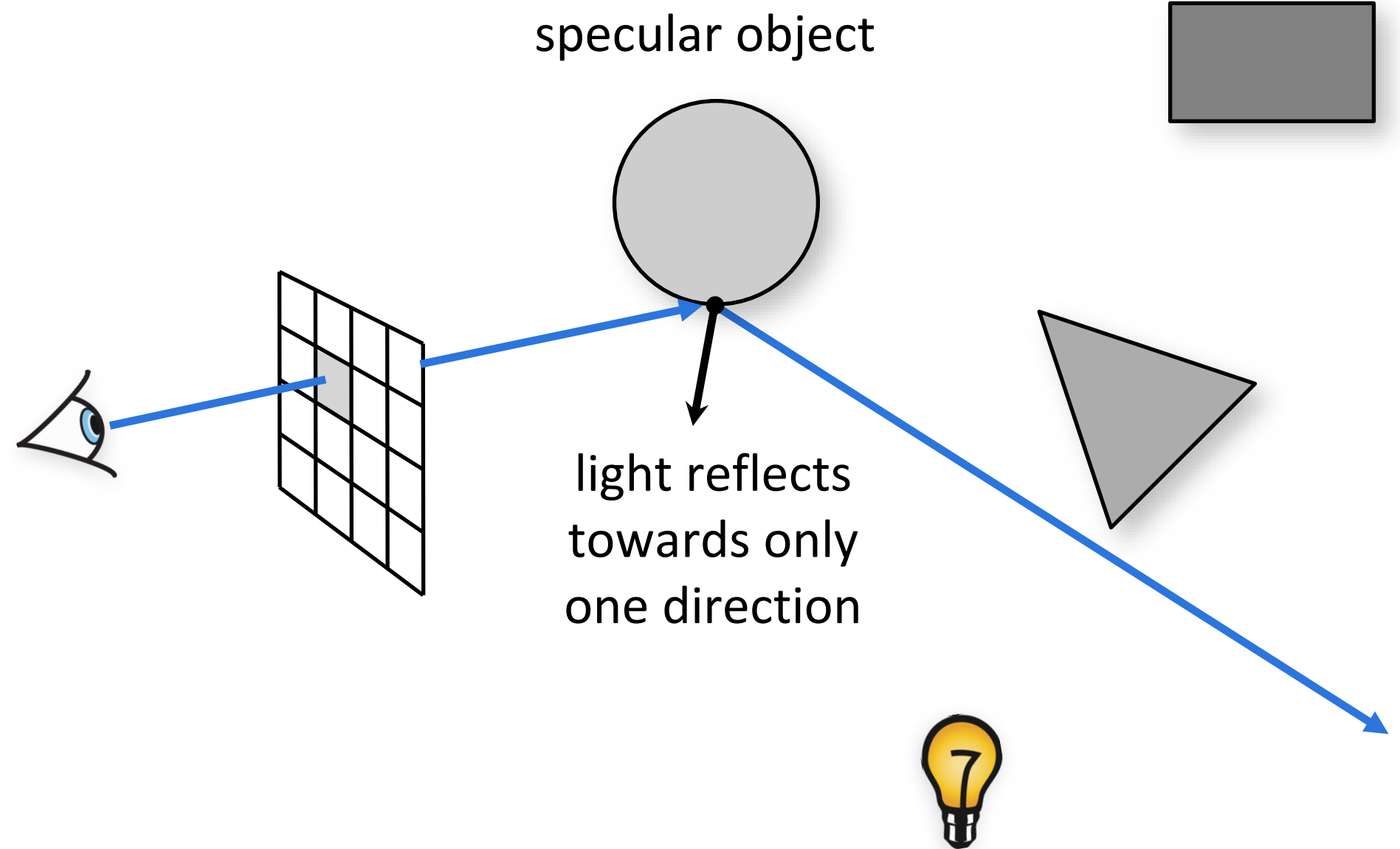
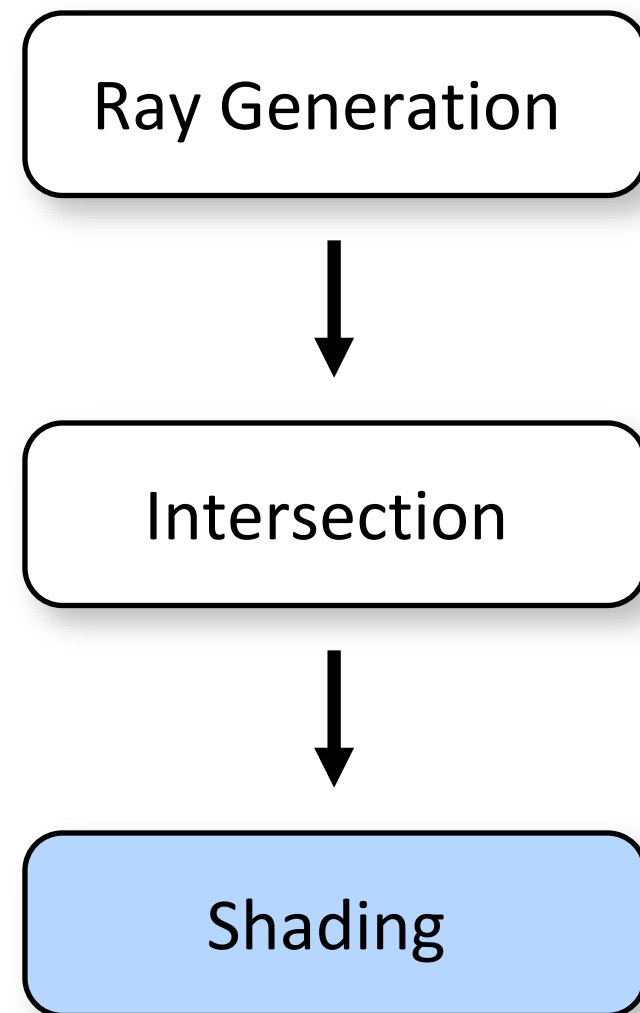


Basic Ray Tracing Pipeline



What direction should we trace a new ray towards?

Basic Ray Tracing Pipeline



What direction should we trace a new ray towards?

- Just use law of mirror reflection, no need for random selection!

Mirror reflection

Consider perfectly shiny surface

- there's a reflection of other objects

Can render this using recursive ray tracing

- to find out mirror reflection color ask:
“what color is seen from surface point in reflection direction?”

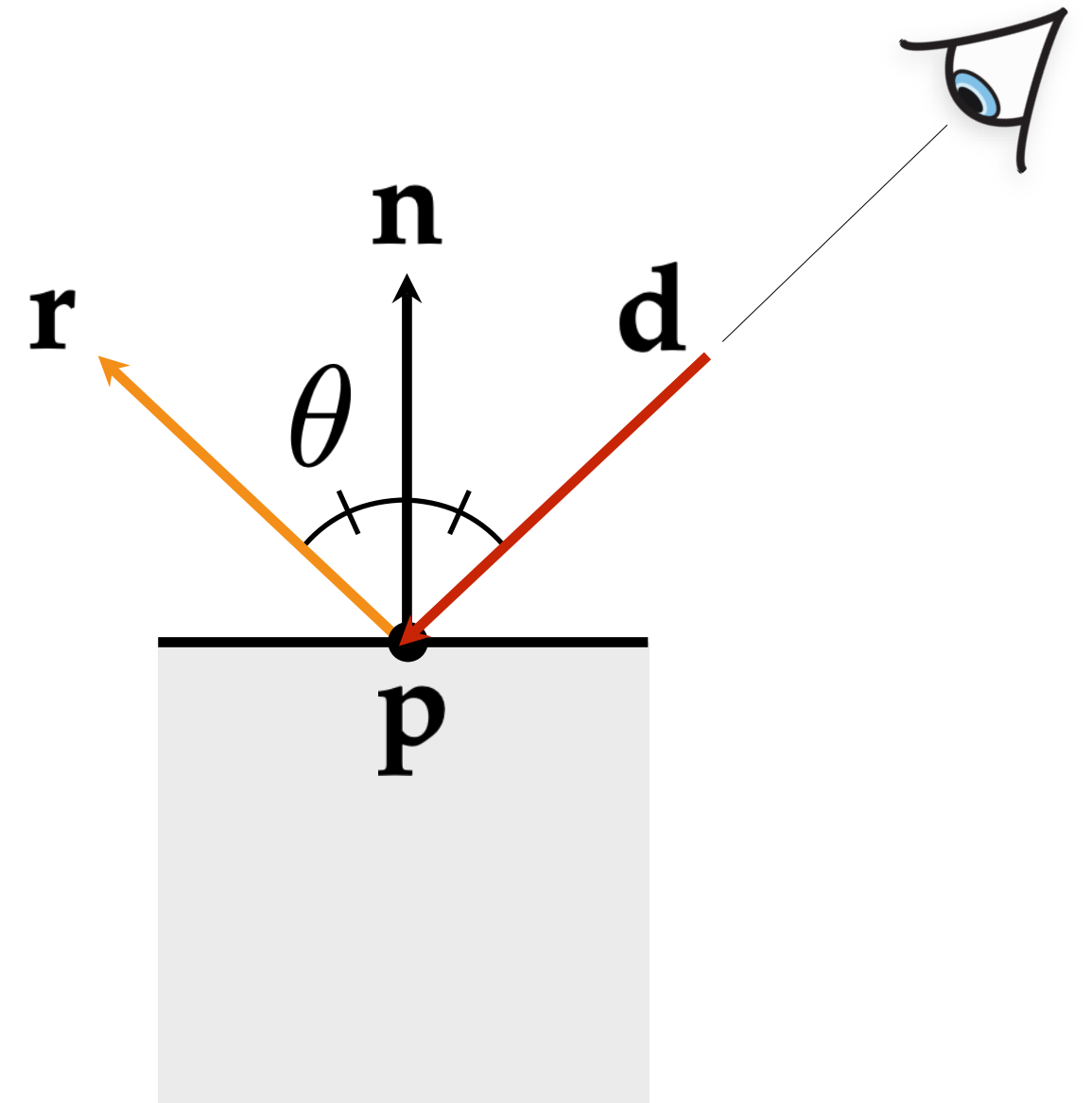
Mirror reflection

Evaluated by tracing a new ray:

$$L_r = k_r \text{trace}(\mathbf{p}, \mathbf{r})$$

Diagram illustrating the components of the mirror reflection equation:

- L_r : reflected light
- k_r : reflection scaled down/tinted
- $\text{trace}(\mathbf{p}, \mathbf{r})$: recursive call



Mirror reflection

Evaluated by tracing a new ray:

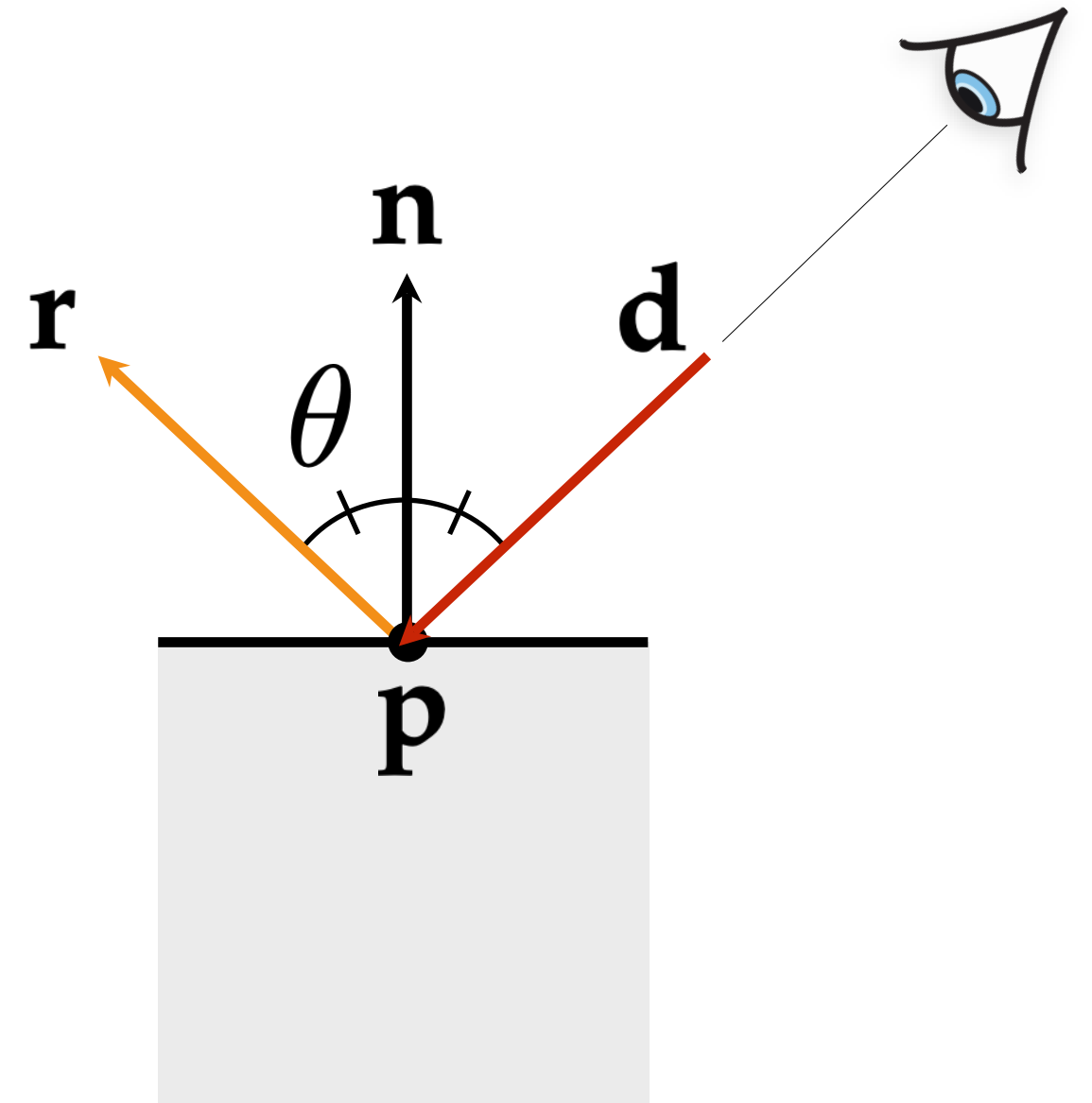
$$L_r = k_r \text{trace}(\mathbf{p}, \mathbf{r})$$

Diagram illustrating the components of the mirror reflection equation:

- L_r : reflected light
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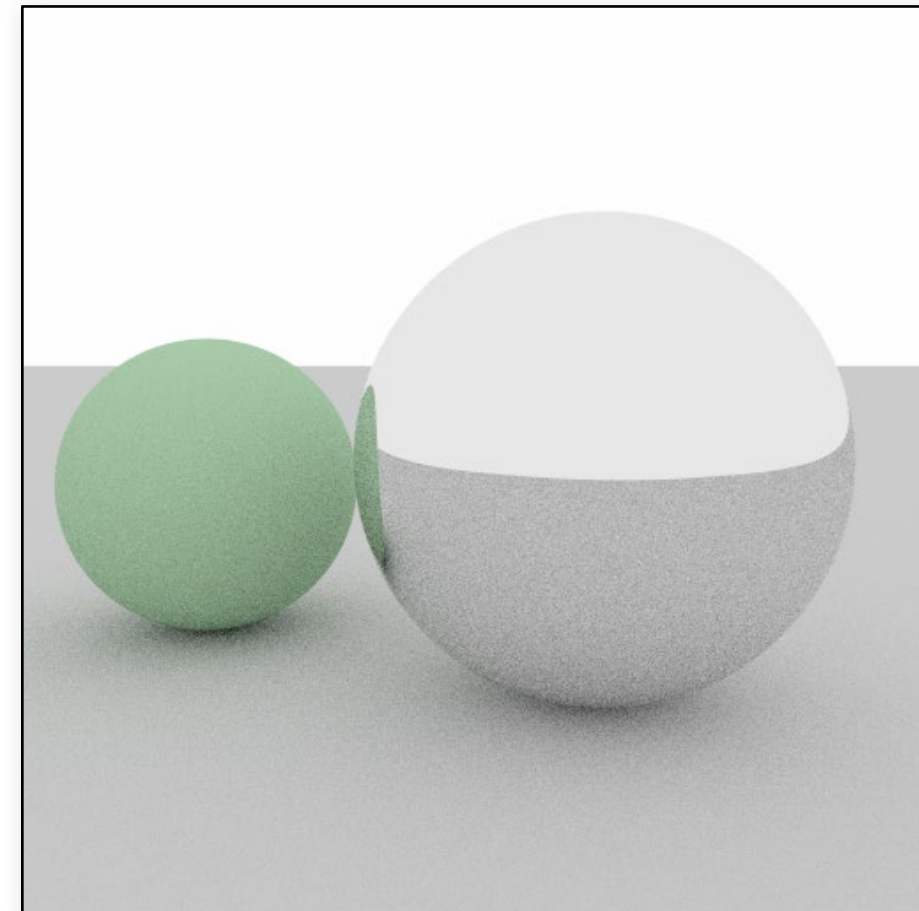
Implementation details:

- don't self-intersect ($t_{\min} > \epsilon$)
- don't recurse indefinitely

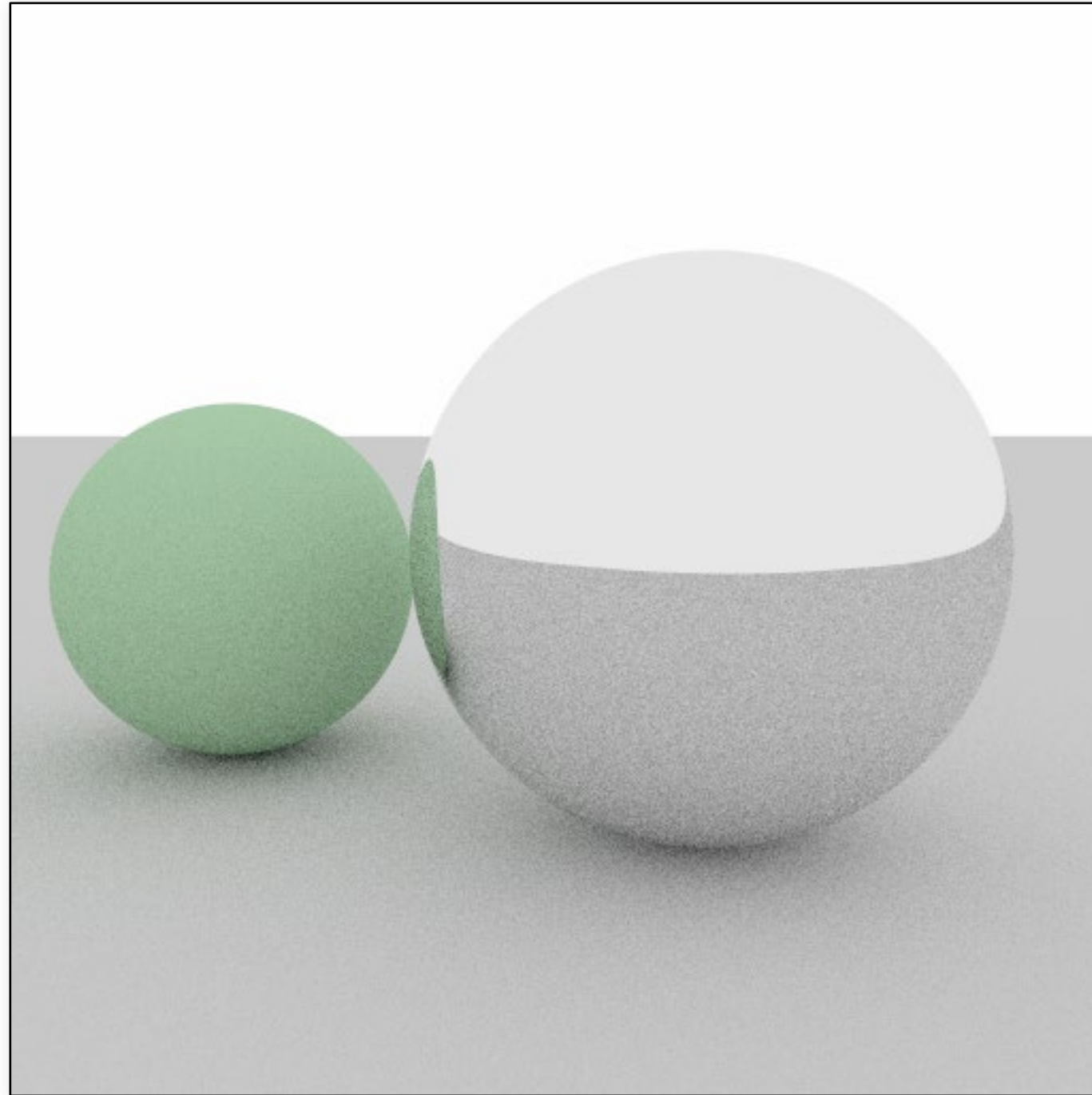


Same pseudo-code

```
Scene::trace(Ray ray)
    hit = surfaces.intersect(ray);
    if hit
        [col, sRay] = hit->mat->scatter(ray)
        return col * trace(sRay);
    else
        return backgroundColor;
```

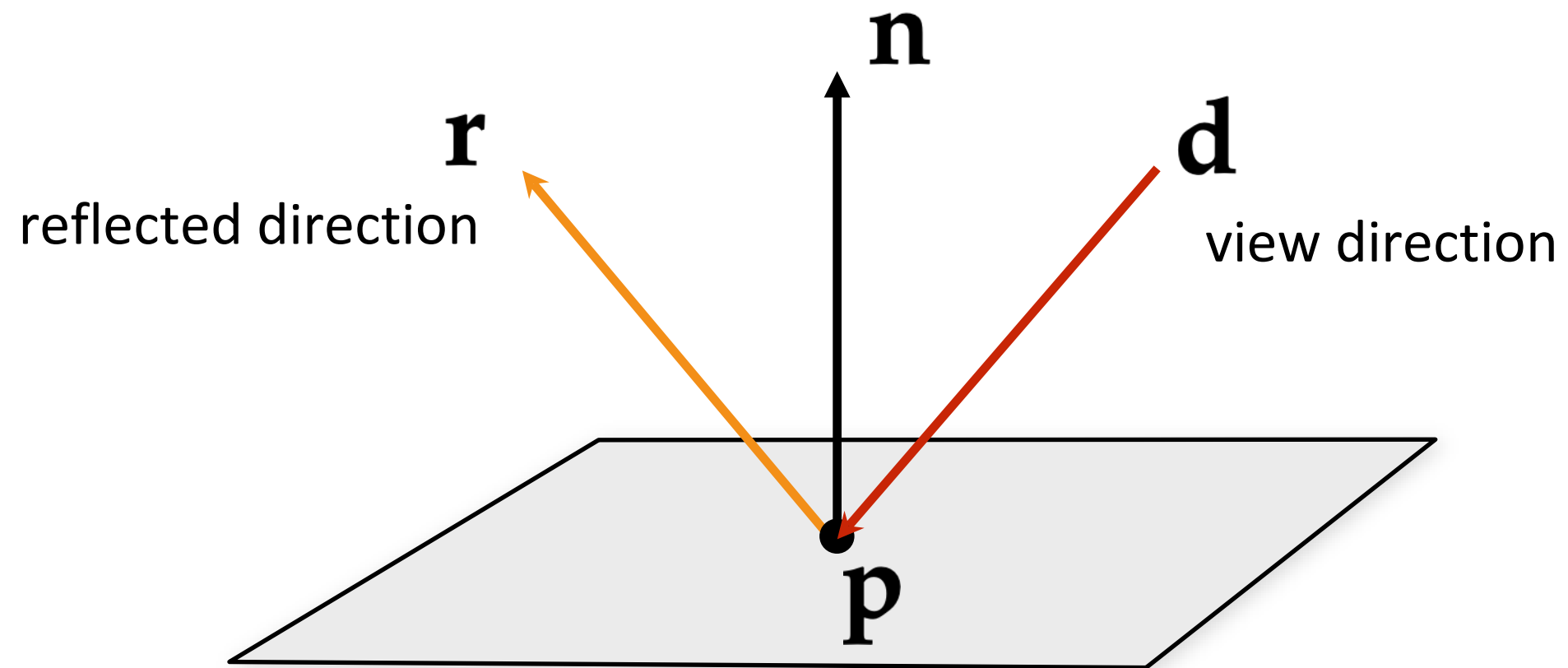


Diffuse & mirror spheres



Mirror reflection

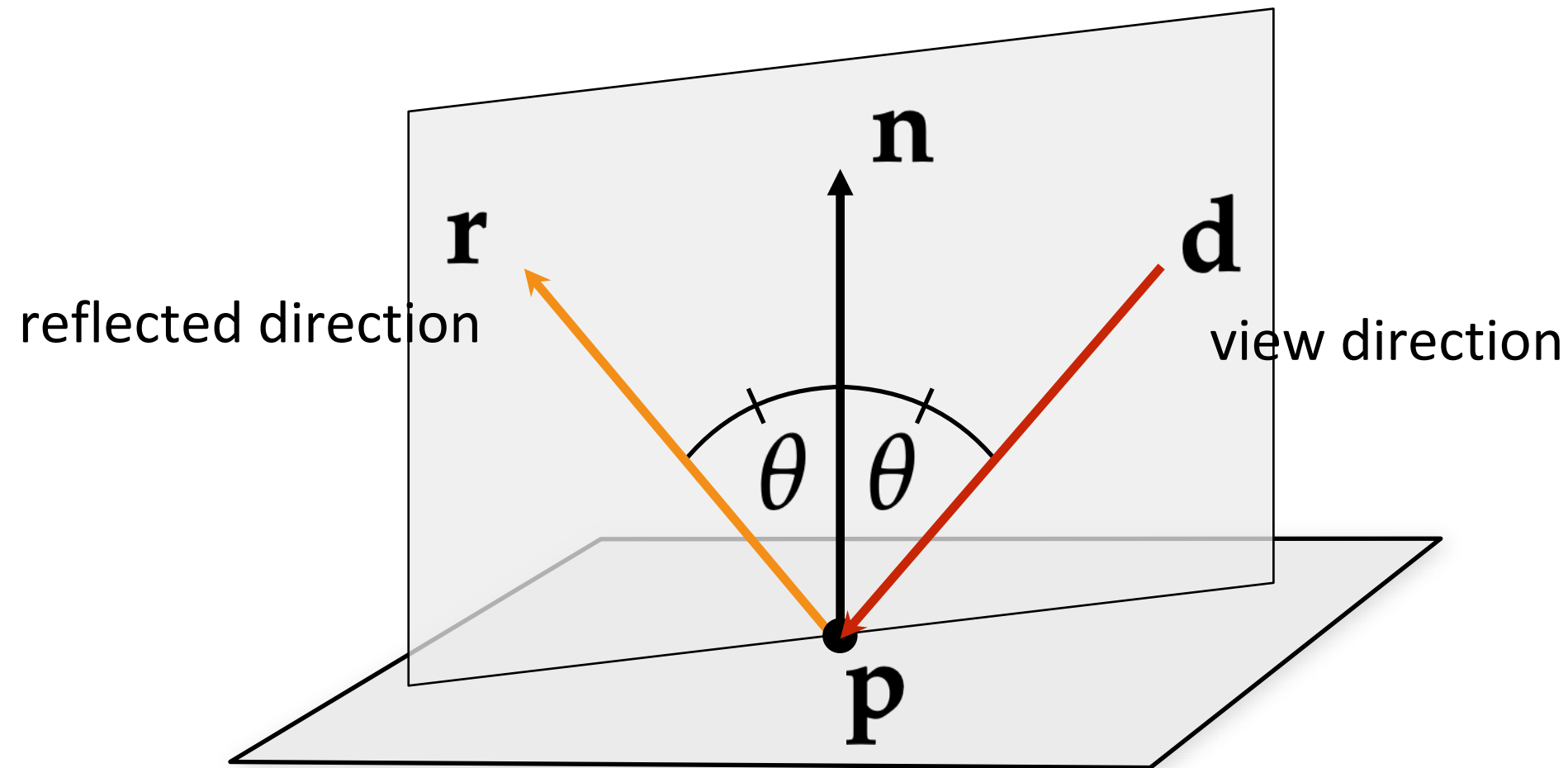
Assume \mathbf{n} is unit length



What two properties defined reflection direction?

Mirror reflection

Assume \mathbf{n} is unit length

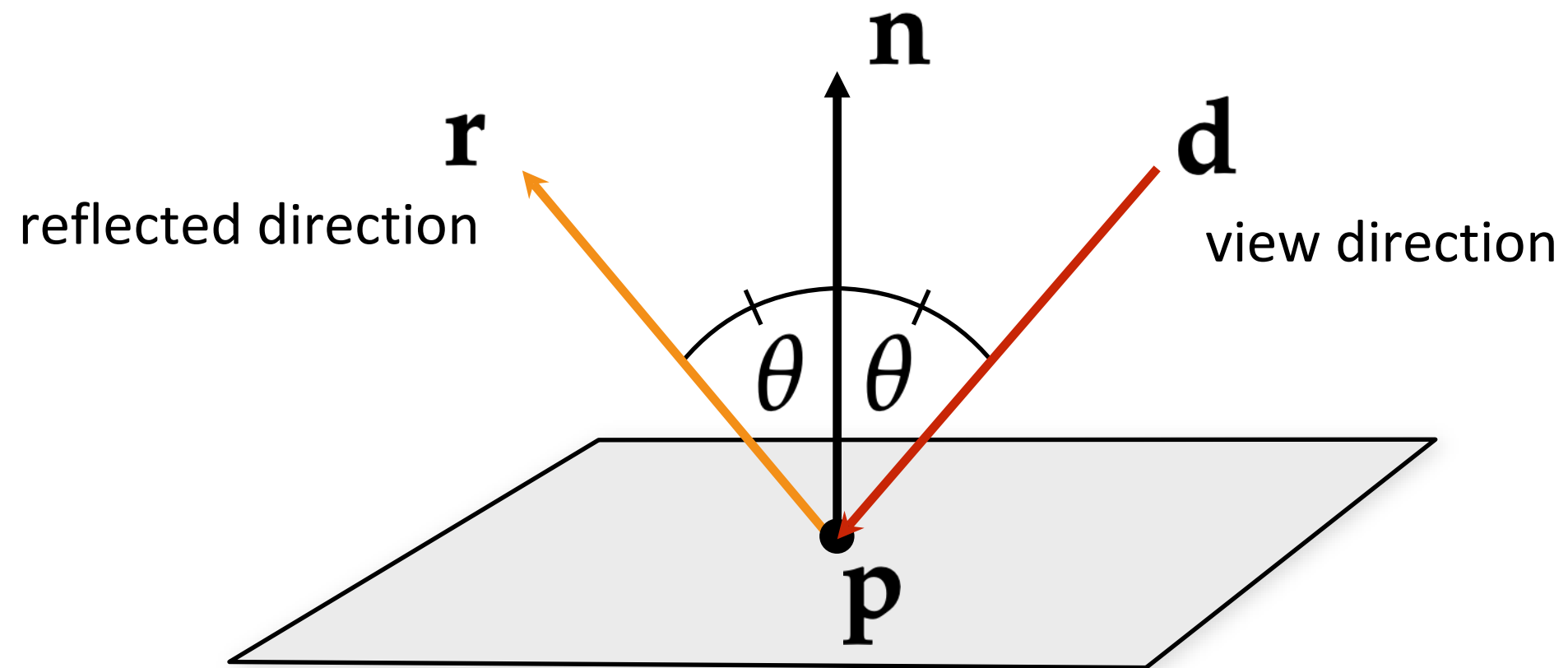


What two properties defined reflection direction?

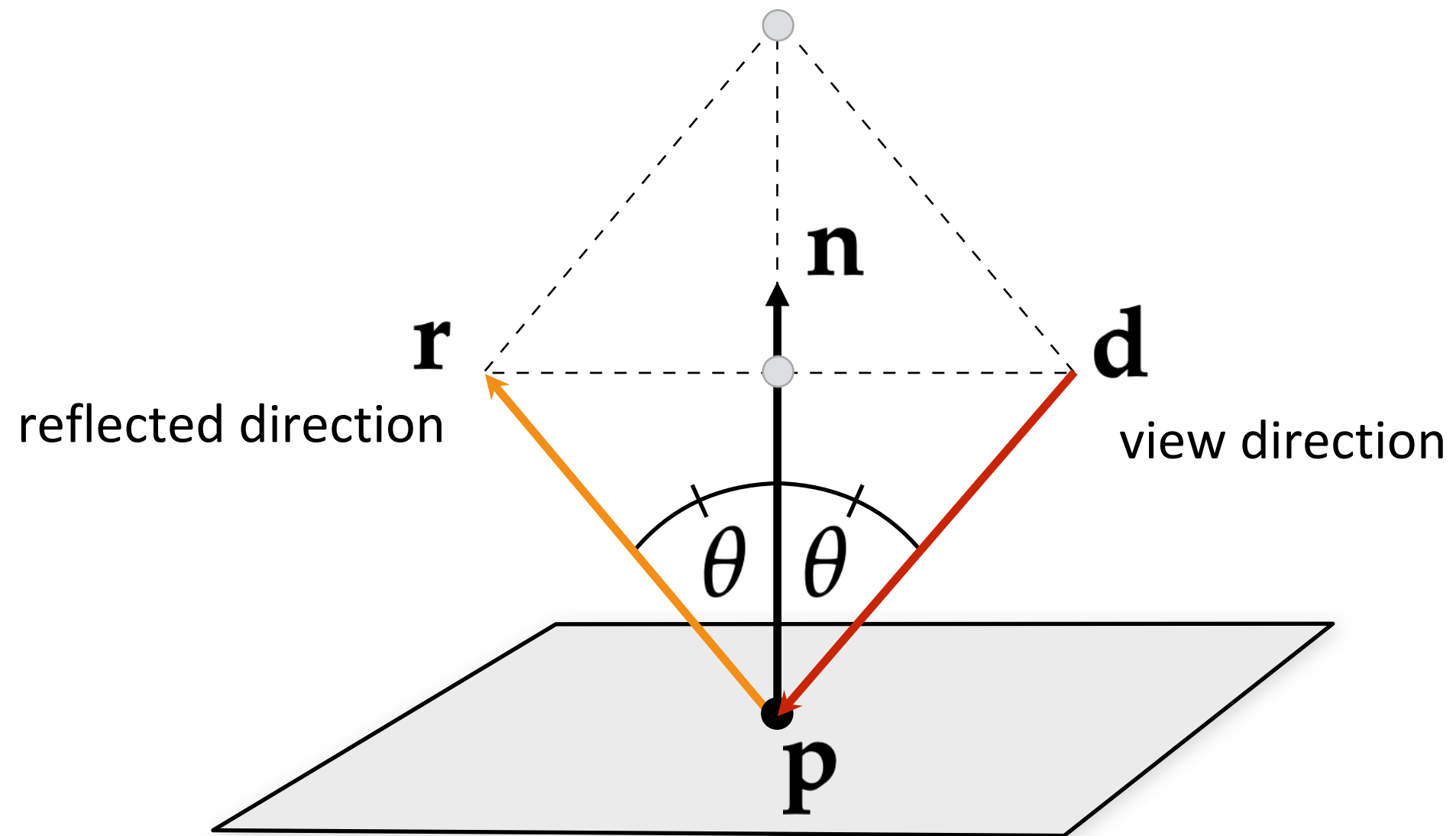
- co-planar view direction, reflected direction, and normal direction
- equal angles between normal-view directions, and normal-reflected directions

Mirror reflection

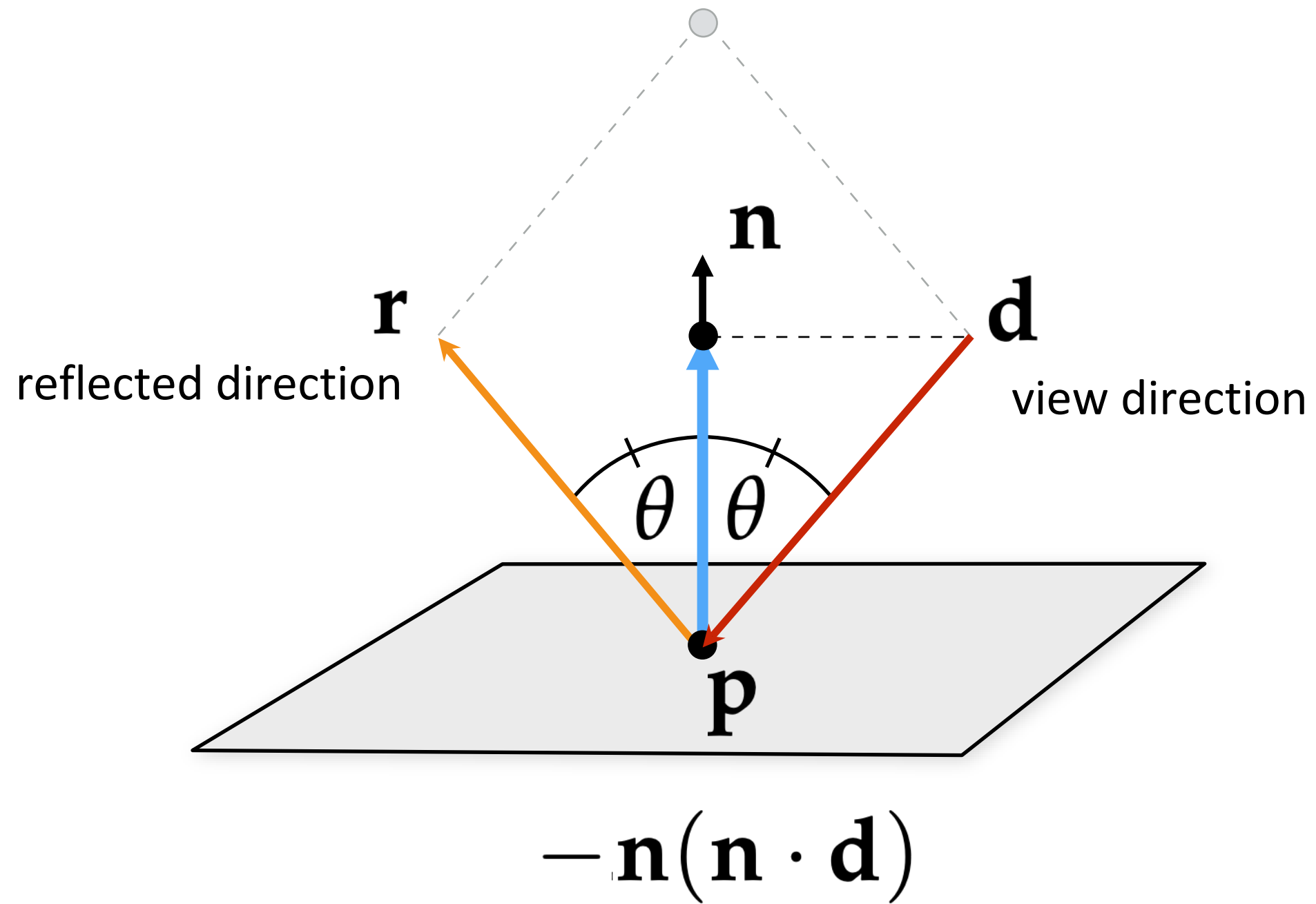
Assume \mathbf{n} is unit length



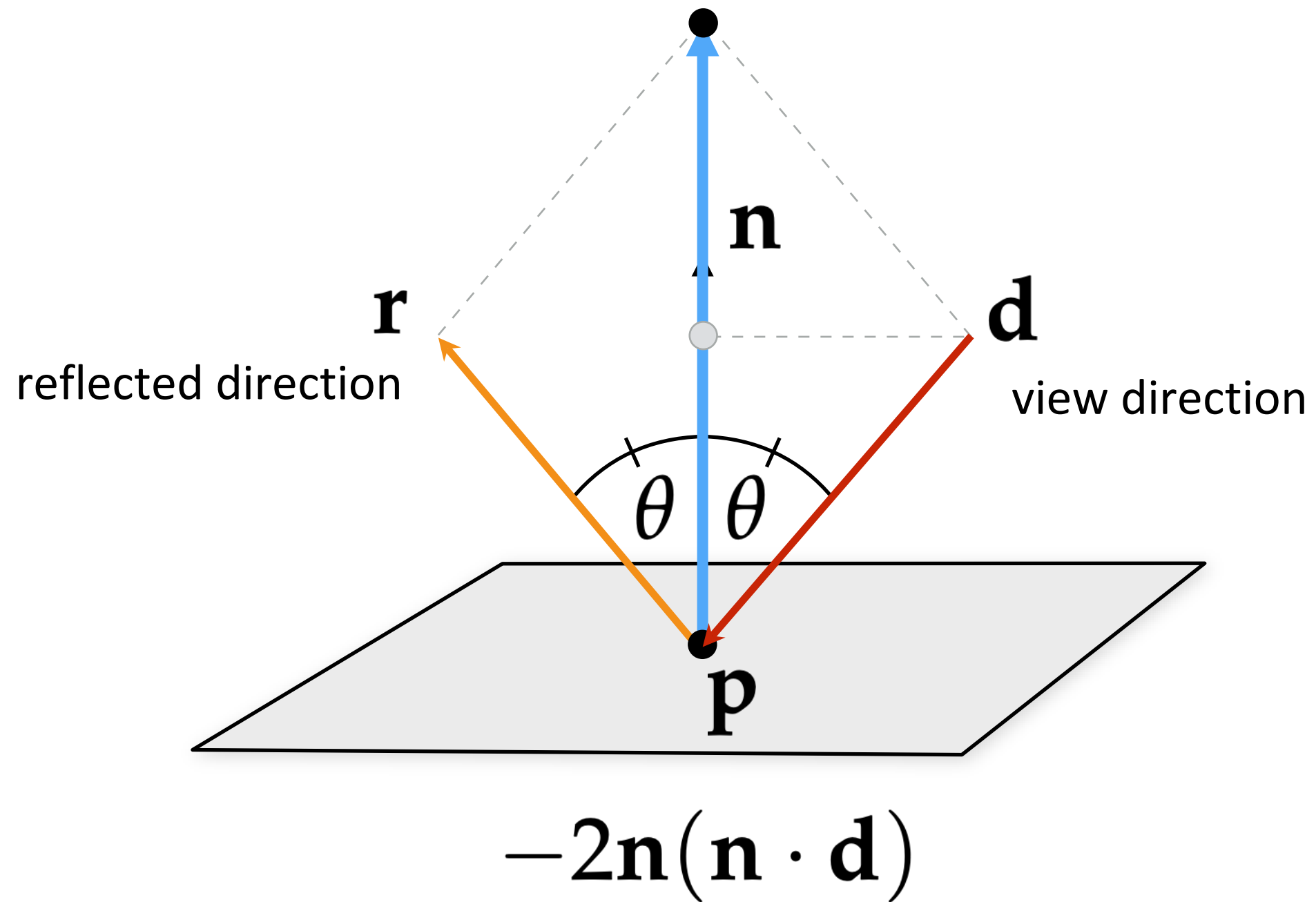
Mirror reflection



Mirror reflection

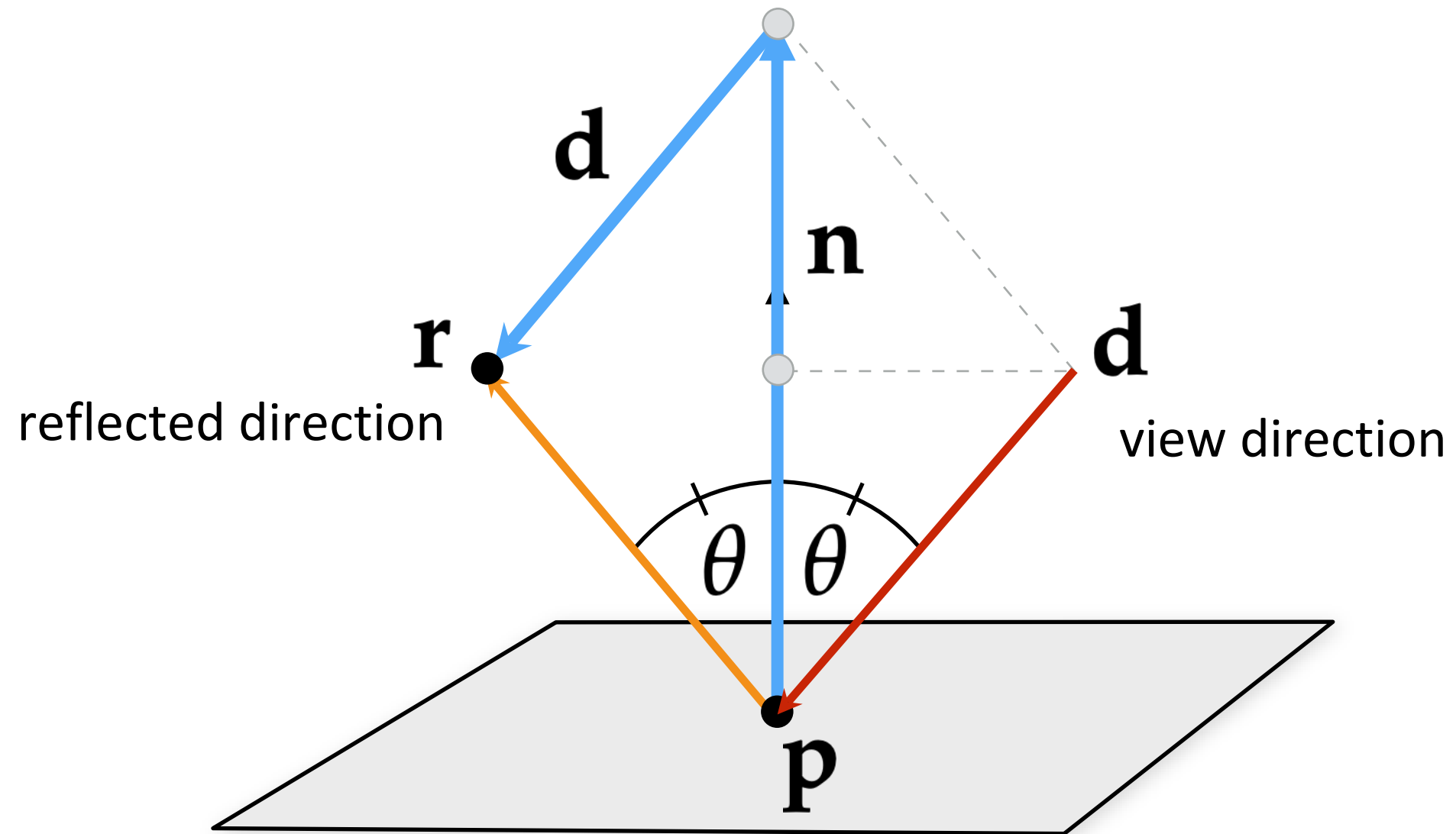


Mirror reflection



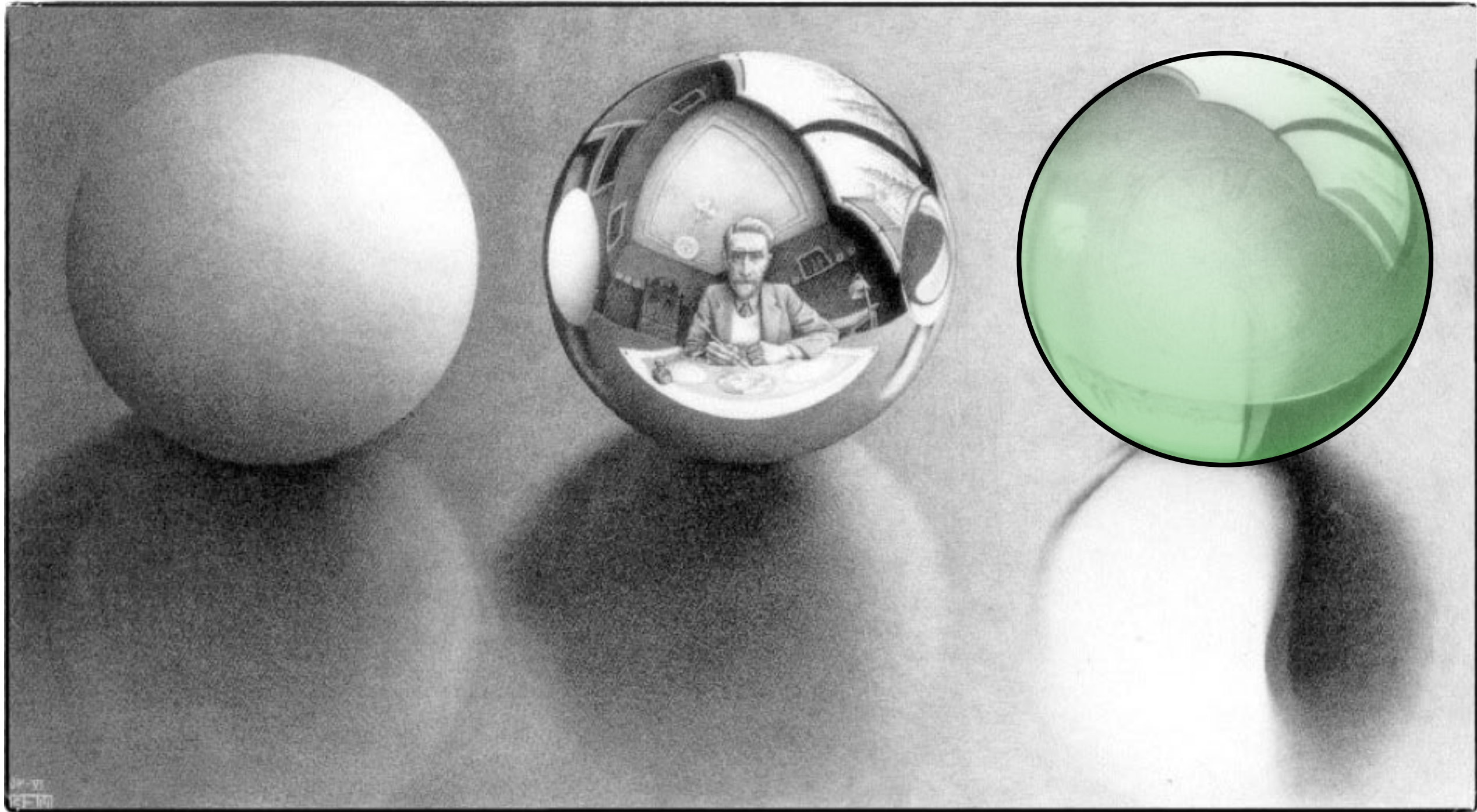
Mirror reflection

Assumes \mathbf{n} is unit length



$$\mathbf{r} = -2\mathbf{n}(\mathbf{n} \cdot \mathbf{d}) + \mathbf{d}$$

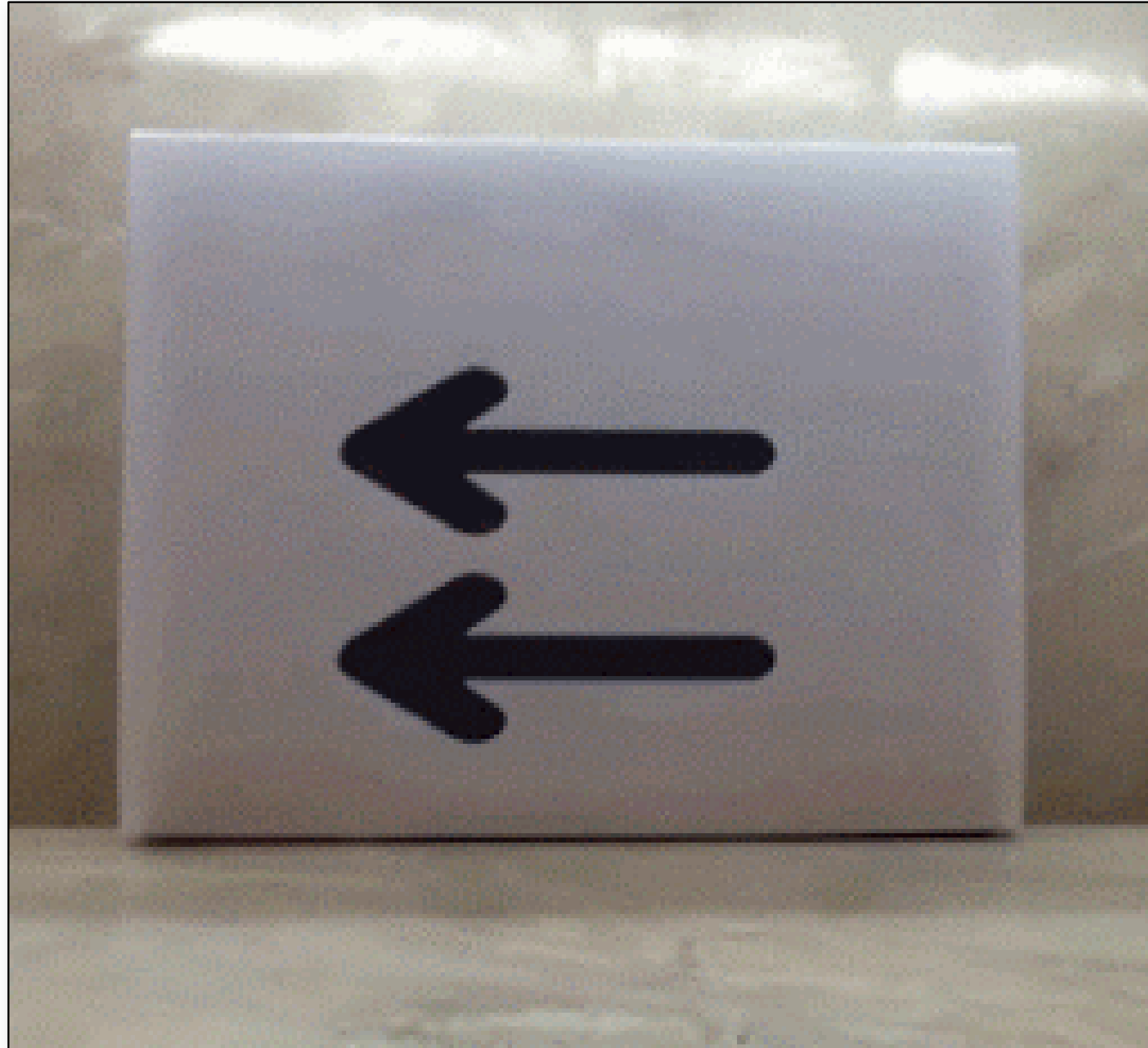
Specular refraction



Refraction



Refraction



Index of Refraction

Speed of light in vacuum / speed of light in medium

Some values of η	
Vacuum	1
Air at STP	1.00029
Ice	1.31
Water	1.33
Crown glass	1.52 - 1.65
Diamond	2.417

Specular transmission/refraction

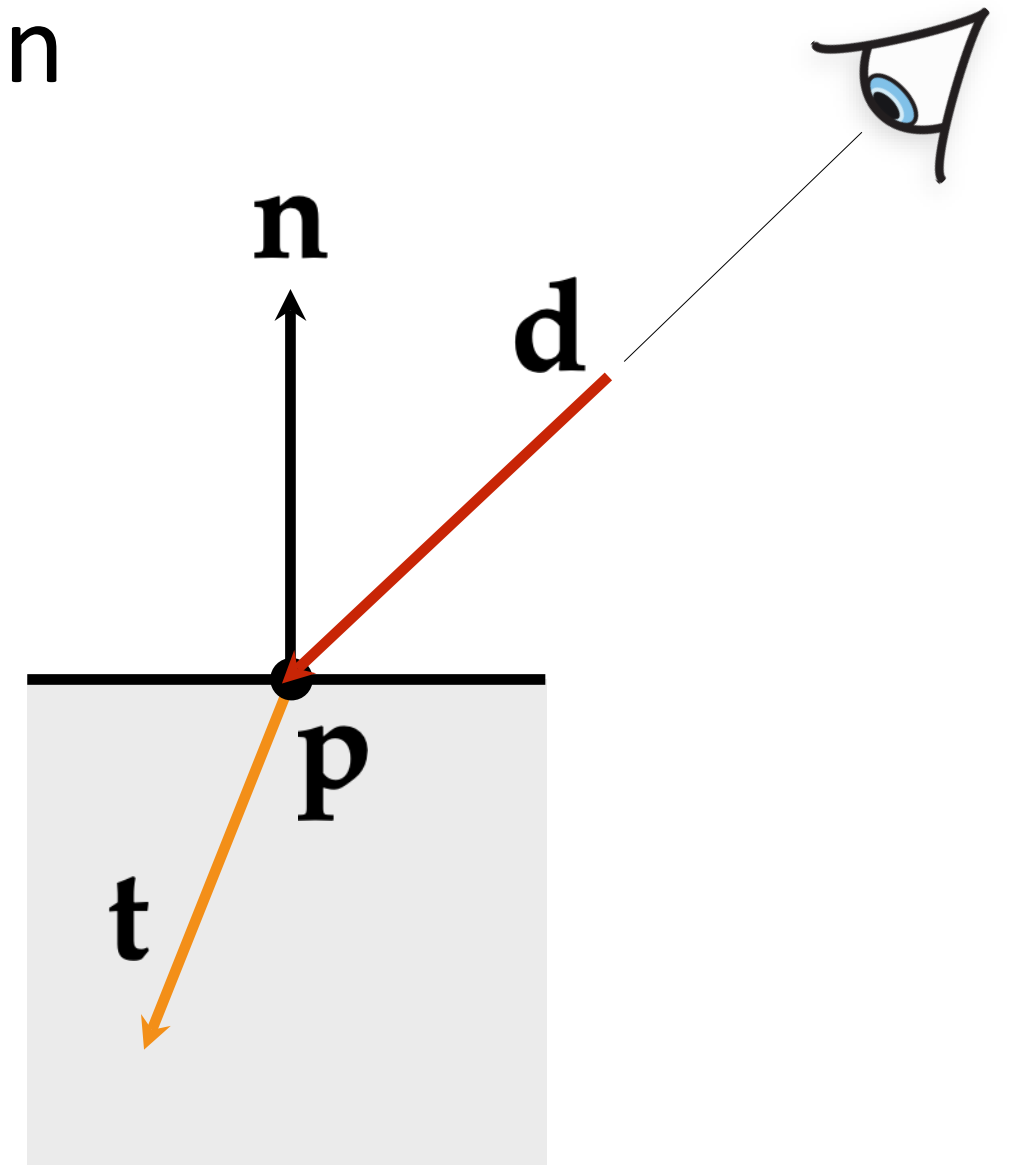
Materials like water, glass, etc., also refract/bend light

Trace a recursive ray in the refraction direction

$$L_t = k_t \text{trace}(\mathbf{p}, \mathbf{t})$$

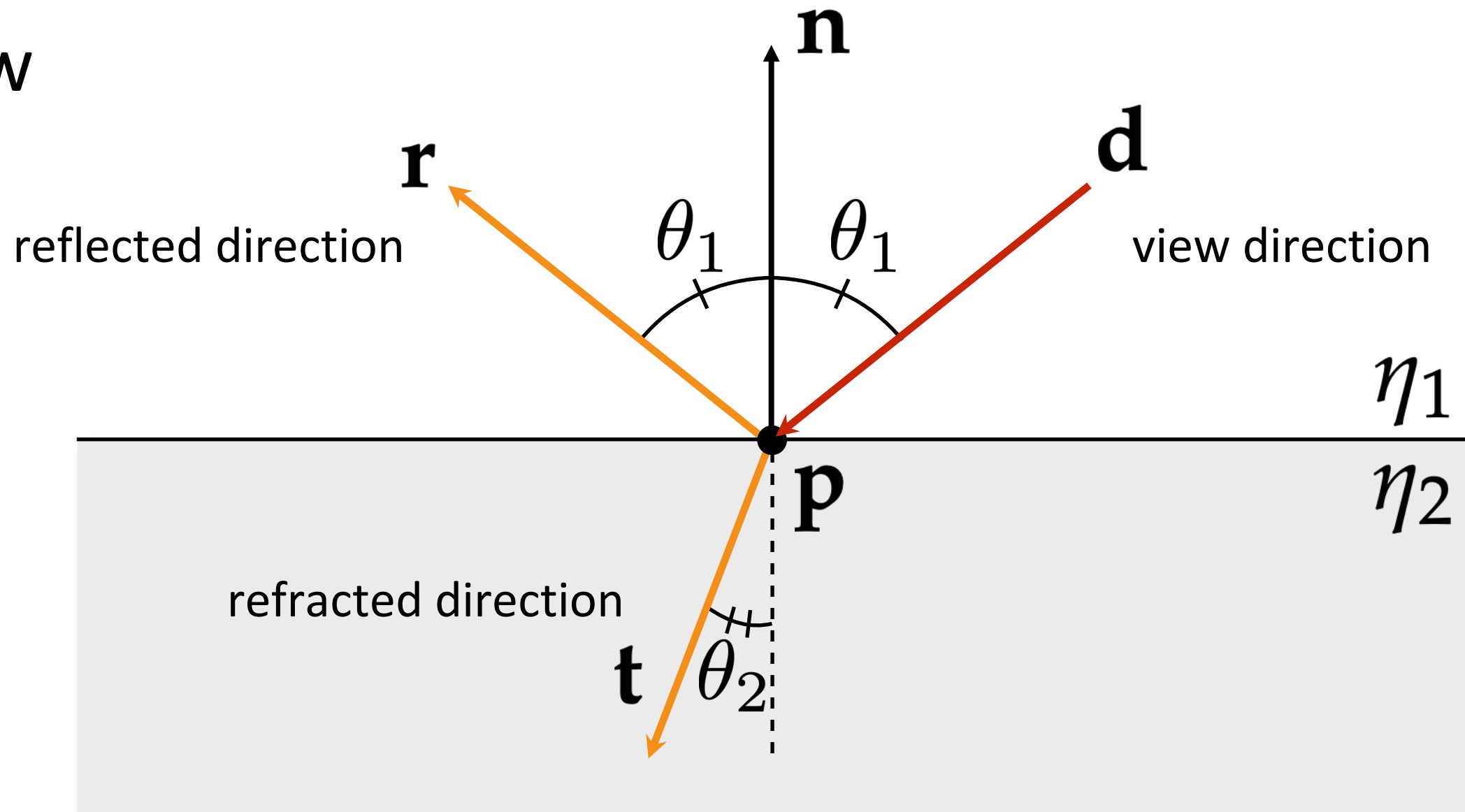
Diagram illustrating the recursive ray tracing equation for refraction:

- L_t : refracted light
- k_t : refraction scaled down
- $\text{trace}(\mathbf{p}, \mathbf{t})$: recursive call



Specular transmission/refraction

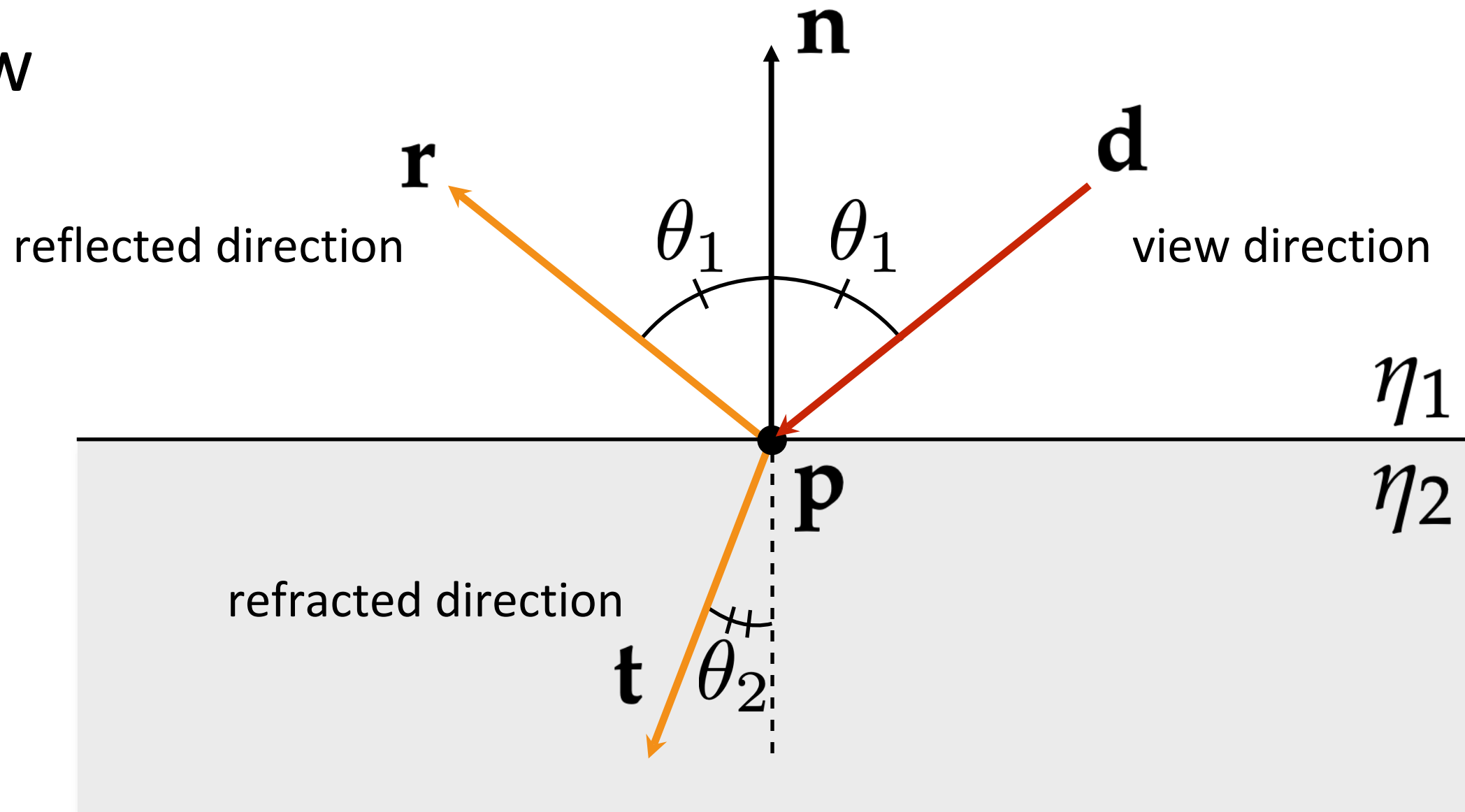
Snell's law



$$\eta_1 \sin \theta_1 = \eta_2 \sin \theta_2$$

Specular transmission/refraction

Snell's law



$$\mathbf{t} = \eta_1/\eta_2 (\mathbf{d} - (\mathbf{d} \cdot \mathbf{n}) \mathbf{n}) - \mathbf{n} \sqrt{1 - \eta_1^2/\eta_2^2 (1 - (\mathbf{d} \cdot \mathbf{n})^2)}$$

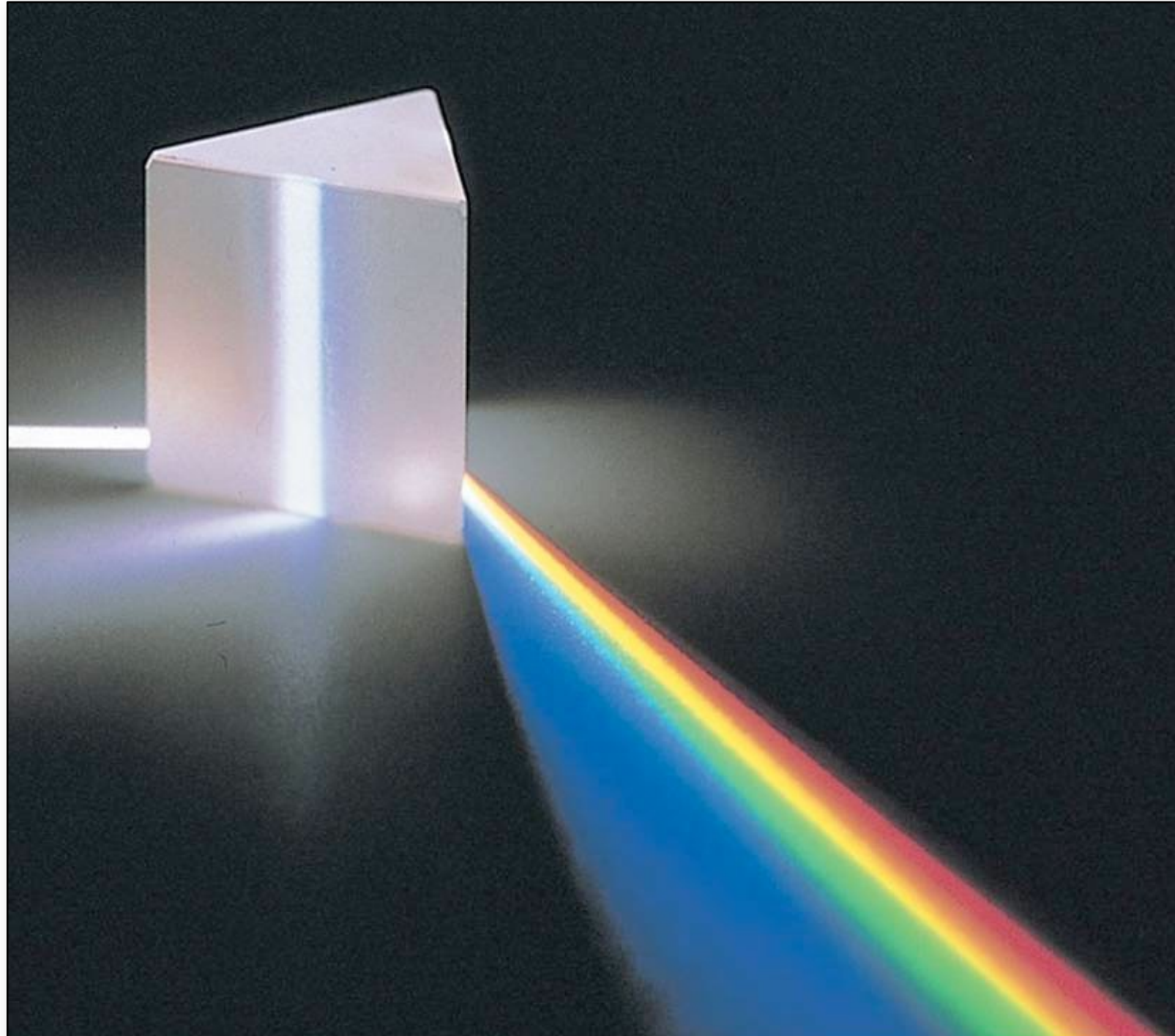
Index of Refraction

Speed of light in vacuum / speed of light in medium

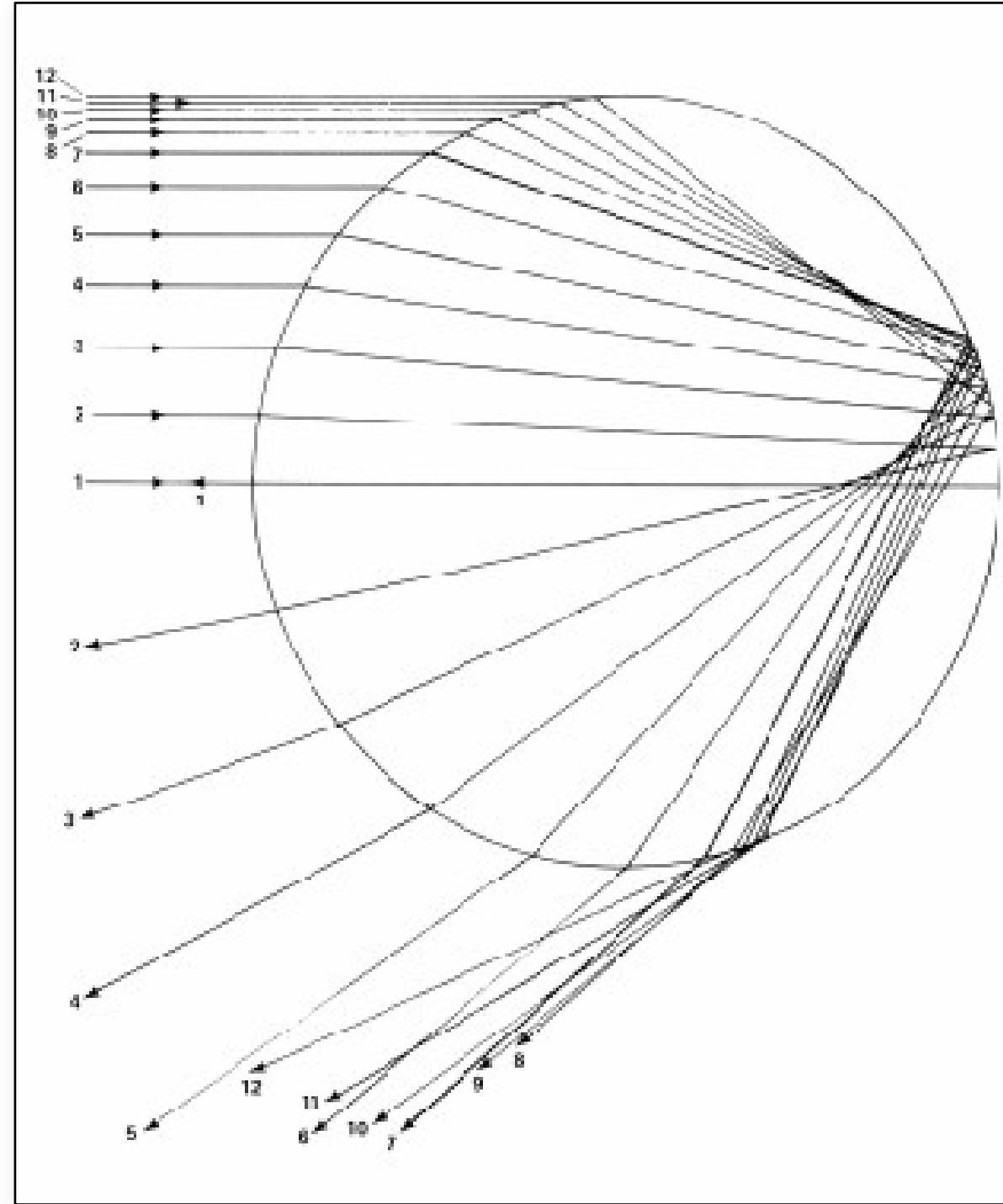
Some values of η	
Vacuum	1
Air at STP	1.00029
Ice	1.31
Water	1.33
Crown glass	1.52 - 1.65
Diamond	2.417

These are actually wavelength dependent!

Dispersion



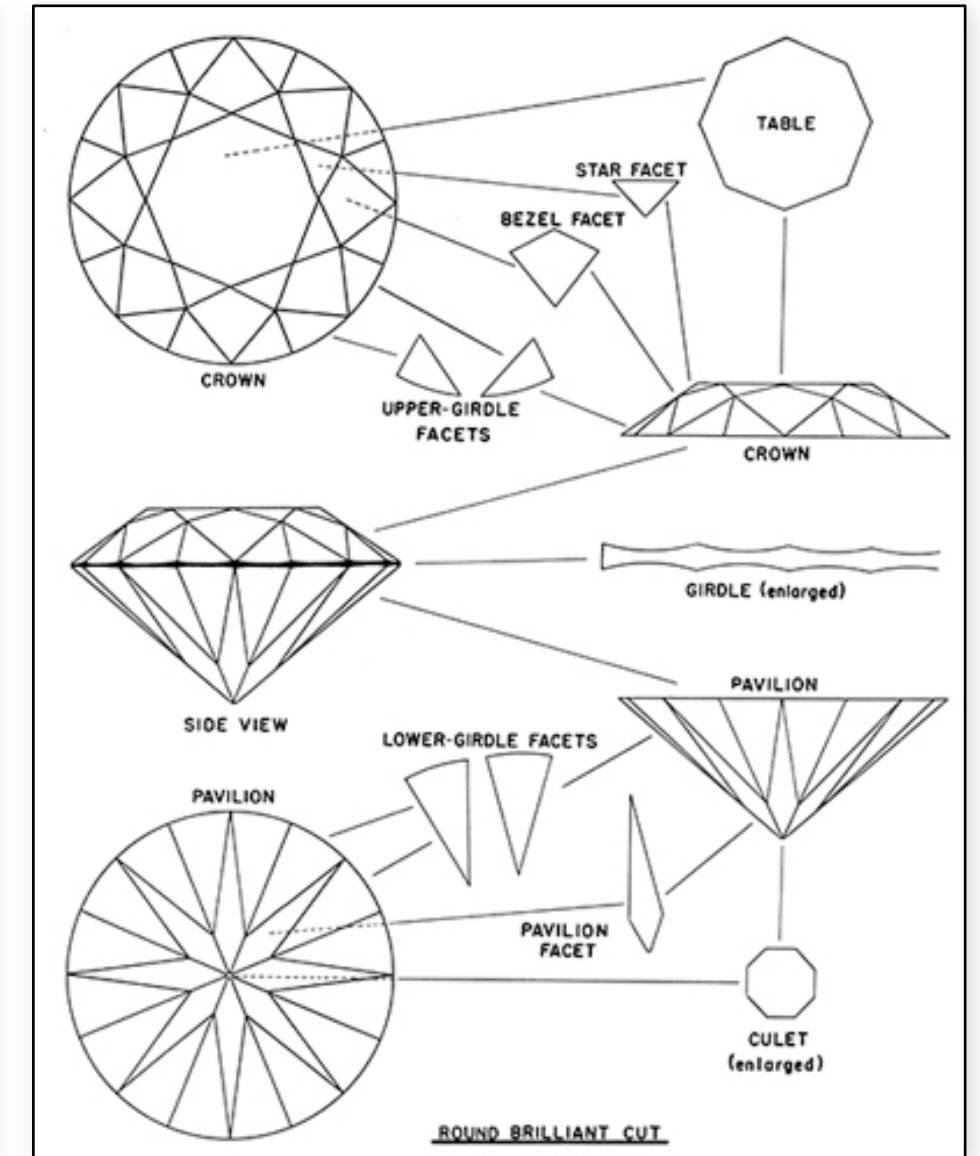
Refraction in a Waterdrop



Double rainbow all the way across the sky!



Dispersion



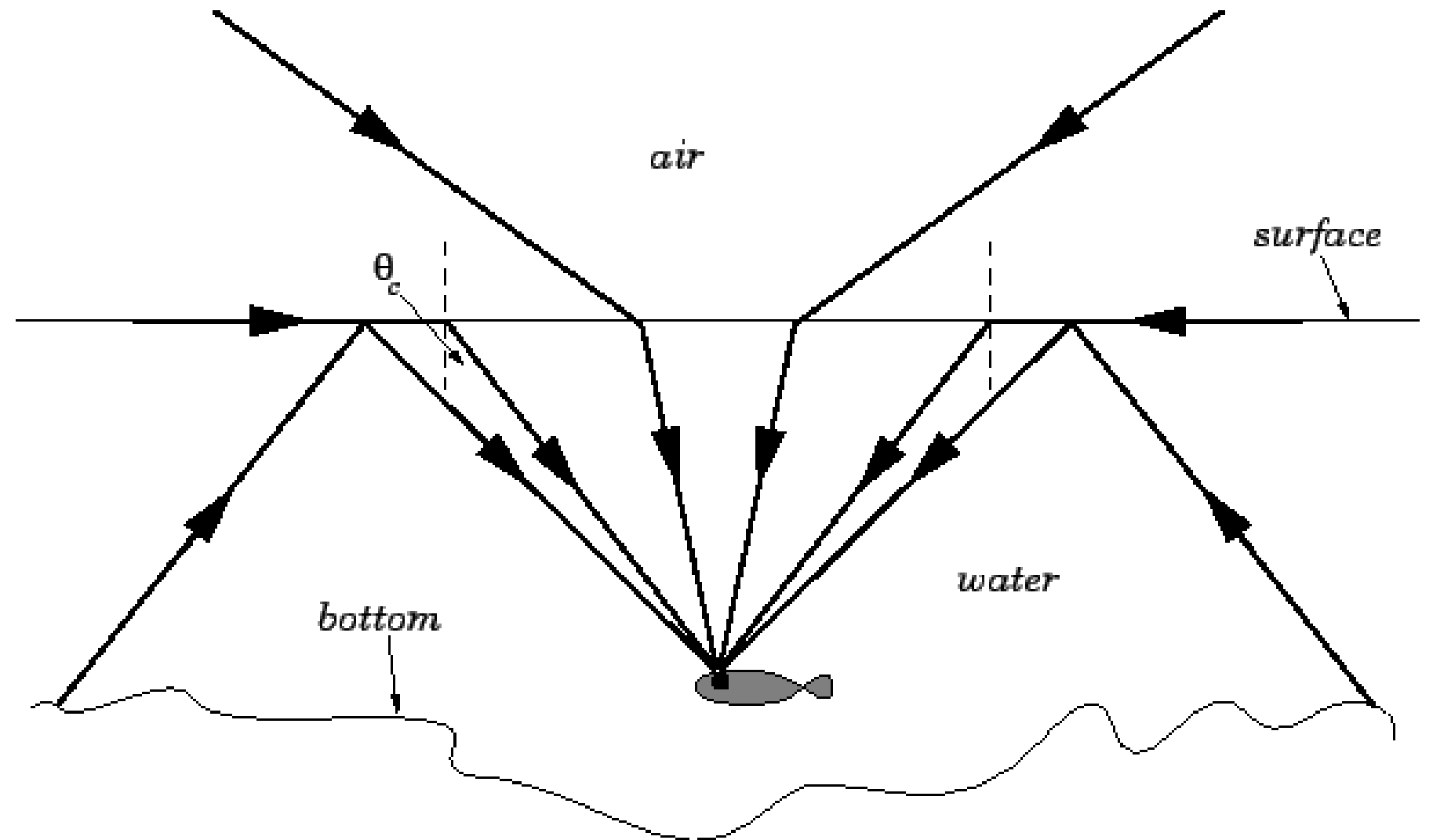
What is this dark circle?



What is this dark circle?

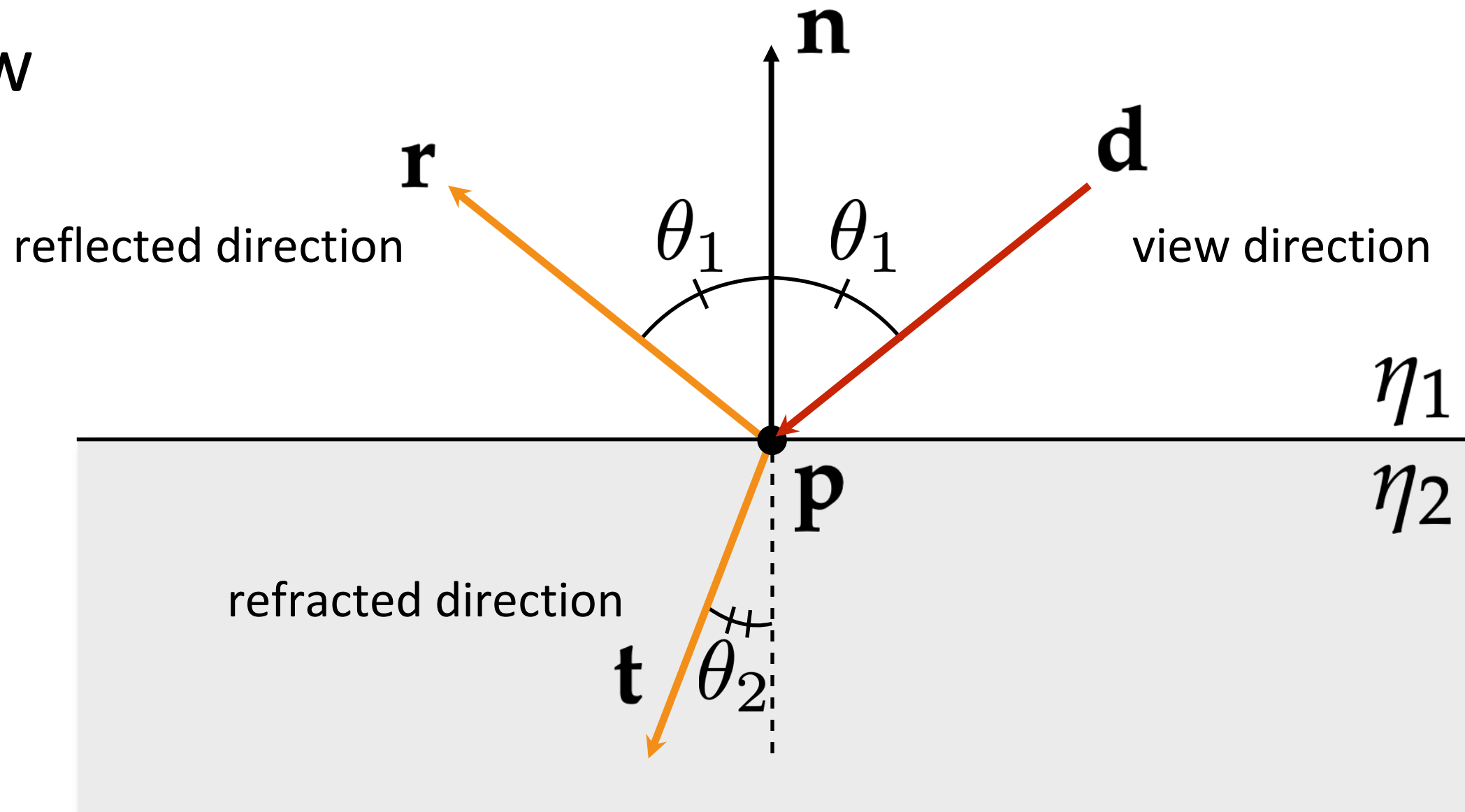


Called
“Snell's window”
Caused by total
internal reflection



Recall...

Snell's law

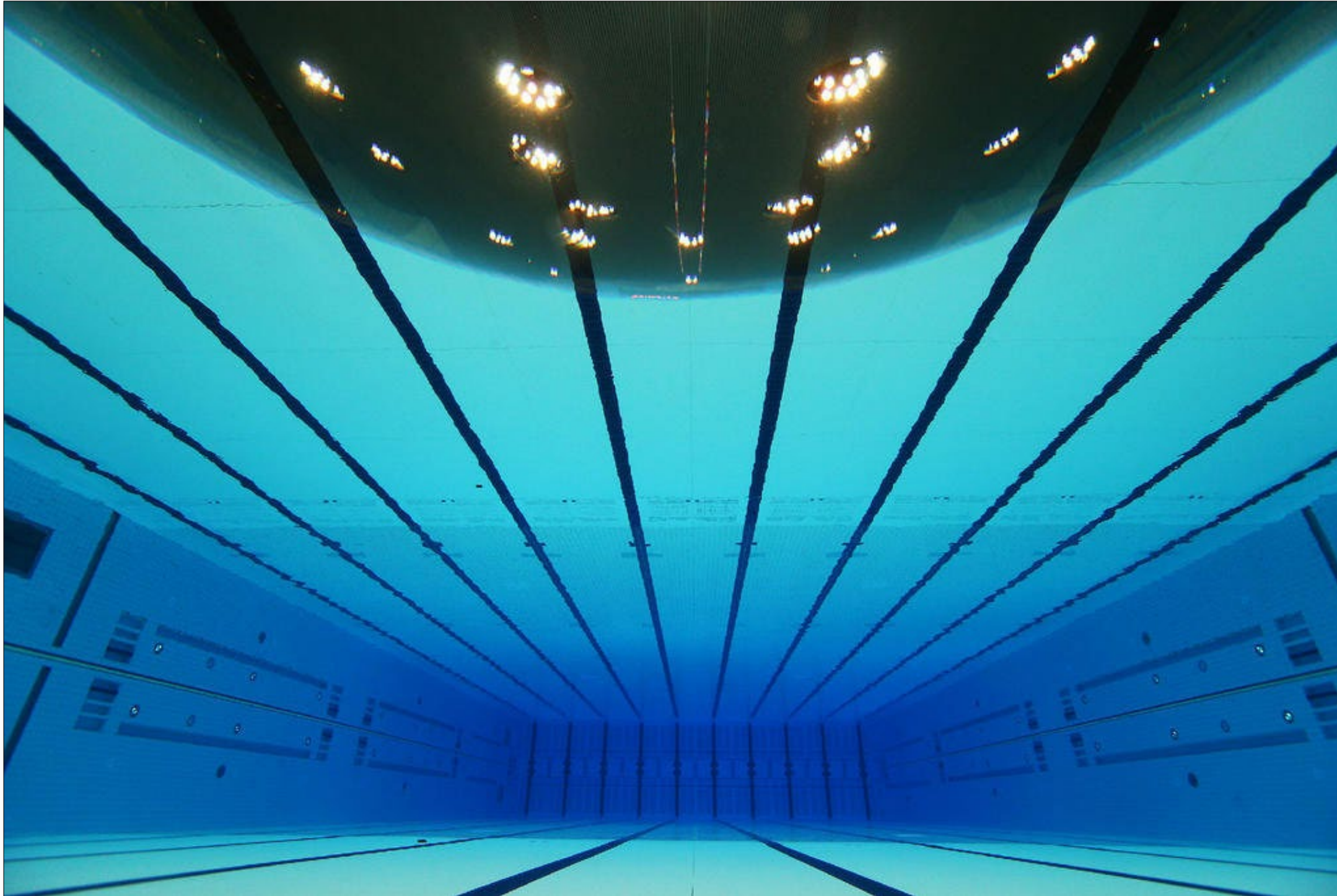


$$\mathbf{t} = \eta_1/\eta_2 (\mathbf{d} - (\mathbf{d} \cdot \mathbf{n}) \mathbf{n}) - \mathbf{n} \sqrt{1 - \eta_1^2/\eta_2^2 (1 - (\mathbf{d} \cdot \mathbf{n})^2)}$$

When can total internal reflection happen?

Can only happen when the ray starts in the higher index medium

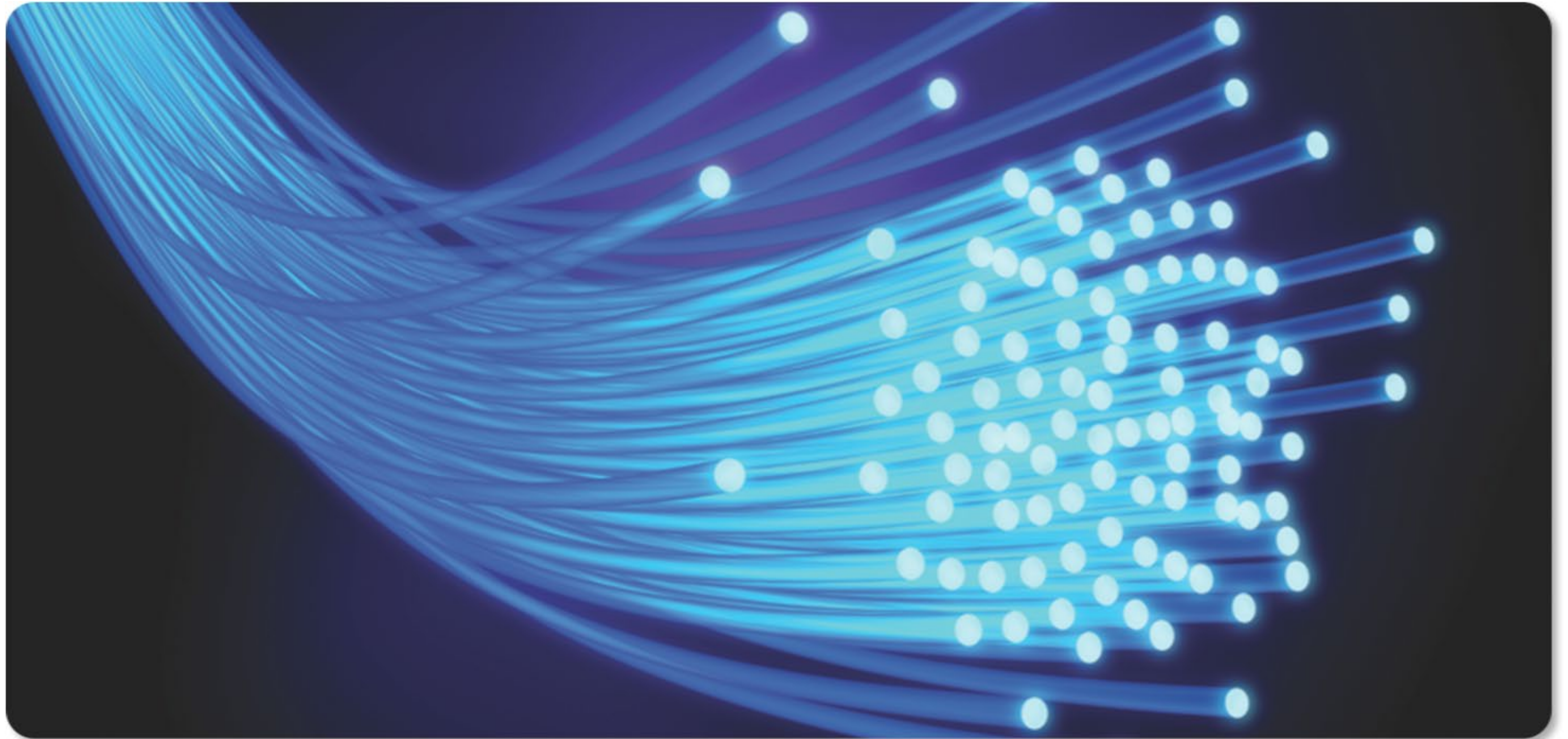
Total Internal Reflection



Total Internal Reflection



Total Internal Reflection



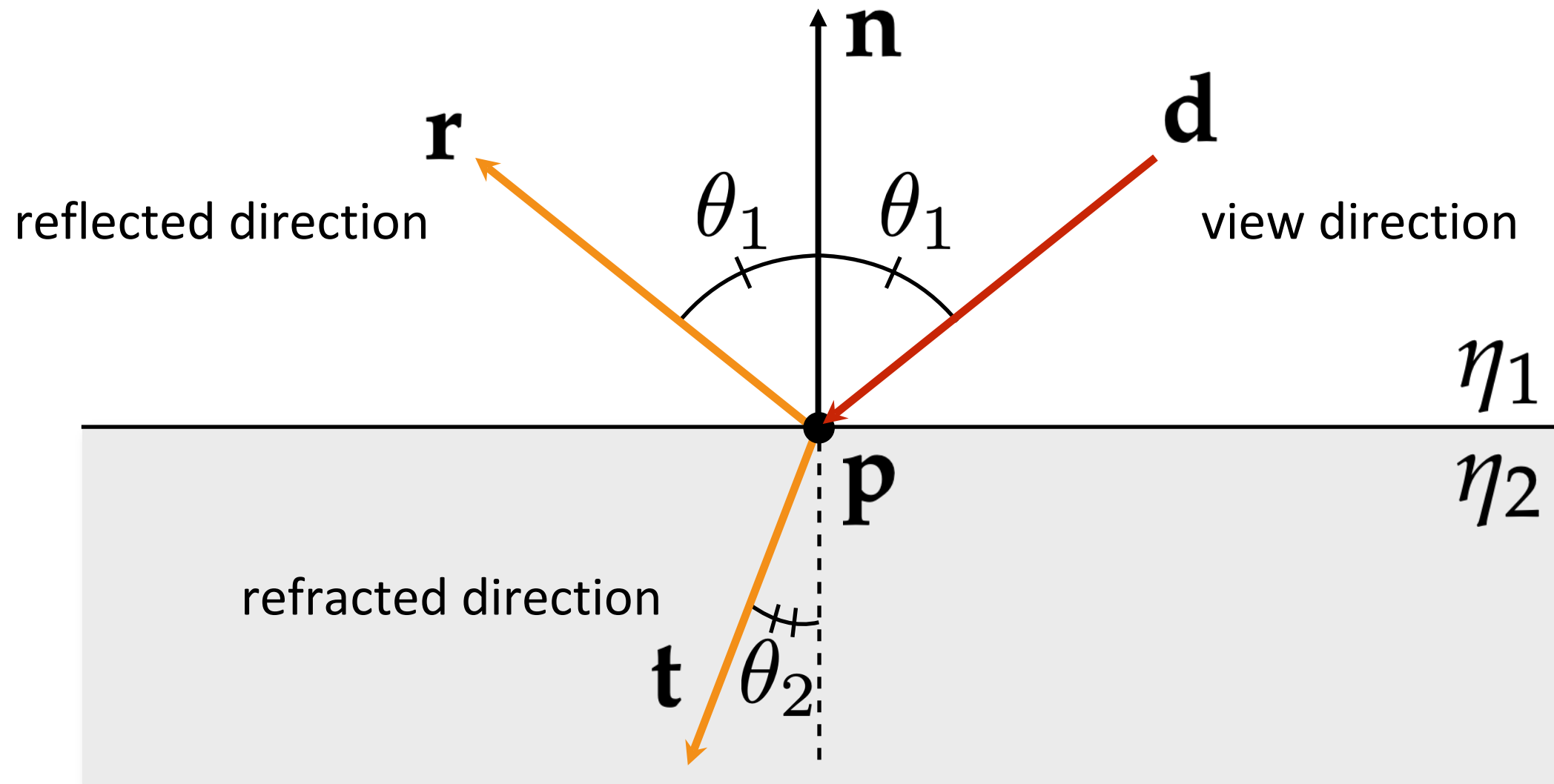
Total Internal Reflection



Reflection vs. Refraction

How much light is reflected vs. refracted?

- in reality determined by “Fresnel equations”



Fresnel Equations

Reflection and refraction from smooth *dielectric* (e.g. glass) surfaces

Reflection from *conducting* (e.g. metal) surfaces

Derived from Maxwell equations

Involves polarization of the wave

Fresnel Equations for Dielectrics

Reflection of light polarized parallel and perpendicular to the plane of refraction

$$\rho_{||} = \frac{\eta_2 \cos \theta_1 - \eta_1 \cos \theta_2}{\eta_2 \cos \theta_1 + \eta_1 \cos \theta_2}$$

$$\rho_{\perp} = \frac{\eta_1 \cos \theta_1 - \eta_2 \cos \theta_2}{\eta_1 \cos \theta_1 + \eta_2 \cos \theta_2}$$

$$\text{reflected: } F_r = \frac{1}{2} \left(\rho_{||}^2 + \rho_{\perp}^2 \right)$$

$$\text{refracted: } F_t = 1 - F_r$$

What's happening in this photo?



Polarizing Filter



Polarization



Without Polarizer



With Polarizing Filter

Polarization



Without Polarizer



With Polarizing Filter

Effect of Polarization



Effect of Polarization



Fresnel Equations for Dielectrics

Reflection of light polarized parallel and perpendicular to the plane of refraction

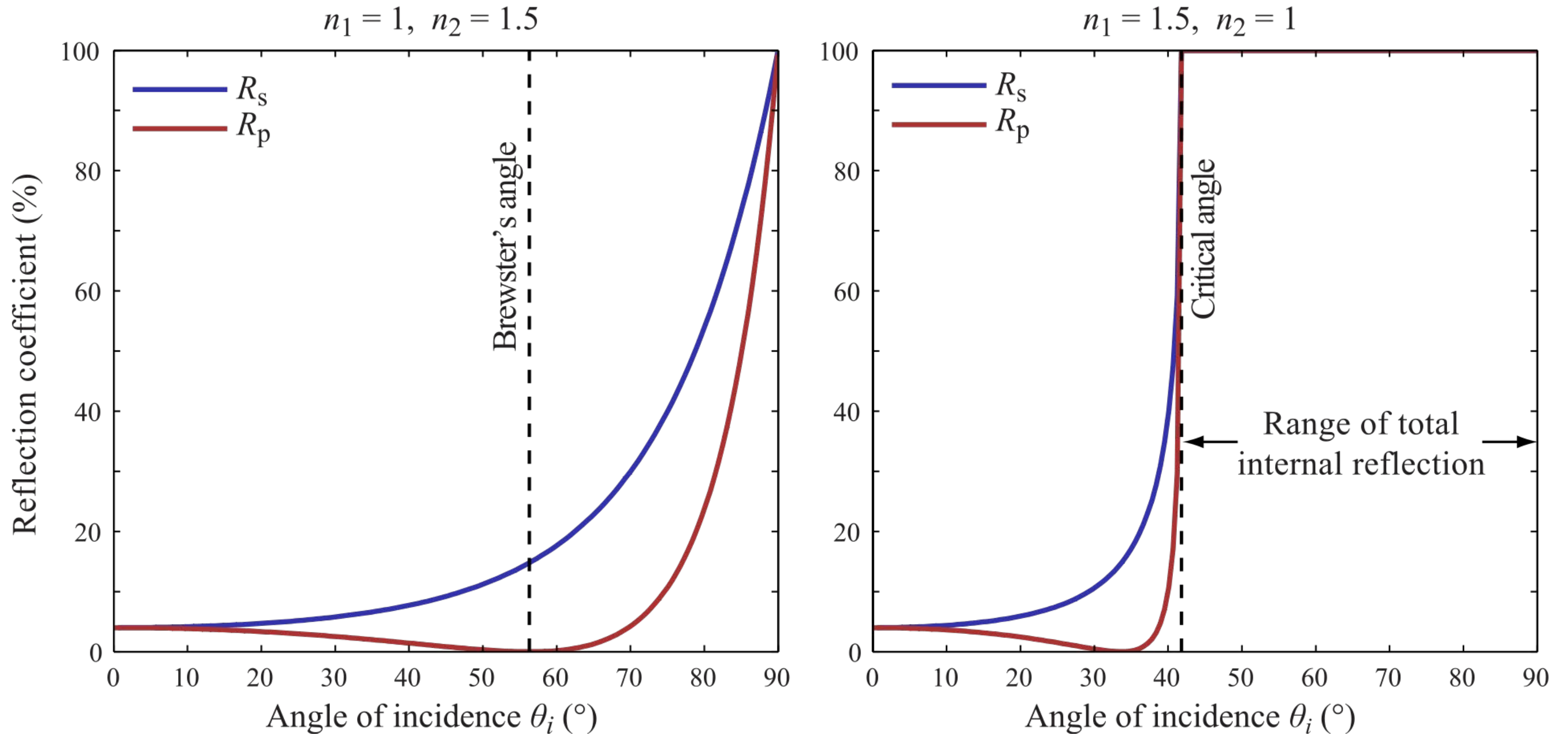
$$\rho_{||} = \frac{\eta_2 \cos \theta_1 - \eta_1 \cos \theta_2}{\eta_2 \cos \theta_1 + \eta_1 \cos \theta_2}$$
$$\rho_{\perp} = \frac{\eta_1 \cos \theta_1 - \eta_2 \cos \theta_2}{\eta_1 \cos \theta_1 + \eta_2 \cos \theta_2}$$

reflected: $F_r = \frac{1}{2} (\rho_{||}^2 + \rho_{\perp}^2)$

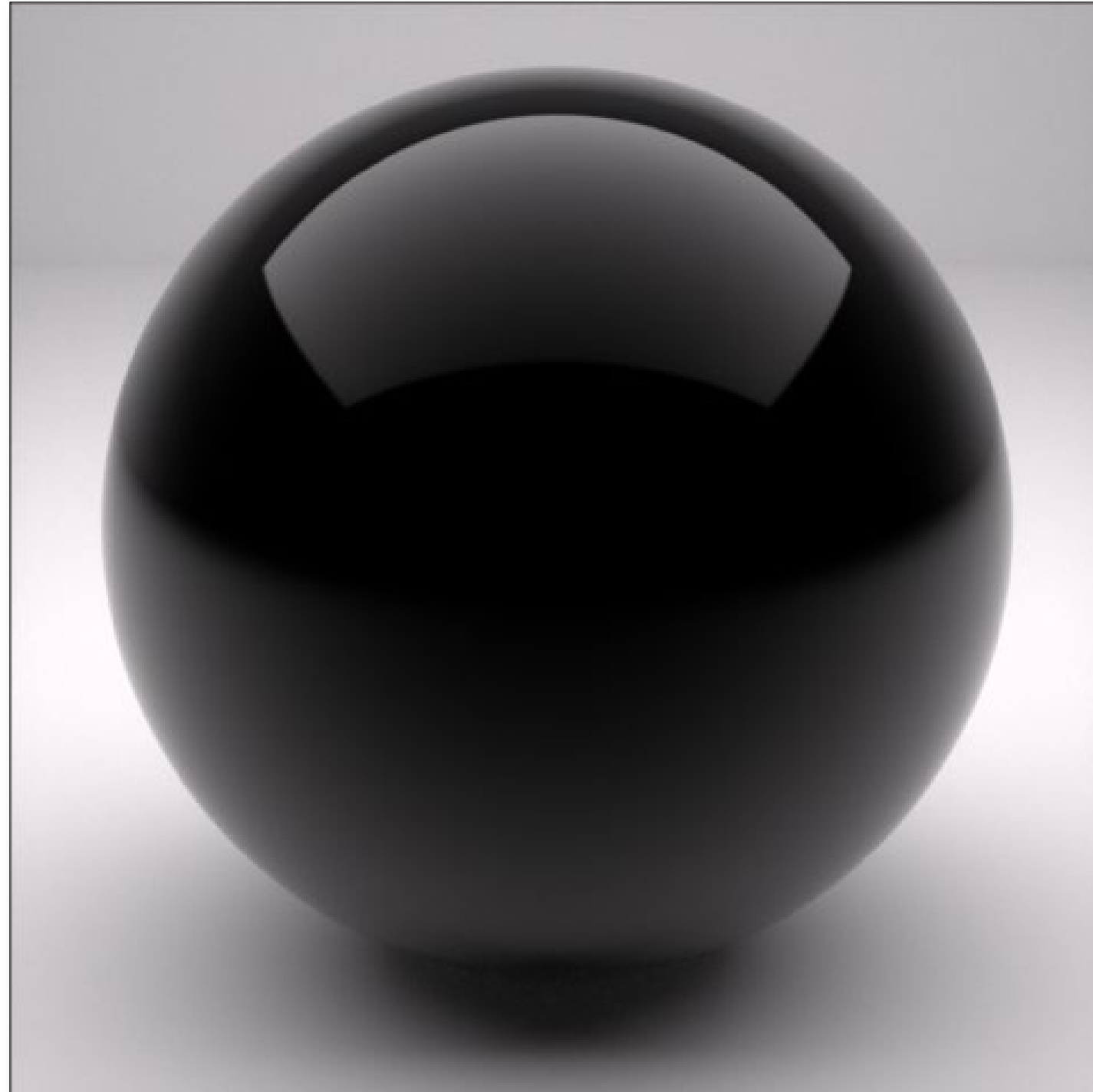
refracted: $F_t = 1 - F_r$

- The Shirley book uses a faster approximation (Schlick), but to get full accuracy you'd need to use these equations

Fresnel equations for glass



Fresnel reflection

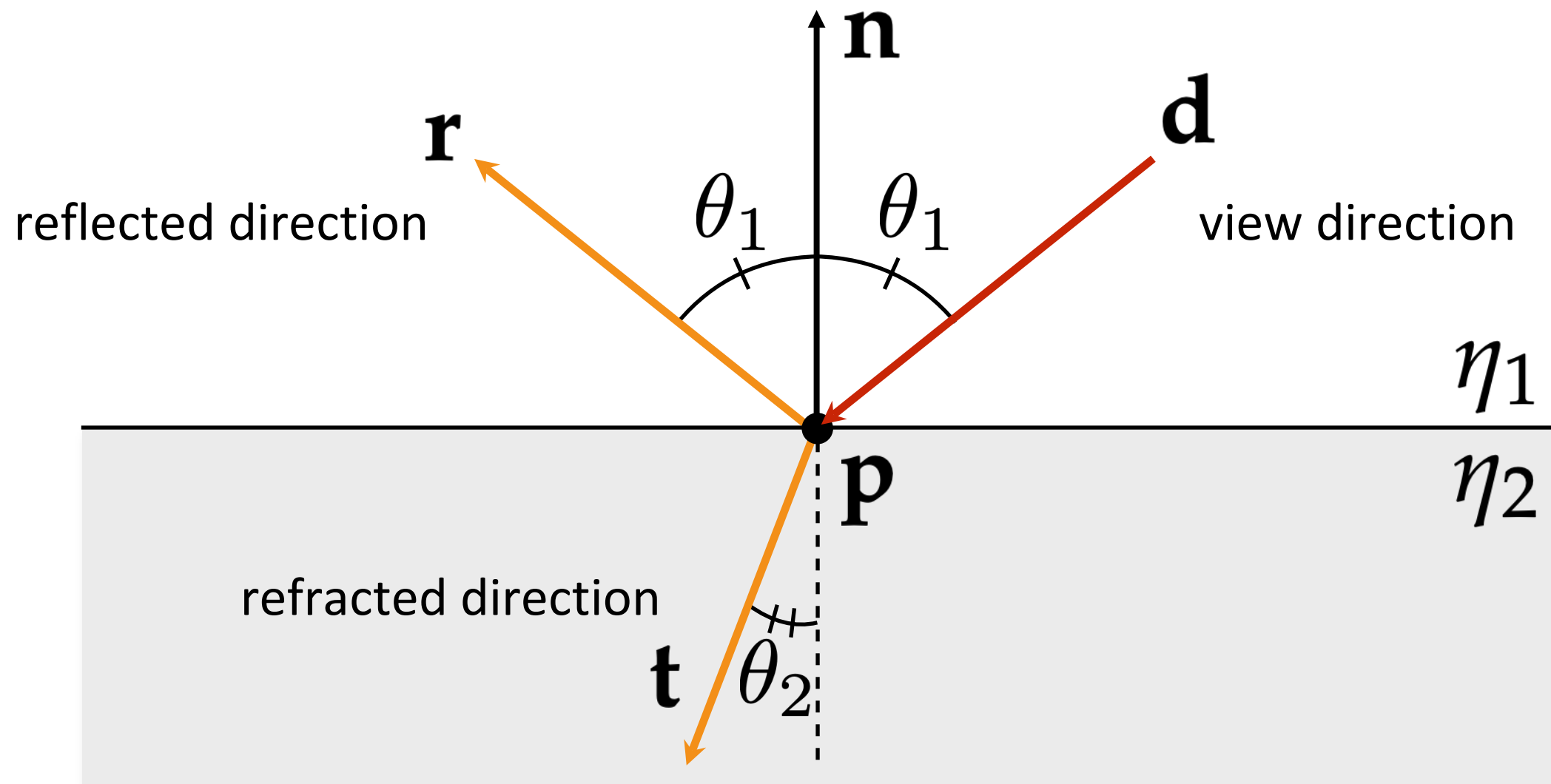


Fresnel reflection/refraction



Reflection vs. Refraction

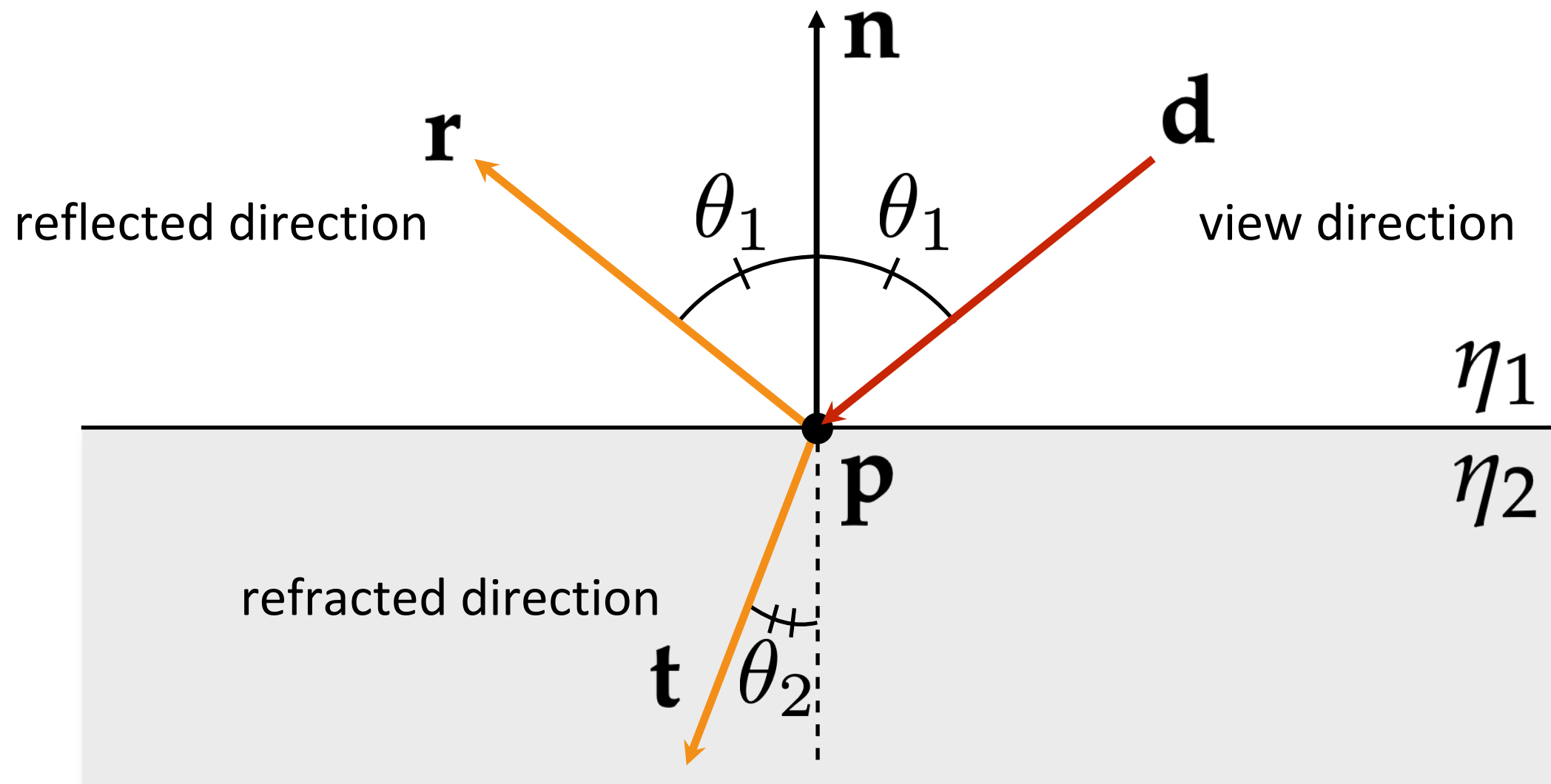
During ray tracing, how do we decide whether to reflect or refract?



Reflection vs. Refraction

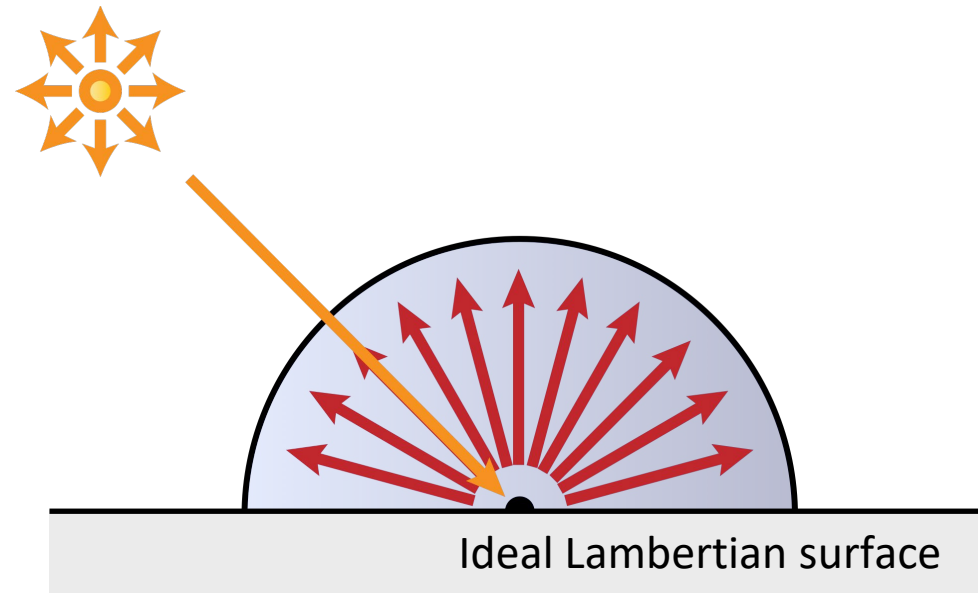
During ray tracing, how do we decide whether to reflect or refract?

- Randomly! Using Fresnel coefficients as probabilities.

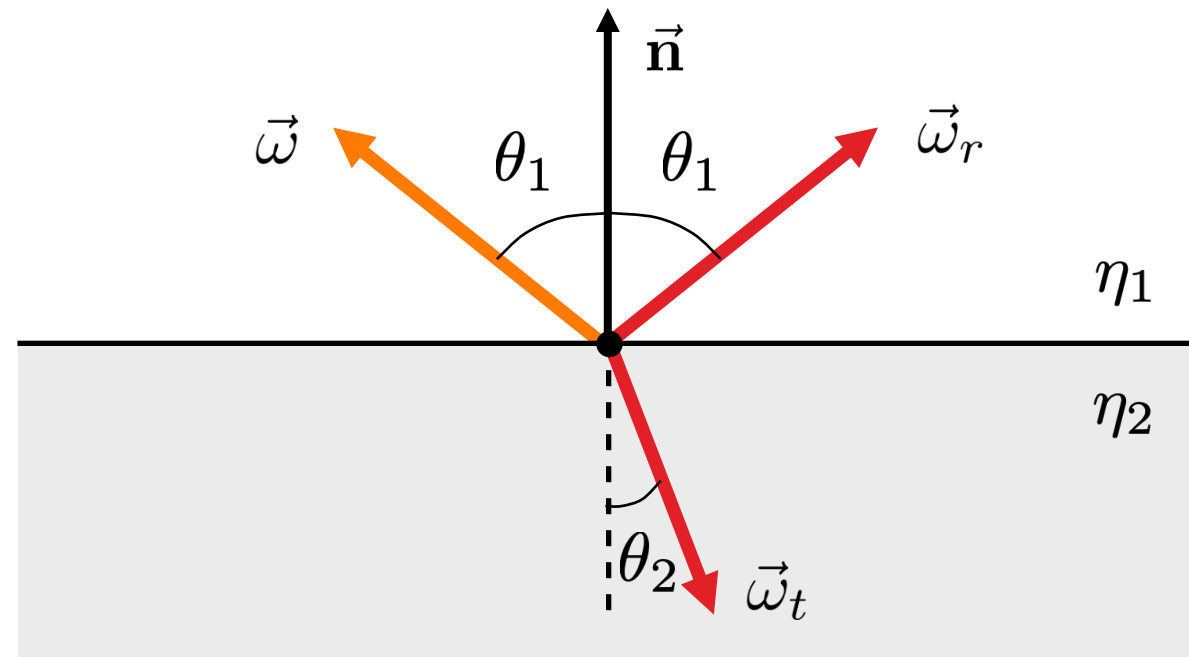


So Far: Idealized BRDF Models

Diffuse



Specular Reflection and Refraction



Real materials are more complex

Rough materials

In reality, most materials are neither perfectly diffuse nor specular, but somewhere in between

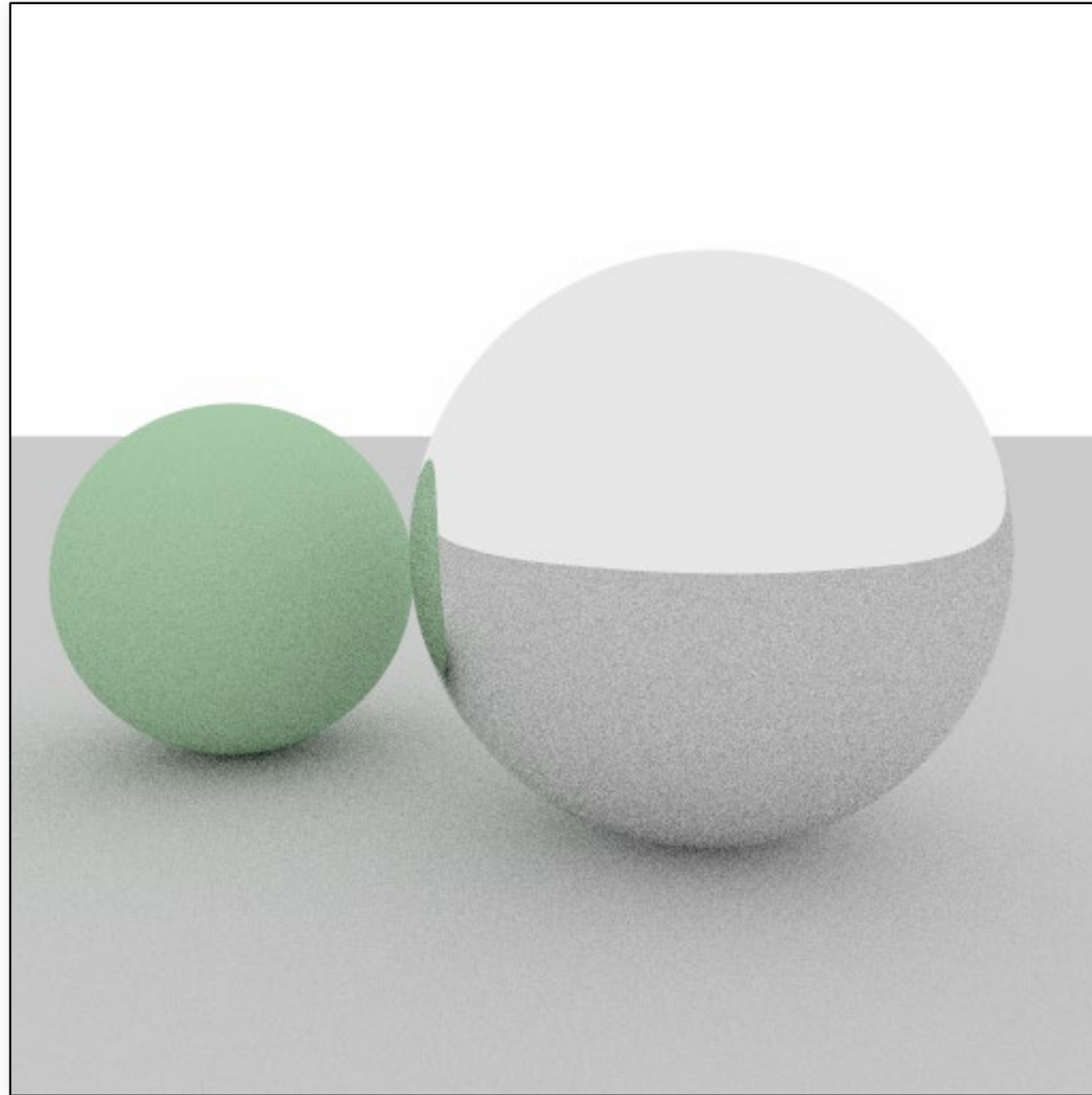
- Imagine a shiny surface scratched up at a microscopic level

We will look at a more principled way to handle this later.

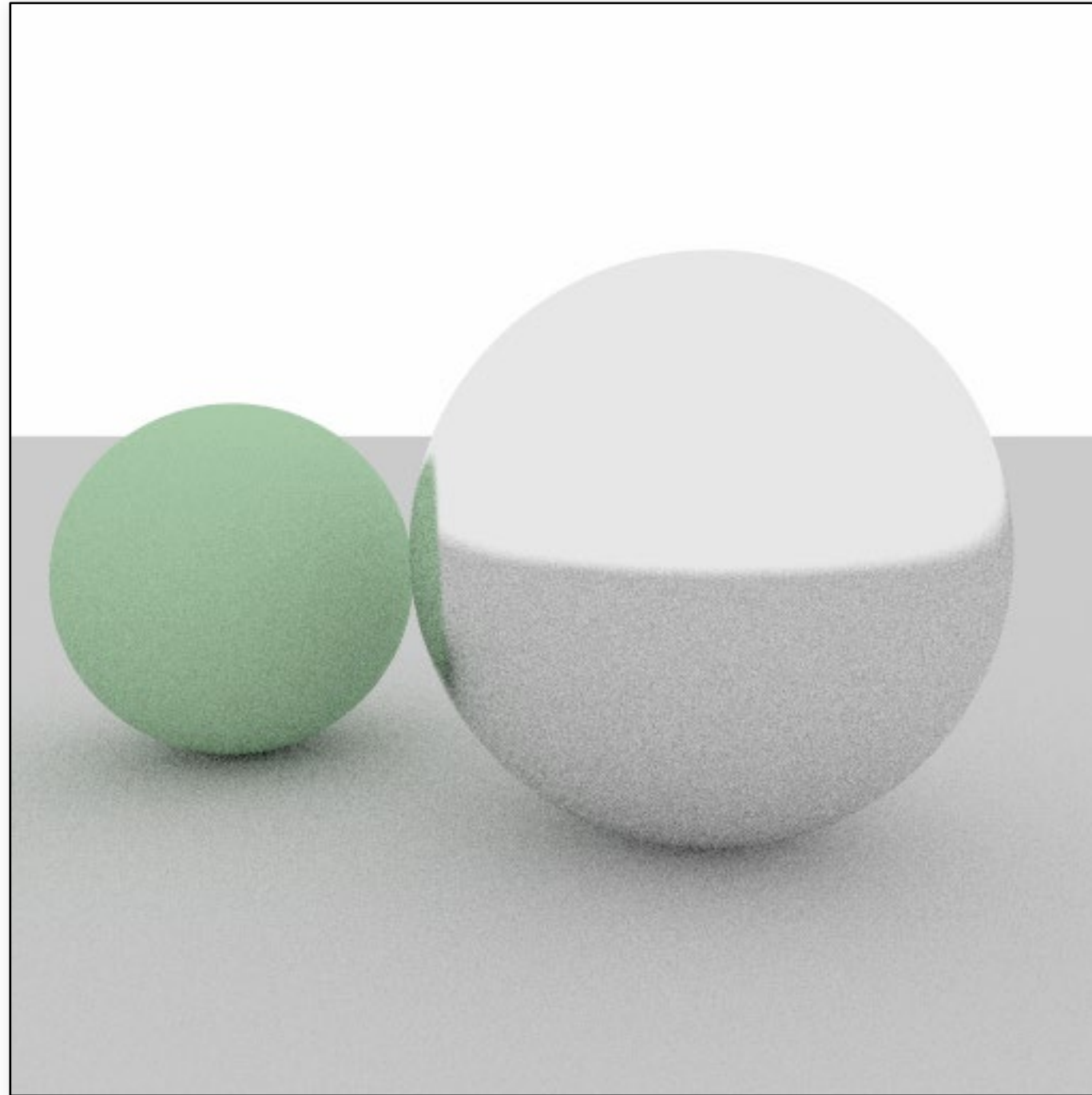
For now, we can easily approximate one important characteristic: blurred reflections

- Compute reflection direction, then add a random offset to it
- Sample random offset from sphere. Scale it to increase/decrease fuzziness

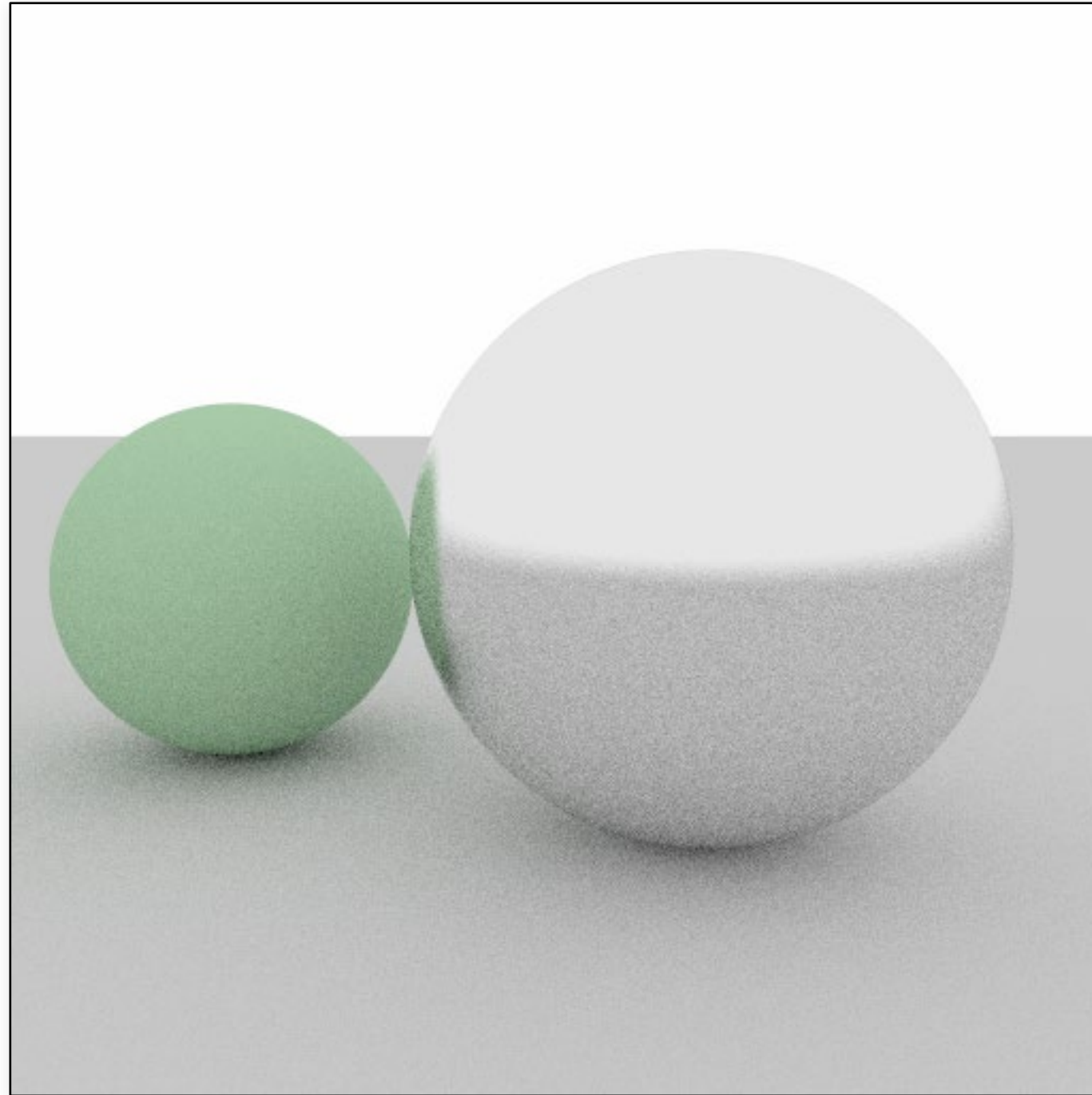
Diffuse & mirror spheres



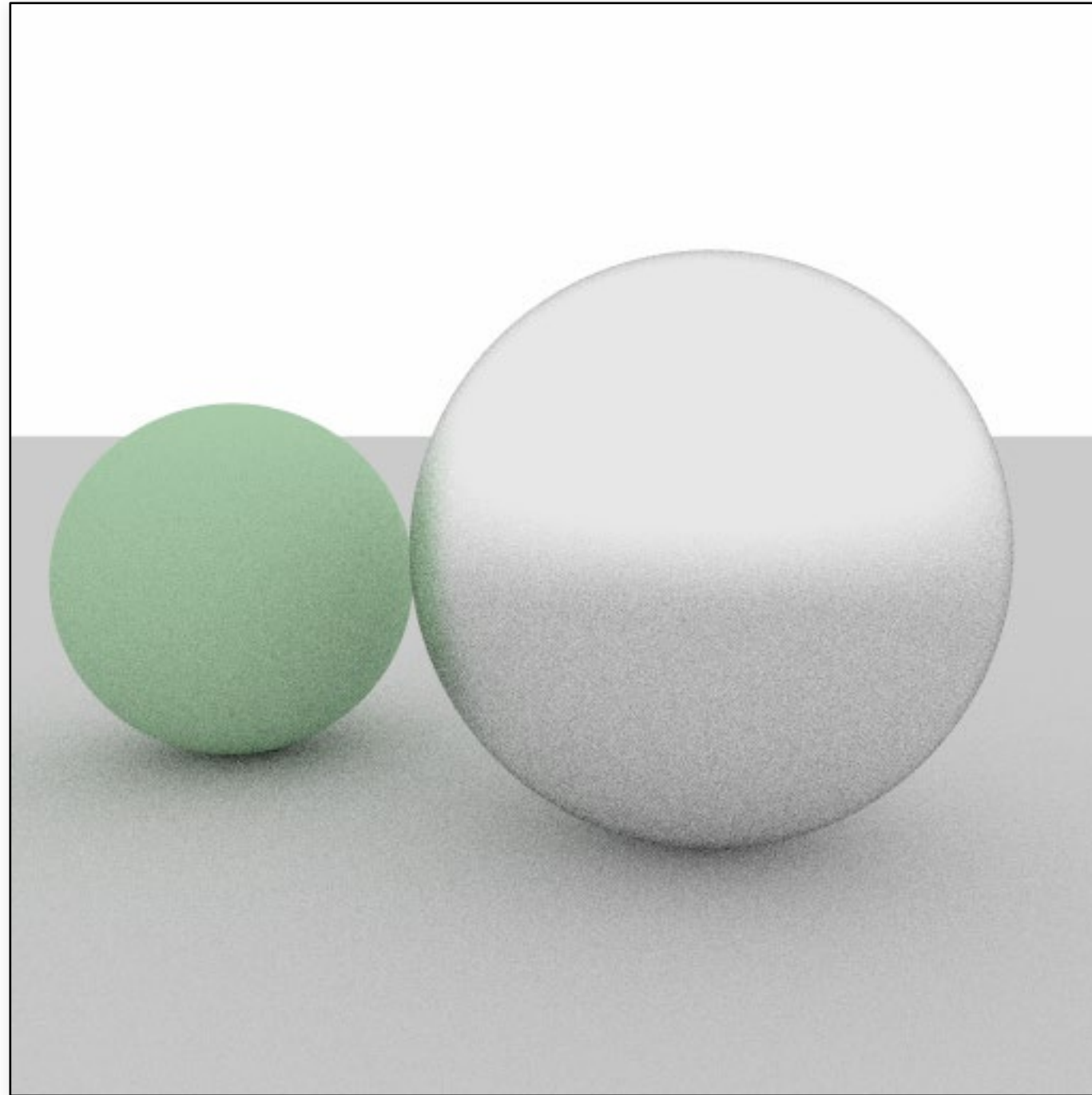
Diffuse & rough mirror spheres



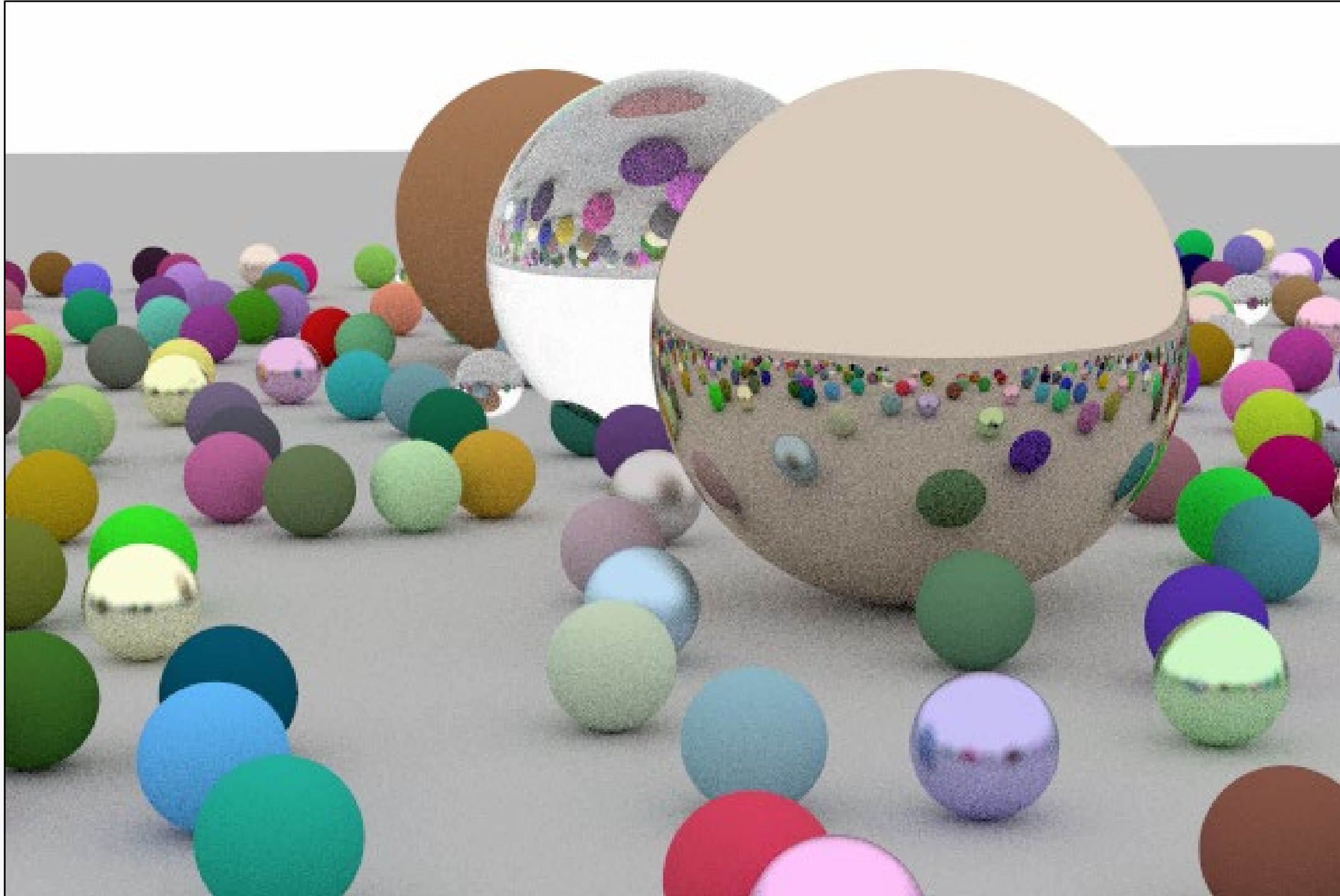
Diffuse & rough mirror spheres



Diffuse & rough mirror spheres



Putting it together



Lighting

Lighting

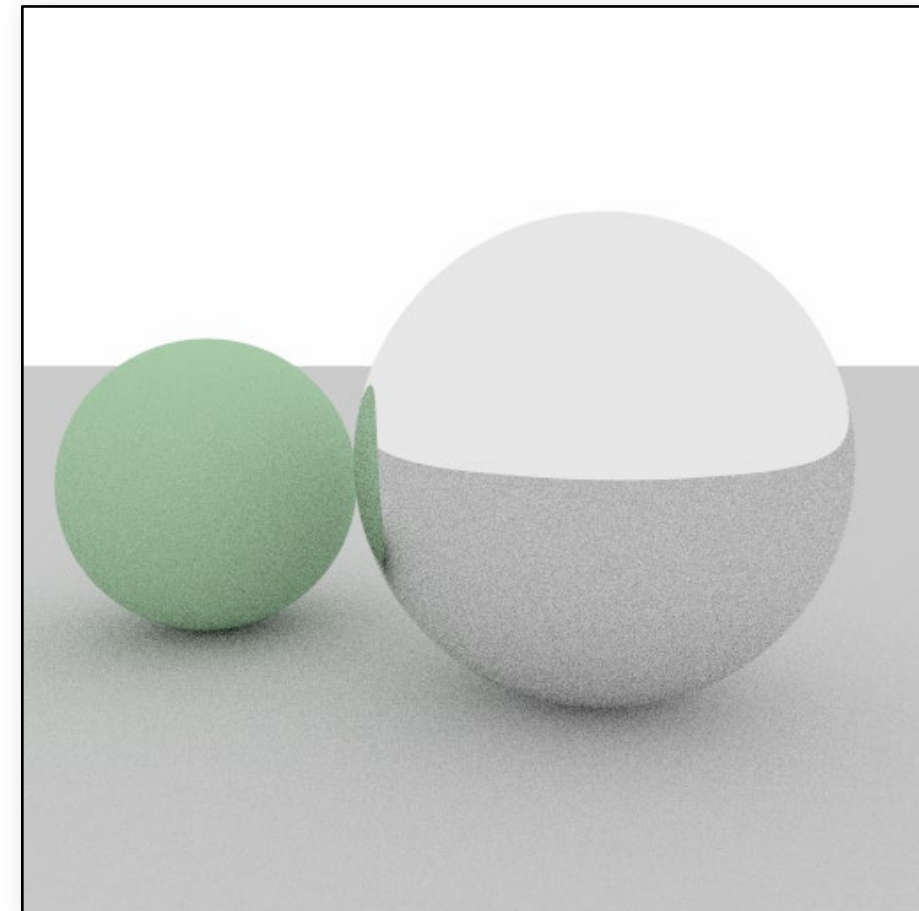
So far, the sky or background has been the only source of emitted light

But we can easily make any surface a light source!

- Just return an emitted color when a ray hits that surface
- Add a function to material that returns emitted color
 - Returns black (all zeros) for regular (non-emissive) surfaces
 - Color will often be greater than (1,1,1)
- Also possible for surfaces to emit & scatter (but not common)

Pseudo-code

```
Scene::trace(Ray ray)
    hit = surfaces.intersect(ray);
    if hit
        [col, sRay] = hit->mat->scatter(ray)
        return col * trace(sRay);
    else
        return backgroundColor;
```



Pseudo-code

```
Scene::trace(Ray ray)
    hit = surfaces.intersect(ray);
    if hit
        emit = hit->mat->emit(ray)
        [col, sRay] = hit->mat->scatter(ray)
        return emit + col * trace(sRay);
    else
        return backgroundColor;
```

